

The curvature tensor

Covariant derivative of contravariant vector

The covariant derivative of a (contravariant) vector is

$$V^\mu{}_{;v} = \partial_v V^\mu + \left\{ \begin{matrix} \mu \\ v\sigma \end{matrix} \right\} V^\sigma. \quad (19.1)$$

We used this in Eq. 18.2 without explaining it. Where does it come from? We know that the derivative of a scalar is a covariant vector,

$$\overset{df}{\varphi}_{,\mu} = \partial_\mu \varphi.$$

Now, suppose the scalar is the contraction of 2 vectors:

$$\varphi = A_\mu B^\mu \quad (19.2)$$

then by definition (and the product rule)

$$\begin{aligned} \partial_v \varphi &= A_{\mu,v} B^\mu + A_\mu B^\mu{}_{,v} \equiv A_{\mu;v} B^\mu + A_\mu B^\mu{}_{;v} \\ &= \left[A_{\mu;v} + \left\{ \begin{matrix} \sigma \\ v\mu \end{matrix} \right\} A_\sigma \right] B^\mu + A_\mu B^\mu{}_{;v} \end{aligned} \quad (19.3)$$

From Eq. 19.3 we have

$$A_{\mu;v} B^\mu + A_\mu B^\mu{}_{;v} = A_{\mu;v} B^\mu + \left\{ \begin{matrix} \sigma \\ v\mu \end{matrix} \right\} A_\sigma B^\mu + A_\mu B^\mu{}_{;v}$$

or

$$A_\mu B^\mu{}_{;v} = A_\mu B^\mu{}_{,v} + \left\{ \begin{matrix} \sigma \\ v\mu \end{matrix} \right\} A_\sigma B^\mu = A_\mu \left[B^\mu{}_{,v} + \left\{ \begin{matrix} \mu \\ v\sigma \end{matrix} \right\} B^\sigma \right] \quad (19.4)$$

and since A_μ is arbitrary, we may say

$$B^\mu{}_{;v} = B^\mu{}_{,v} + \left\{ \begin{matrix} \mu \\ v\sigma \end{matrix} \right\} B^\sigma. \quad (19.5)$$

Covariant derivative of tensor

By similar manipulations, we can identify the covariant derivative of a contravariant second-rank tensor—we write

$$\varphi = A_\mu B_\nu T^{\mu\nu} \quad (19.6)$$

and use the product rule again to write

$$\begin{aligned} \partial_\kappa \varphi &= A_{\mu,\kappa} B_\nu T^{\mu\nu} + A_\mu B_{\nu,\kappa} T^{\mu\nu} + A_\mu B_\nu T^{\mu\nu}{}_{,\kappa} \\ &\equiv A_{\mu;\kappa} B_\nu T^{\mu\nu} + A_\mu B_{\nu;\kappa} T^{\mu\nu} + A_\mu B_\nu T^{\mu\nu}{}_{;\kappa} \end{aligned} \quad (19.7)$$

to find

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$$T^{\mu\nu}{}_{;\kappa} = T^{\mu\nu}{}_{,\kappa} + \left\{ \begin{matrix} \mu \\ \kappa\sigma \end{matrix} \right\} T^{\sigma\nu} + \left\{ \begin{matrix} \nu \\ \kappa\sigma \end{matrix} \right\} T^{\mu\sigma} \quad (19.8)$$

and so forth.

Geodesic coordinates

Suppose we locally change coordinates to a system $x'^{\mu} = b^{\mu}_{\sigma} x^{\sigma}$, with (linear) transformation coefficients b^{μ}_{σ} chosen such that the new metric at that point is

$$g'^{\mu\nu} \stackrel{df}{=} g^{\alpha\beta} b^{\mu}_{\alpha} b^{\nu}_{\beta} = \eta^{\mu\nu}. \quad (19.9)$$

Equations 19.9 constitute 10 inhomogeneous equations for 16 unknowns b^{μ}_{σ} , whose determinant,

$$\det[g^{\alpha\beta}] = \frac{1}{-g},$$

is non-zero. Therefore they can always be solved, leaving 6 free parameters. These are in fact the 6 parameters of the Lorentz transformation (3 boost, 3 rotation) which, as we already know, leave the Minkowski metric unchanged.

Moreover, we can specify the coordinates further so that in the new system, all first derivatives of the new metric, $g'_{\mu\nu, \kappa}$ vanish at the point a^{σ} . The coordinates that do this are

$$\begin{aligned} x'^{\mu} = & b^{\mu}_{\sigma} x^{\sigma} + \frac{1}{2} \Gamma^{\mu}_{\sigma\kappa} (x^{\sigma} - a^{\sigma})(x^{\kappa} - a^{\kappa}) + \\ & + \frac{1}{3!} \Lambda^{\mu}_{\sigma\kappa\lambda} (x^{\sigma} - a^{\sigma})(x^{\kappa} - a^{\kappa})(x^{\lambda} - a^{\lambda}) + \dots \end{aligned}$$

where the coefficients $\Gamma^{\mu}_{\sigma\kappa}$ and $\Lambda^{\mu}_{\sigma\kappa\lambda}$ are manifestly symmetric in their lower indices, hence represent 20 and 80 independent parameters, respectively.

