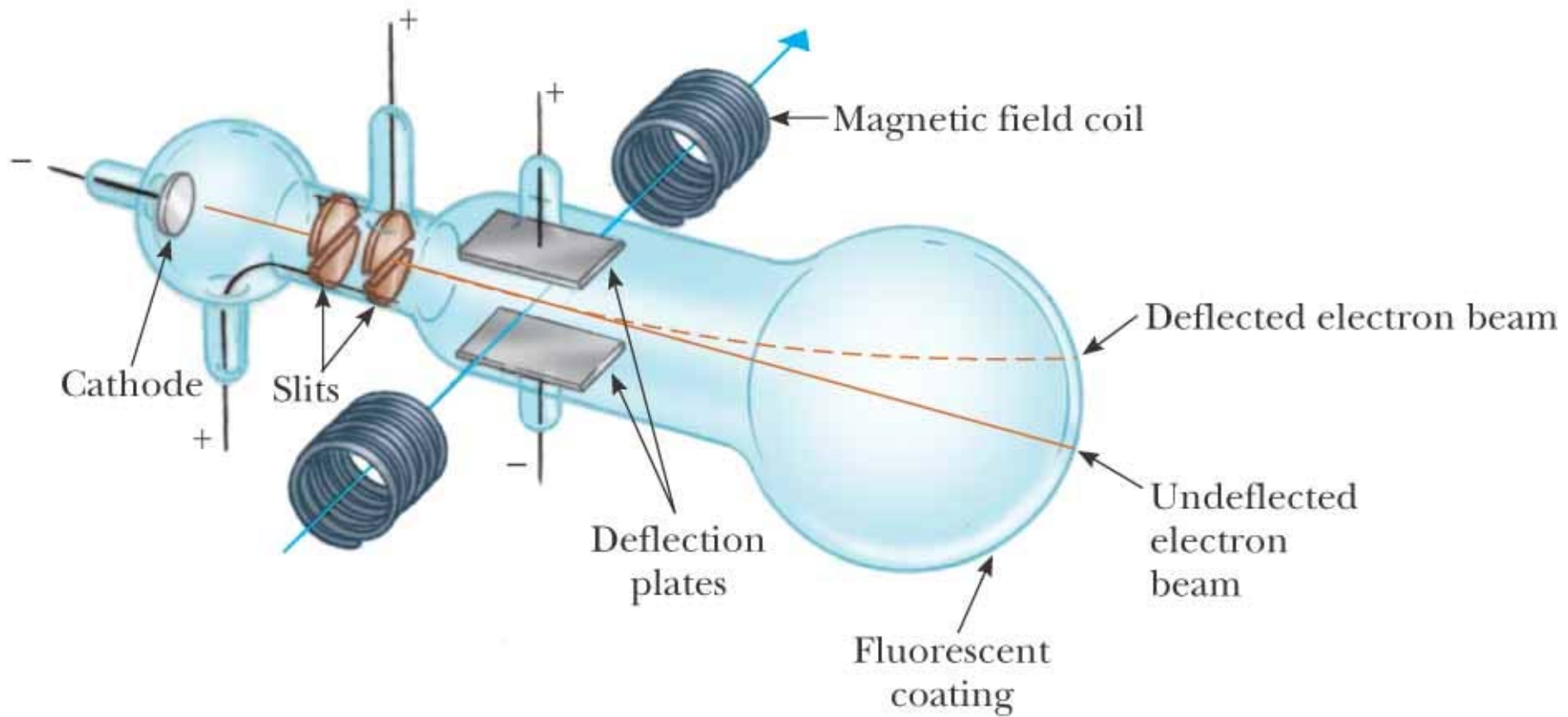


Introduction

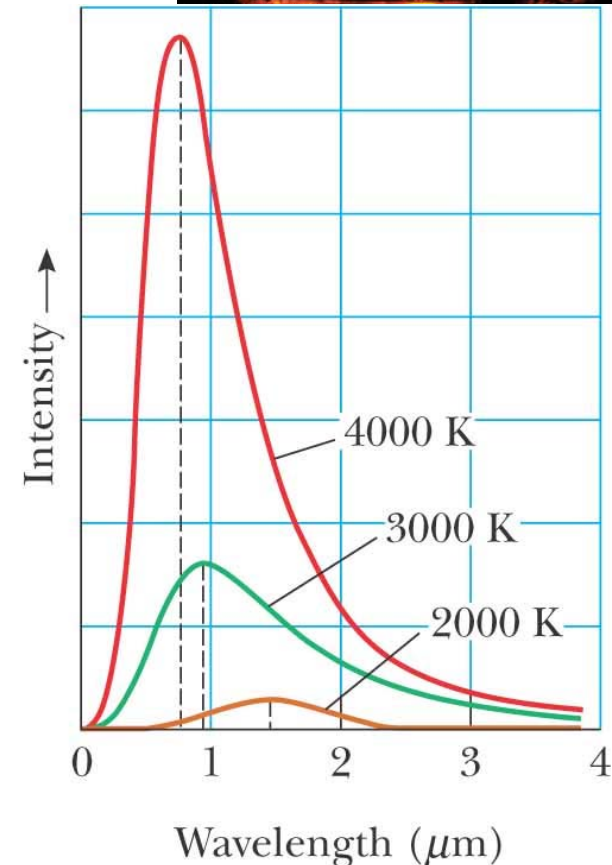
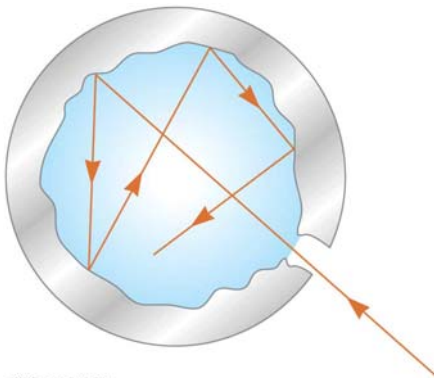
- By the end of the 19th century with Newtonian mechanics and Maxwell's equations, physics almost seemed complete
- But some puzzling observations:
 - Michelson-Morley experiment:
 - **Special and General relativity**
 - Atomic line spectra
 - Properties of atoms
 - Blackbody Radiation
 - Photoelectric effect
 - **“Old Quantum theory” and Quantum mechanics.**



(a)

Blackbody Radiation

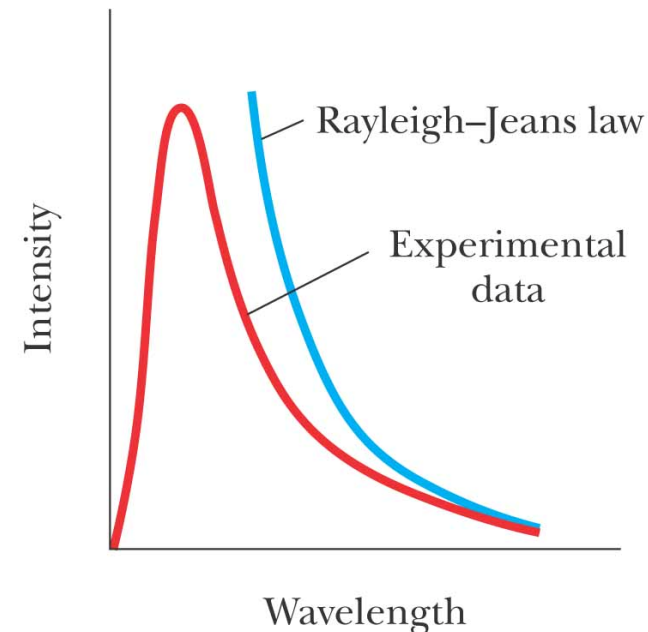
- A Blackbody is an ideal system that absorbs all radiation incident on it. Emission of radiation by a blackbody is independent of the properties of its wall, but depends only on its temperature
- When the temperature of the blackbody increases, the spectrum of emitted radiation shifts to lower λ .



- Rayleigh-Jeans law was an attempt to explain blackbody radiation based on classical ideas:
 - Blackbody is modeled in terms of modes of oscillations of the e/m field in the cavity due to the accelerating charges on cavity walls.
 - # of Degrees of Freedom for oscillations go up as λ goes down
 - Equipartition of energy: Each DOF carries an energy $k_B T$: where k_B is the Boltzman constant.
 - Intensity (power per unit area):

