

## Gravitation and Cosmology

Lecture 26: Dimensional analysis of neutron stars

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# Dimensional analysis of neutron stars

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The equations for the Ricci tensor were

$$R_{tt} = \frac{-B''}{2A} + \frac{B'}{4A} \left( \frac{B'}{B} + \frac{A'}{A} \right) - \frac{1}{r} \frac{B'}{A} = -4\pi G (3p + \rho) B \quad (25.7t)$$

$$R_{rr} = \frac{-B''}{2B} + \frac{B'}{4B} \left( \frac{B'}{B} + \frac{A'}{A} \right) + \frac{1}{r} \frac{A'}{A} = 4\pi G (\rho - p) A \quad (25.7r)$$

$$R_{\theta\theta} = -1 + \frac{r}{2A} \left( \frac{B'}{B} - \frac{A'}{A} \right) + \frac{1}{A} = -4\pi G (\rho - p) r^2 \quad (25.7\theta)$$

$$R_{\varphi\varphi} = \sin^2\theta R_{\theta\theta} \quad (25.7\varphi)$$

### Size and mass

We shall now derive some qualitative results using dimensional analysis. Recall

$$-r^2 \frac{dp}{dr} = GM(r) \rho(r) \left( 1 + \frac{p(r)}{\rho(r)} \right) \left( 1 + \frac{4\pi r^2 p(r)}{M(r)} \right) \left( 1 - \frac{2GM(r)}{r} \right)^{-1} \quad (25.16)$$

and take for the equation of state of a pure neutron gas

$$\rho(r) = \frac{1}{\pi^2} \int_0^{k_F(r)} dk k^2 (k^2 + m^2)^{1/2} \quad (25.17a)$$

$$p(r) = \frac{1}{3\pi^2} \int_0^{k_F(r)} dk k^4 (k^2 + m^2)^{-1/2} \quad (25.17b)$$

We can rewrite Eq. 25.17b as

$$p(r) = \frac{1}{3} \rho(r) - \frac{m^2}{3\pi^2} \int_0^{k_F(r)} dk k^2 (k^2 + m^2)^{-1/2} \quad (26.1)$$

Now change to the dimensionless variable  $\sinh\theta = \frac{k}{m}$ ; let

$$\rho_c = \frac{m^4}{3\pi^2}, \quad \theta_c = \sinh^{-1}(k/m) \quad (26.2)$$

then

$$\rho(r) = 3\rho_c \int_0^{\theta_c} d\theta \cosh^2\theta \sinh^2\theta \quad (26.3a)$$

$$p(r) = \rho_c \int_0^{\theta_c} d\theta \sinh^4\theta. \quad (26.3b)$$

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Rotational frequency

From Eq. 26.3a,b we have

$$p(r) = \rho_c F\left(\frac{\rho(r)}{\rho_c}\right)$$

where  $F(x)$  is some transcendental function.

The dimensional quantities in the theory are therefore  $\rho_c$  and  $2G$ . From them we can construct a radius (recall  $\hbar = c = 1$ )

$$R_c = (2G \rho_c)^{-1/2} \quad (26.4)$$

and a mass

$$M_c = \frac{R_c}{2G} \approx 3.5 M_\odot. \quad (26.5)$$

In general, the mass of the star must be  $M_c$  times a function of the dimensionless ratio  $\frac{\rho(0)}{\rho_c}$ :

$$M = M_c f\left(\frac{\rho(0)}{\rho_c}\right)$$

and the radius of the star must be  $R_c$  times a dimensionless function:

$$R = R_c g\left(\frac{\rho(0)}{\rho_c}\right).$$

### Rotational frequency

Dimensional analysis also gives us a handle on the rotational frequencies of neutron stars: clearly the maximum frequency occurs when the centrifugal acceleration and gravitational acceleration are comparable,  $\frac{GM}{R^2} \approx R\omega^2$ :

$$\omega_{\max} \approx \left(\frac{GM}{R^3}\right)^{1/2} \approx \frac{1}{\sqrt{2}} \frac{c}{R_c} \approx 2 \times 10^4 \text{ sec}^{-1} = \frac{2\pi}{\tau_{\min}}$$

or

$$\tau_{\min} \approx 0.3 \times 10^{-3} \text{ sec};$$

so that the observed pulsars with millisecond periods agree well with this. White dwarf periods, however, are necessarily much longer. Therefore pulsars must be neutron stars.

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### Some features of stellar structure

The total number of neutrons is

$$N = \int d^3r \left( A(r) B(r) \right)^{1/2} J_N^0(r) \quad (26.6)$$

where the (conserved) neutron current is  $J_N^\mu(r)$ . We can define the “proper” number density as

$$n(r) = U_\mu J_N^\mu = \left( B(r) \right)^{1/2} J_N^0(r) \quad (26.7)$$

hence

$$N = 4\pi \int dr r^2 \left( A(r) \right)^{1/2} n(r). \quad (26.8)$$

We can define the energy content of the star as

$$E_{tot} = T + V = M(R) - Nm \quad (26.9)$$

and the local (non-gravitational) energy density as

$$\varepsilon(r) = \rho(r) - m n(r) \quad (26.10)$$

which gives

$$\begin{aligned} E &= 4\pi \int dr r^2 \left[ \rho(r) - m \left( A(r) \right)^{1/2} n(r) \right] \\ &= 4\pi \int dr r^2 \left[ \varepsilon(r) + m n(r) \left( 1 - \left( A(r) \right)^{1/2} \right) \right] \\ &= 4\pi \int dr r^2 \left[ \varepsilon(r) + \left( \rho(r) - \varepsilon(r) \right) \left( 1 - \left( A(r) \right)^{1/2} \right) \right] \\ &= 4\pi \int dr r^2 \left[ \varepsilon(r) \left( A(r) \right)^{1/2} + \rho(r) \left( 1 - \left( A(r) \right)^{1/2} \right) \right]. \end{aligned} \quad (26.11)$$

We can make the identifications

$$T = 4\pi \int dr r^2 \varepsilon(r) \left( A(r) \right)^{1/2} \quad (26.12a)$$

$$V = 4\pi \int dr r^2 \rho(r) \left[ 1 - \left( A(r) \right)^{1/2} \right] \quad (26.12 b)$$

and to leading order we see that

$$T \approx 4\pi \int dr r^2 \varepsilon(r) \quad (26.13 a)$$

and since  $A(r) \approx 1 + \frac{MG}{r}$ ,

$$V = -4\pi \int dr r^2 \rho(r) \frac{GM(r)}{r} \equiv -4\pi \int dM(r) \frac{GM(r)}{r} \quad (26.13 b)$$

Equations 26.13a,b are precisely what we would have written down based on Newtonian mechanics and Newtonian gravitation.

## **Gravitation and Cosmology**

Some features of stellar structure