Coherence Theory: Spatial

Last time, developed theory for incoherent sources

Temporal incoherence:

Wave fluctuates randomly in time

Today, generalize to spatial incoherence:

Wave fluctuates randomly in space

Get spatial fluctuations from extended sources

Outline:
- Review temporal coherence
- Interference with extended sources
- van Cittert-Zernike theorem
- Mutual coherence function
- Michelson interferometer

Material from Hecht Ch 12

Next time: quantum optics

Last class: review for final

Info on final posted on discussion board

Review Temporal Coherence

Random waves characterized by coherence time τ_c

Over times $\ll \tau_c$: oscillations are regular

Over times $\gg \tau_c$: wave fluctuates

Characterize with temporal coherence function

$$\Gamma(\tau) = \langle E(t+\tau)E^*(t)\rangle$$

or complex degree of temporal coherence
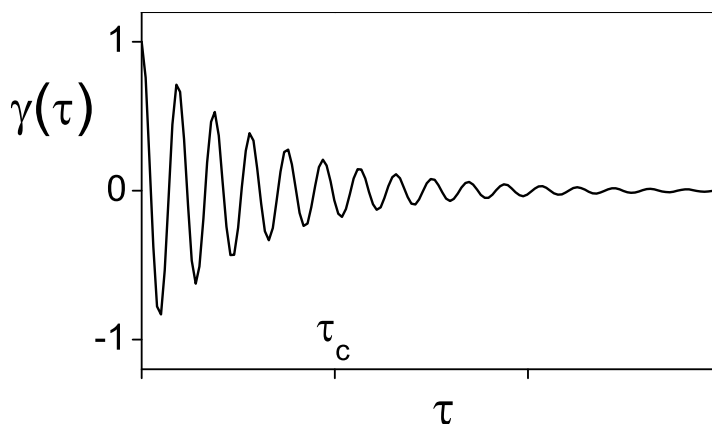
$$\gamma(\tau) = \frac{\Gamma(\tau)}{\Gamma(0)}$$

Have $|\gamma(0)| = 1$

Amplitude decreases over time scale τ_c

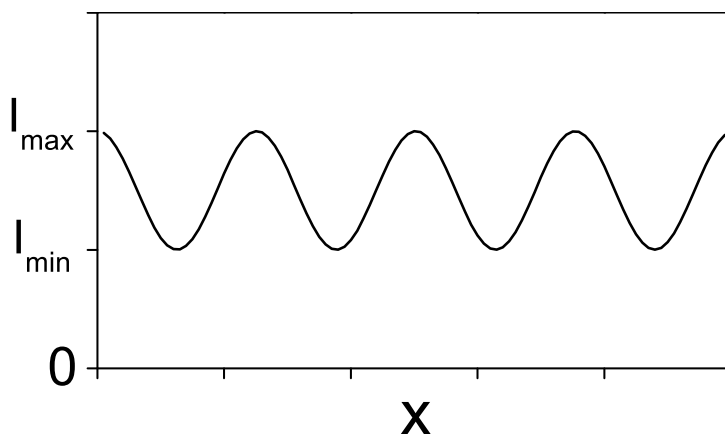Phase of γ oscillates at average frequency ω_0

Typical behavior:

For interferometer with path-length difference h:

Get fringe visibility $\mathcal{V} = \left| \gamma\left(\dfrac{h}{c}\right) \right|$

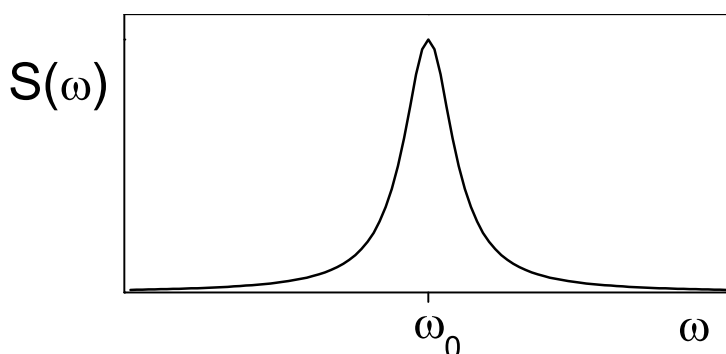where $\mathcal{V} = \dfrac{I_{\text{max}} - I_{\text{min}}}{I_{\text{max}} + I_{\text{min}}}$

Γ also related to spectrum of light

Power spectral density

$$S(\omega) = \frac{1}{2\eta_0} \int_{-\infty}^{\infty} \Gamma(\tau) e^{-i\omega\tau} d\tau$$