$$I(\lambda, T) = \frac{2\pi k_B T}{\lambda^4}$$

Ultraviolet Catastrophe



Max Planck's theory of Blackbody radiation

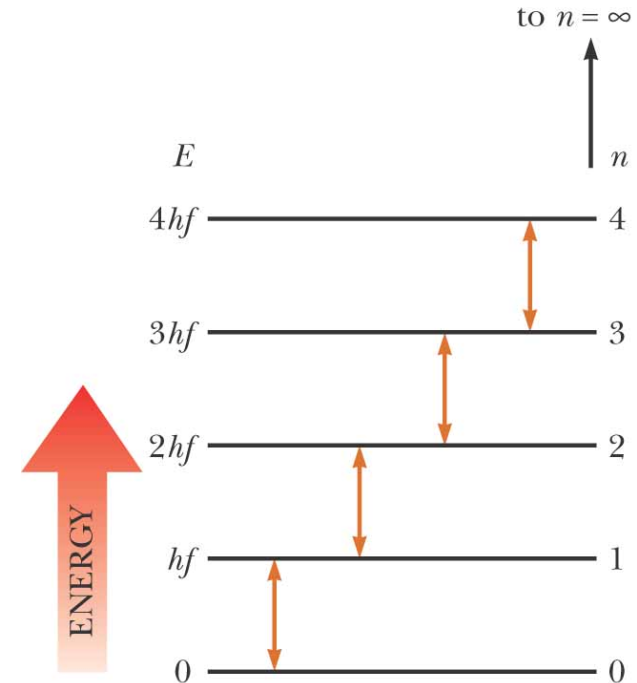
Planck's assumptions on the nature of the oscillators in the cell walls:

- The energy of an oscillator can only have a **discrete** set of values E_n :

$$E_n = nh\nu$$

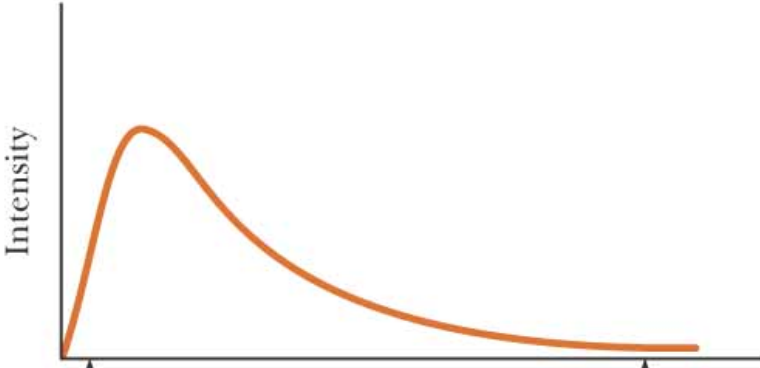
n is the **quantum number**, ν is the frequency of oscillations and h is a new parameter Planck introduced: it is called **Planck's Constant** now. Energy is said to be **quantized** and each discrete value of energy corresponds to a **quantum state**.

The oscillators emit or absorb energy when making a transition from one quantum state to another. The amount of energy emitted or absorbed is equal to the energy difference between initial and final quantum states:



- As λ increases the gap between energy levels goes down
- Boltzman's distribution law:
 - the probability of a state of being occupied is proportional to:

$$e^{-E/k_B T}$$

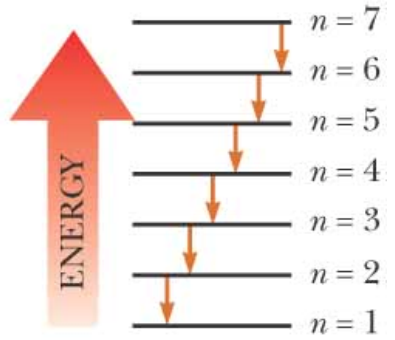
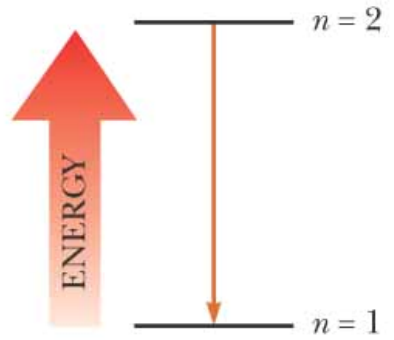


- At short wavelengths:
- large energy separation
 - low probability of excited states
 - few downward transitions