Problem:

An object with 3 indices that run from 0 to 3 obviously has 64 components. Show that if the object is fully symmetric in the 3 indices, then there are but 20 independent components.

Hence show that $\Lambda^{\mu}_{\sigma\kappa\lambda}$ has 80 independent components.

Problem:

Find the relation between the coefficients $\Gamma^{\mu}_{\sigma\kappa}$ and the Christoffel symbols $\left\{ \begin{matrix} \mu \\ \sigma\kappa \end{matrix} \right\} (a)$ defined at the point a^{σ} in terms of the (derivatives of the) old metric $g_{\mu\nu}$.

Since there are 20 first derivatives of the metric tensor, we can obviously choose the coefficients $\Gamma^{\mu}_{\sigma\kappa}$ to set the derivatives of the new metric tensor equal to zero at one point. Since then the

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Christoffel symbol vanishes, the motion of a test body in these new coordinates is unaccelerated, *i.e.* freely falling.

What about the second derivatives of the metric tensor? There are 100 distinct components,

$$g'_{\mu\nu, \kappa\sigma} ,$$

(two distinct pairs of symmetric indices, giving $10 \times 10 = 100$), but only 80 independent parameters $\Lambda^{\mu}_{\sigma\kappa\lambda}$, hence there will be 20 quantities involving second derivatives of the metric tensor, that cannot be made to vanish at a point by a coordinate transformation.

In a frame where the first derivatives of the metric tensor can be chosen to vanish at a point, the Christoffel symbols also vanish at that point, hence

$$\frac{d^2 x'^{\mu}}{d^2 \tau} = 0 . \quad (19.10)$$

The new coordinates at that point are freely falling, hence the name *geodesic*.

We have spoken before of parallel transport, and concluded that when a vector A_{μ} is transported an amount δx^{κ} parallel to itself, the change in A_{μ} , arising from the change in the coordinates, is

$$\delta A_{\mu} = \left\{ \begin{matrix} \sigma \\ \kappa\mu \end{matrix} \right\} A_{\sigma} \delta x^{\kappa} . \quad (19.11)$$

We can think of the covariant derivative as the difference between the ordinary derivative and the change that would occur if the vector were merely parallel-transported; hence the change in a *contravariant* vector under parallel transport is

$$\delta A^{\mu} = - \left\{ \begin{matrix} \mu \\ \kappa\sigma \end{matrix} \right\} A_{\sigma} \delta x^{\kappa} . \quad (19.12)$$

Finally, we note that the 4-velocity $U^{\mu} = \frac{dx^{\mu}}{d\tau}$ is always parallel-transported; moreover, the contraction of the velocity and A_{μ} is a scalar that is invariant under parallel transport

$$\delta (U^{\mu} A_{\mu}) = \left[U^{\sigma} \left\{ \begin{matrix} \mu \\ \sigma\kappa \end{matrix} \right\} A_{\mu} - A_{\mu} \left\{ \begin{matrix} \mu \\ \sigma\kappa \end{matrix} \right\} U^{\sigma} \right] \delta x^{\kappa} = 0 . \quad (19.13)$$

To put it another way, the angle between a vector that is parallel-transported and the velocity is always constant.

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Let us parallel-transport a vector around an infinitesimal closed curve parameterized by $x^\mu(p)$: if $\tilde{V}_\mu(0) \equiv V_\mu(x^\kappa(0))$, we find

$$\tilde{V}_\mu(p) - \tilde{V}_\mu(0) = \int_0^p dp' \frac{d\tilde{V}_\mu(p')}{dp'} = \int_0^p dp' \left\{ \begin{matrix} \sigma \\ \kappa\mu \end{matrix} \right\} (p') \tilde{V}_\sigma(p') \frac{dx^\kappa}{dp'}. \quad (19.14)$$

Let $x^\mu(p=0) = a^\mu$; we can then expand in Taylor's series about $p=0$:

$$\tilde{V}_\mu(p') \approx V_\mu(a) + \left\{ \begin{matrix} \sigma \\ \kappa\sigma \end{matrix} \right\} (a) V_\sigma(a) (x^\kappa(p') - a^\kappa) \quad (19.15)$$

and

$$\left\{ \begin{matrix} \sigma \\ \kappa\sigma \end{matrix} \right\} (x(p')) \approx \left\{ \begin{matrix} \sigma \\ \kappa\sigma \end{matrix} \right\} (a) + (x^\lambda(p') - a^\lambda) \frac{\partial}{\partial x^\lambda} \left\{ \begin{matrix} \sigma \\ \kappa\sigma \end{matrix} \right\} (a). \quad (19.16)$$