Then $S(\omega)\Delta\omega = $ total irradiance
in frequency range ω to $\omega + \Delta\omega$

Extended Sources (Hecht 12.2)

So far, considered waves from point source
   (recall plane wave = point source at ∞)

No problem writing down $E(\mathbf{r})$
Have spherical wave, plane wave, or dipole pattern
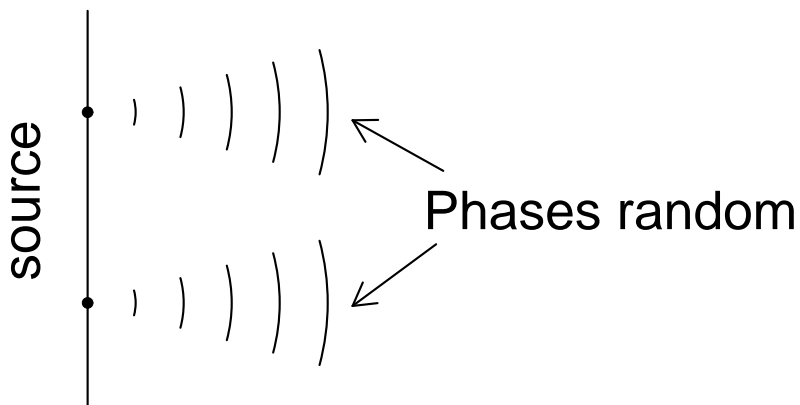
What if source is extended object?
   = collection of many points

Suppose source monochromatic, frequency ω

Still have incoherence:
   phase of source varies from point to point

Phases random

Average over phases:
No interference between fields from different points
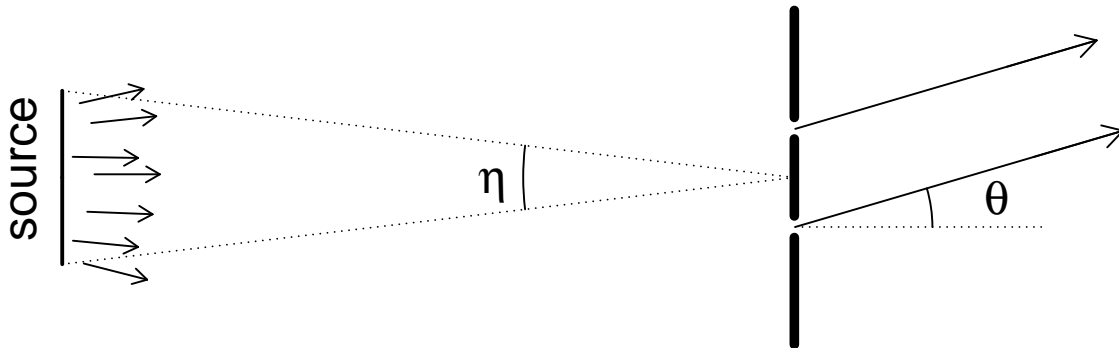
Add irradiances to get total pattern

Question: If the phases are random but constant, don't you get some unpredictable but steady interference pattern?

Reduces visibility of interference

Example: two slit interference

Distant extended source: subtends angle η

Slit spacing a

For point source at normal incidence,
   have interference pattern

$$I(\theta) \approx I_0 \left[1 + \cos(ka\theta)\right]$$

For point source at angle η' have

$$I(\theta; \eta') \approx I_0 \left[1 + \cos ka(\theta - \eta')\right]$$

For extended source, average over η':

$$I_{\text{tot}}(\theta) = \frac{1}{\eta} \int_{-\eta/2}^{\eta/2} I(\theta; \eta')\, d\eta'$$

We can evaluate this:

$$I_{\text{tot}}(\theta) = I_0 \left[1 + \frac{1}{\eta} \int_{-\eta/2}^{\eta/2} \cos ka(\theta - \eta')\, d\eta' \right]$$

Set $u = ka(\theta - \eta')$

$$I_{\text{tot}}(\theta) = I_0 \left[1 + \frac{1}{ka\eta} \int_{ka(\theta-\eta/2)}^{ka(\theta+\eta/2)} \cos u\, du \right]$$

$$= I_0 \left[1 + \frac{\sin ka(\theta + \eta/2) - \sin ka(\theta - \eta/2)}{ka\eta} \right]$$