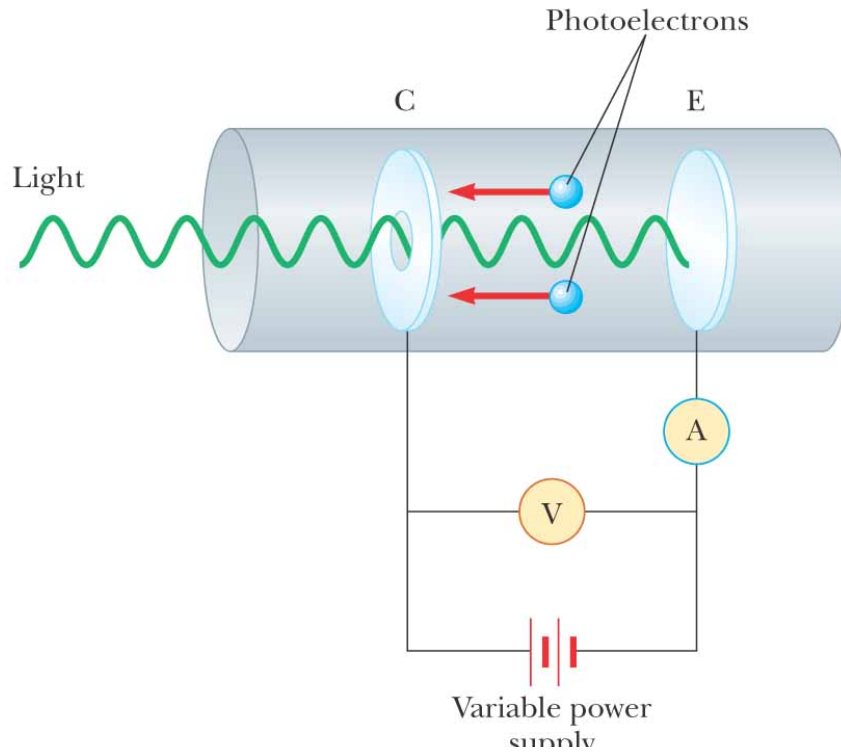
- At long wavelengths:
- small energy separation
 - high probability of excited states
 - many downward transitions

$$I(\lambda, T) = \frac{2\pi h c^2}{\lambda^5 (e^{hc/\lambda k_B T} - 1)}$$

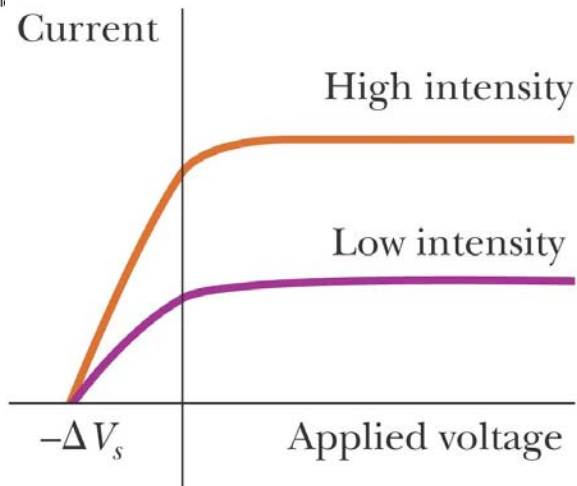
$$h = 6.626 \times 10^{-34} \text{ J}\cdot\text{s}$$



The Photoelectric Effect



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•A classical explanation of the photoelectric effect would predict:

- Kinetic energy of the photoelectrons should increase with light intensity
- Emission of photoelectrons and their kinetic energy should be independent of the light frequency
- When the light intensity is very low there should be a time delay before the photoelectron emission

•Experimental observation:

- Kinetic energy of the photoelectrons is independent of the light intensity but increases with the frequency
- No photoelectrons are emitted below some **cutoff frequency** f_c
- Photoelectrons are emitted instantaneously independent of the intensity.

In 1905 Einstein extended Planck's quantization to e/m waves to explain the photoelectric effect:

- Light and other e/m waves consist of particles (quanta) called photons.
- Each photon has energy $E = hf$, where f is the frequency of the wave.
- An incident light photon gives all its energy to a single electron in the metal

Absorption of light is not continuous: it is delivered in discrete packets

$$K_{\max} = h\nu - W$$

W : work function of the metal; the minimum energy required to remove an electron from the metal

The Compton Effect: Scattering of X-rays from electrons

Einstein: Photons of light: Energy = $h\nu$; momentum = $E/c=h\nu/c$