Therefore

$$\begin{aligned} \tilde{V}_\mu(p') - \tilde{V}_\mu(a) &\approx \int_0^p dp' \left[\left\{ \begin{matrix} \sigma \\ \mu\nu \end{matrix} \right\} (a) + (x^\lambda(p') - a^\lambda) \partial_\lambda \left\{ \begin{matrix} \sigma \\ \mu\nu \end{matrix} \right\} (a) \right] \cdot \\ &\quad \left[V_\sigma(a) + (x^\kappa(p') - a^\kappa) \partial_\kappa \left\{ \begin{matrix} \alpha \\ \kappa\sigma \end{matrix} \right\} (a) V_\alpha(a) \right] \end{aligned} \quad (19.17)$$

Thus, expanding in powers of

$$\delta x^\lambda(p') = x^\lambda(p') - a^\lambda,$$

we find

$$\begin{aligned} \tilde{V}_\mu(p) - V_\mu(a) &\approx \left\{ \begin{matrix} \sigma \\ \nu\mu \end{matrix} \right\} (a) V_\sigma(a) \int_0^p dp' \frac{dx^\nu}{dp'} + \\ &+ \left[\left\{ \begin{matrix} \sigma \\ \nu\mu \end{matrix} \right\} (a) \left\{ \begin{matrix} \alpha \\ \lambda\sigma \end{matrix} \right\} (a) V_\alpha(a) + V_\sigma(a) \partial_\lambda \left\{ \begin{matrix} \sigma \\ \nu\mu \end{matrix} \right\} (a) \right] \int_0^p dp' \frac{dx^\nu}{dp'} \delta x^\lambda(p') + 0(\delta x^\nu \delta x^\lambda) \end{aligned} \quad (19.18)$$

We drop the $0(\delta x^\nu \delta x^\lambda)$ term in Eq. 19.18 because we shall ultimately take the neighborhood of a^λ to be arbitrarily small.

Now, since the parameterization describes a closed curve, we have

$$\int_0^p dp' \frac{dx^\nu}{dp'} = \oint dx^\nu(p') \quad (19.19)$$

Thus we have to evaluate

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$$\oint dx^y (p) (x^\lambda (p) - a^\lambda) \equiv \oint dx^y (p) x^\lambda (p) \quad (19.20)$$

We see immediately that, because $d(x^y x^\lambda) = x^\lambda dx^y + x^y dx^\lambda$ is a perfect derivative,

$$\oint dx^y x^\lambda = -\oint dx^\lambda x^y \quad (19.21)$$

Thus we have an expression of the form

$$\delta \tilde{V}_\mu = \frac{1}{2} R_{\mu\nu\lambda}^\sigma (a) V_\sigma (a) \oint dx^y x^\lambda \quad (19.22)$$

where equation 19.22 defines the *Riemann curvature tensor*:

$$R_{\mu\nu\lambda}^\sigma = \partial_\lambda \left\{ \begin{matrix} \sigma \\ \nu\mu \end{matrix} \right\} - \partial_\nu \left\{ \begin{matrix} \sigma \\ \lambda\mu \end{matrix} \right\} + \left\{ \begin{matrix} \alpha \\ \nu\mu \end{matrix} \right\} \left\{ \begin{matrix} \sigma \\ \lambda\alpha \end{matrix} \right\} - \left\{ \begin{matrix} \alpha \\ \lambda\mu \end{matrix} \right\} \left\{ \begin{matrix} \sigma \\ \nu\alpha \end{matrix} \right\} \quad (19.23)$$

Clearly, $R_{\mu\nu\lambda}^\sigma = 0$ if the space is flat, *i.e.* if the first and second derivatives of the metric tensor vanish, since then the Christoffel symbols and their first derivatives vanish.

Does this mean that in a freely falling system the curvature tensor is zero? No, because while the Christoffel symbols vanish, their (ordinary) derivatives will not. Thus we can, in principle, distinguish between a flat space and a freely falling system in a curved space, by the non-vanishing of the curvature in the latter case.

We note that $R_{\mu\nu\lambda}^\sigma$ is a tensor by construction, since everything else in Eq. 19.22 is a tensor. We also note that it is antisymmetric in $\lambda\nu$. If one lowers the top index to produce the 4th rank covariant tensor

$$R_{\kappa\mu\nu\lambda} = g_{\kappa\sigma} R_{\mu\nu\lambda}^\sigma,$$

we find the latter satisfies four identities:

$$\begin{aligned} R_{\kappa\sigma\mu\nu} &\equiv -R_{\sigma\kappa\mu\nu} \\ R_{\kappa\sigma\mu\nu} &\equiv -R_{\kappa\sigma\nu\mu} \\ R_{\kappa\sigma\mu\nu} &\equiv R_{\mu\nu\kappa\sigma} \\ R_{\kappa\sigma\mu\nu} + R_{\kappa\mu\nu\sigma} + R_{\kappa\nu\sigma\mu} &\equiv 0. \end{aligned}$$

By virtue of these it is possible to show that only 20 of the components of the tensor $R_{\mu\nu\lambda}^\sigma$ are independent (see Ohanian and Ruffini, p. 334ff), hence they may be identified with the 20 non-trivial components of the second derivative of the metric tensor. In fact, up to a constant multiplier $R_{\mu\nu\lambda}^\sigma$ is unique.

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