Expand sines using

$$\sin(\alpha + \beta) = \sin\alpha \cos\beta + \sin\beta \cos\alpha$$
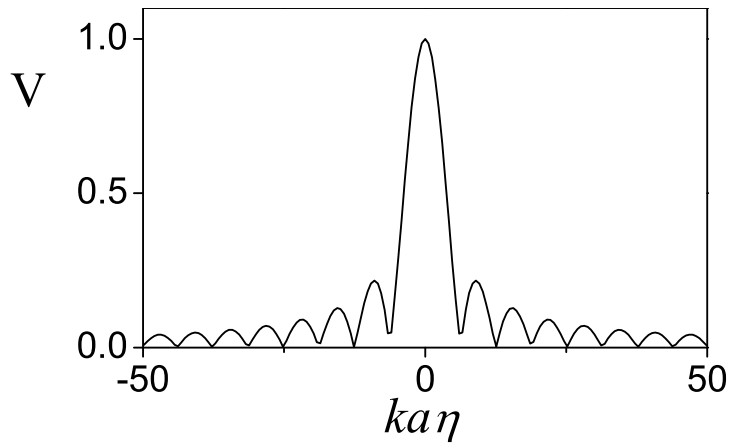
Then $\sin(ka\theta)$ terms cancel, leaves

$$I_{\text{tot}}(\theta) = I_0 \left[1 + \frac{2}{ka\eta} \sin\left(\frac{ka\eta}{2}\right) \cos(ka\theta) \right]$$

$$= I_0 \left[1 + \text{sinc}\left(\frac{ka\eta}{2}\right) \cos(ka\theta) \right]$$

Get visibility

$$\mathcal{V} = \left| \text{sinc}\left(\frac{ka\eta}{2}\right) \right|$$

$$V$$

See that $\mathcal{V} < 1$

Decreases for large η and large a

To have large $\mathcal{V}$, need $a \ll \lambda/\eta$

Leads to idea of spatial coherence:

Field at slit 1 is not completely coherent
    with field at slit 2

    = no definite phase relationship

- Due to indefinite phase between different points
on source

In two-slit example, see that field is coherent
    over distance $\rho_c \approx \lambda/\eta$

For point source $\eta \to 0$ so $\rho_c \to \infty$

Call ρ_c = lateral (or transverse) coherence length

   $a \ll \rho_c$, can treat as point source

   $a \gg \rho_c$, no interference

Generally true that

$$\boxed{\rho_c \approx \frac{\lambda}{\eta}}$$

for source subtending angle η

Question: For a source consisting of two points separated by angle η, the interference pattern is lost when the slit separation $a = \lambda/2\eta$. Is there a simple way to see this?

Example: star α-Centauri

   distance $d = 4.1 \times 10^{16}$ m

   diameter $D = 1.7 \times 10^9$ m

So $\eta = D/d = 4 \times 10^{-8}$ rad

   $\rho_c \approx 12$ m for visible light ($\lambda = 500$ nm)

Could observe interference with slits 12 m apart

   (hard to achieve in practice)

Also:

Require 12 m diameter telescope to resolve disc

   (over distances < 12 m, acts like point source)

Spatial Coherence Function (Hecht 12.3)

General way to calculate visibility:

Suppose interferometer samples source fields
$$E_1 = E(\mathbf{r}_1)$$
$$E_2 = E(\mathbf{r}_2)$$

In two slit interferometer, $\mathbf{r}_1$ and $\mathbf{r}_2$
= positions of slits

Let fields propagate, acquire phases ϕ_1 and ϕ_2
before overlapping

Final field

$$E_{\text{tot}} = E_1 e^{i\phi_1} + E_2 e^{i\phi_2}$$

Interference depends on $\epsilon = \phi_1 - \phi_2$:

$$\left\langle |E_{\text{tot}}|^2 \right\rangle = \left\langle |E_1|^2 \right\rangle + \left\langle |E_2|^2 \right\rangle + \left\langle E_1 E_2^* \right\rangle e^{i\epsilon} + \left\langle E_1^* E_2 \right\rangle e^{-i\epsilon}$$

In two-slit example, $\epsilon = ka\theta$
from extra propagation distance

Define *spatial coherence function*

$$\Gamma_{12} = \left\langle E_1 E_2^* \right\rangle = \left\langle E(\mathbf{r}_1) E^*(\mathbf{r}_2) \right\rangle$$

Analogous to $\Gamma(\tau)$

Here average =
average over random phases of source

Write $\Gamma_{12} = |\Gamma_{12}|e^{i\phi}$:

Then

$$\left\langle |E_{\text{tot}}|^2 \right\rangle = \Gamma_{11} + \Gamma_{22} + |\Gamma_{12}|e^{i(\phi+\epsilon)} + |\Gamma_{12}|e^{-i(\phi+\epsilon)}$$

$$= \Gamma_{11} + \Gamma_{22} + 2|\Gamma_{12}|\cos(\phi+\epsilon)$$

Here $\Gamma_{11} = \left\langle |E_1|^2 \right\rangle = 2\eta_0 I_1$
  similar for Γ_{22}

So Γ_{12} determines interference pattern

For point source, don't need to average:

$$|\Gamma_{12}| = |E_1||E_2| = \sqrt{\Gamma_{11}\Gamma_{22}}$$

So define *complex degree of spatial coherence*

$$\gamma_{12} = \frac{\Gamma_{12}}{\sqrt{\Gamma_{11}\Gamma_{22}}}$$

Have $0 < |\gamma_{12}| < 1$

Visibility of interference pattern is

$$\mathcal{V} = \frac{2\sqrt{I_1 I_2}}{I_1 + I_2}|\gamma_{12}|$$

just as for temporal coherence

van-Cittert–Zernike Theorem (Hecht 12.3.1)

For temporal coherence, can't calculate $\Gamma(\tau)$

But can calculate Γ_{12} from source geometry

Characterize source by brightness $B(X, Y)$

Recall $B = \dfrac{\text{power}}{\text{solid angle} \cdot \text{area}}$ emitted by source

So $B(X, Y) \, dX dY = \text{W/srad}$
 emitted by area $dX \, dY$ on source

Suppose $B(X, Y)$ has transform $\mathcal{B}(k_x, k_y)$

If $\mathbf{r}_1$ and $\mathbf{r}_2$ are a distance d from source
 with $d \gg |\mathbf{r}_1 - \mathbf{r}_2|$ and $d \gg$ source size

Then find

$$\boxed{\gamma_{12} = \frac{1}{\mathcal{L}} \mathcal{B}\left[\frac{k(x_1 - x_2)}{d}, \frac{k(y_1 - y_2)}{d}\right]}$$

with $\mathcal{L} = \mathcal{B}(0, 0) = \iint B(X, Y) \, dX \, dY$

Same as far-field diffraction pattern produced by aperture equal to source

Won't prove, but saw example already:

Line source length b subtends angle $\eta = b/d$

From previous calculation

$$\gamma_{12} = \text{sinc}\left(\frac{ka\eta}{2}\right) = \text{sinc}\left[\frac{k(x_1 - x_2)b}{2d}\right]$$

Same as single-slit diffraction pattern

So circular source diameter D gives

$$\gamma_{12} = \frac{4d}{k\rho D} J_1\left(\frac{k\rho D}{2d}\right)$$

for $\rho = |\mathbf{r}_1 - \mathbf{r}_2|$

Mutual Coherence Function (Hecht 12.3)

In general, both spatial and temporal fluctuations

Define *mutual coherence function*

$$\Gamma_{12}(\tau) = \langle E(\mathbf{r}_1, t + \tau) E^*(\mathbf{r}_2, t)\rangle$$

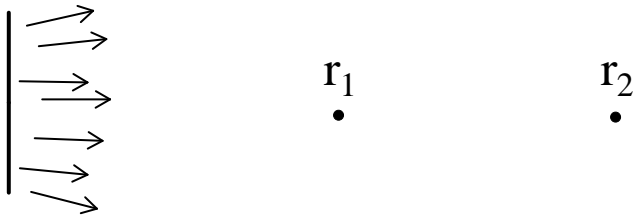Here average over time and source phases

Define *complex degree of coherence*

$$\gamma_{12}(\tau) = \frac{\Gamma_{12}(\tau)}{\sqrt{\Gamma_{11}(0)\Gamma_{22}(0)}}$$

If interferometer samples source field at $\mathbf{r}_1$, $\mathbf{r}_2$
and interfers fields with time delay τ:

Get visibility $\mathcal{V} = |\gamma_{12}(\tau)|$
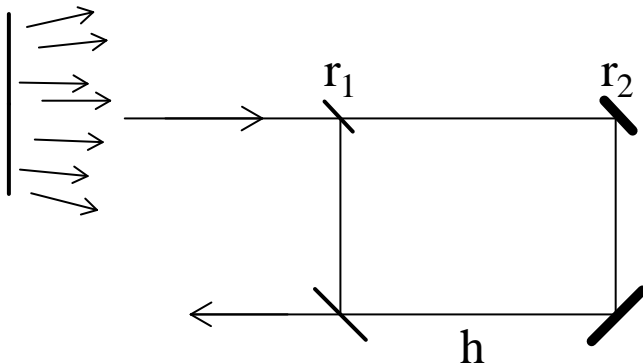
Note that space and time coherence intermixed

Here $\mathbf{r}_1$ sees field at earlier time that $\mathbf{r}_2$
temporal fluctuations contribute to $\gamma_{12}(0)$

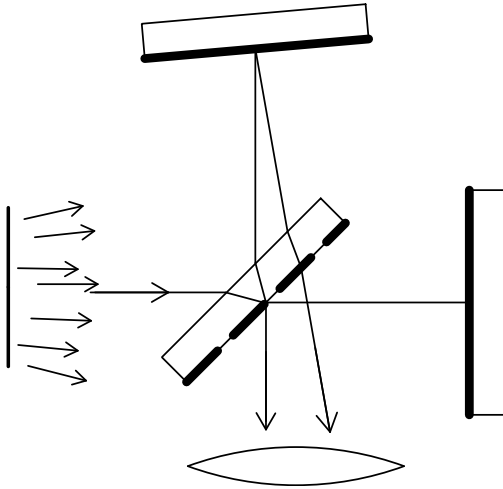Use τ only for time delay after sampling field

Example:

Here want $\gamma_{12}(\tau)$ for $\tau = h/c$

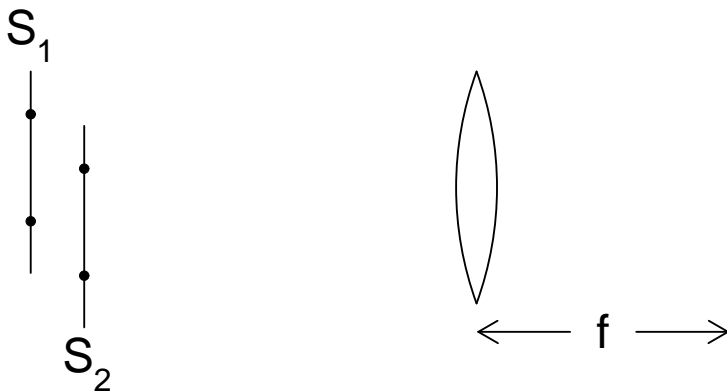Note van Cittert-Zernike theorem only applies to
nearly monochromatic source

Michelson Interferometer

Michelson is special case:

    spatial coherence unnecessary

To analyze, look at virtual sources produced by mirrors:

S_1

S_2

$\longleftarrow f \longrightarrow$
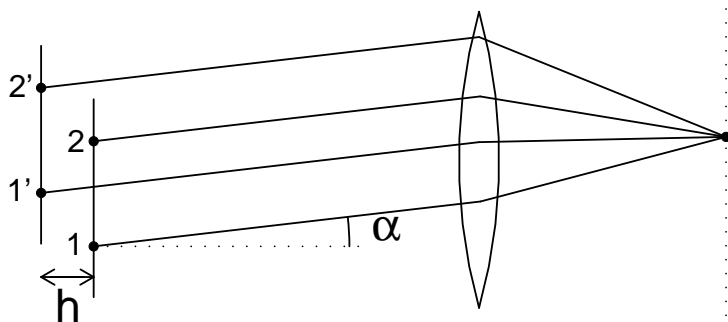
Sources separated by path length difference h

And displaced by θd

    d = arm length, θ = mirror tilt

Detect at focal plane of lens

All rays at angle α imaged to point $\mathbf{r}$

Light from $\mathbf{r}_1$ interfers only with light from $\mathbf{r}'_1$

Path difference 11' $= h/\cos\alpha$

Same as path difference 22'

So light from all points reaching $\mathbf{r}$ has same phase

See interference pattern as function of α

In Michelson demo:
   source = frosted glass illuminated with laser
   lens = camera lens
   detector = CCD

Note, no interference observed without lens

Lens is good trick to recover interference

Summary:

- Extended source usually decreases $\mathcal{V}$

- Characterize spatial coherence with Γ_{12}
$$\Gamma_{12} = \langle E(\mathbf{r}_1)E(\mathbf{r}_2)^* \rangle$$

- Lateral coherence length $\rho_c \approx \lambda/\eta$
η = angle subtended by source

- Get Γ_{12} from van Cittert-Zernike

- Both time and space coherence: $\Gamma_{12}(\tau)$

- Michelson works with extended source
when using lens