## Ray Optics II

Last time, developed idea of ray optics approximation to wave theory

Introduced paraxial approximation:
rays with $\theta \ll 1$
Will continue to use

Started disussing imaging and lenses:
Thin lens equation $\frac{1}{s_{o}}+\frac{1}{s_{i}}=\frac{1}{f}$
Basic equation of paraxial optics

Example: Suppose a thin lens of focal length $f=-100 \mathrm{~cm}$ is placed 50 cm in front of a small light bulb. Where will the image of the bulb be located?


## Solution:

Real object upstream of lens, so $s_{o}=+50 \mathrm{~cm}$.
Then $\frac{1}{s_{i}}=\frac{1}{f}-\frac{1}{s_{o}}=\frac{1}{(-100 \mathrm{~cm})}-\frac{1}{50 \mathrm{~cm}}=-0.03 \mathrm{~cm}^{-1}$
So $s_{i}=-33.3 \mathrm{~cm}$
Negative, so image located 33.3 cm before lens

$$
=16.7 \mathrm{~cm} \text { from object }
$$

Today: continue with imaging

- Imaging extended objects
- Multiple Ienses
- Mirrors
- Apertures

Next time:
Techniques for dealing with multi-lens systems

Finite Imaging (Hecht 5.2.3)
Before, all pictures showed point-to-point imaging


Points are on optic axis
$=$ symmetry axis of lens

Finite object $=$ collection of points must deal with points that are off axis

Picture:


Construct image using ray diagram
Three simple rays:

- ray through lens center: undeviated
- ray through front focal point:
becomes horizontal
- horizontal ray: hits back focal point


Image where rays intersect

Each point in object plane $\rightarrow$ point in image plane Here, image inverted:
object height $y_{o}>0$
maps to image height $y_{i}<0$
Define magnification $m=\frac{y_{i}}{y_{o}}$

Get magnification from diagram:


Triangles are similar, so $\frac{s_{i}}{y_{i}}=-\frac{s_{o}}{y_{o}}$
Then $m=\frac{y_{i}}{y_{o}}=-\frac{s_{i}}{s_{o}}$
with $s_{i}$ determined by thin lens equation

For negative parameters, follow sign conventions:


Since $f<0$, "front" and "back" focal points reversed

See that $s_{i}<0$ and $m>0$

Ray diagrams good tool for understanding

## Lens systems

If more than one lens:
apply thin lens law in succession
Consider two lenses $f_{1}, f_{2}$, separated by $d$


Find image distance $s_{i}$

First lens: image distance $s_{i 1}$

$$
\frac{1}{s_{i 1}}=\frac{1}{f_{1}}-\frac{1}{s_{o}}
$$

Second lens: object distance $s_{o 2}=d-s_{i 1}$
Then final image at $s_{i}: \frac{1}{s_{i}}=\frac{1}{f_{2}}-\frac{1}{s_{o 2}}$
Combine, get

$$
s_{i}=\frac{f_{2} d s_{o}-f_{1} f_{2} d-f_{1} f_{2} s_{o}}{s_{o} d-s_{o} f_{1}-s_{o} f_{2}-d f_{1}+f_{1} f_{2}}
$$

Not very enlightening

Simple case $d \rightarrow 0$ :

$$
s_{o 2}=-s_{i 1}
$$

Then $\frac{1}{s_{i}}=\frac{1}{f_{2}}+\frac{1}{s_{i 1}}=\frac{1}{f_{2}}+\left(\frac{1}{f_{1}}-\frac{1}{s_{o}}\right)$
So

$$
\frac{1}{s_{O}}+\frac{1}{s_{i}}=\frac{1}{f_{1}}+\frac{1}{f_{2}}
$$

like single lens with $\frac{1}{f_{\mathrm{tot}}}=\frac{1}{f_{1}}+\frac{1}{f_{2}}$

Define $\frac{1}{f}=$ power of lens
units: diopters $\left(\equiv \mathrm{m}^{-1}\right)$
Lens powers add for adjacent lenses
Question: My eyeglass prescription is -4 diopters. What is the focal length of my lenses?

Next time, see better ways to treat lens systems For now, move on to other elements

Mirrors (Hecht 5.4)
Flat mirrors just deflect rays


Makes ray tracing hard

Best approach:
(1) Use law of reflection to find path of optic axis
(2) Trace rays as if axis were straight

(3) Result correct relative to axis

Mirrors can be curved too act like a lens

Get perfect imaging with aspheric mirror


Parabola $=$ points equidistant from focus and line so $\mathcal{S}=$ constant for all rays

Aspheric surfaces still expensive:
spherical mirrors more common
work for paraxial rays

Analyze with law of reflection or Fermat's principle Result:

$$
\frac{1}{s_{o}}+\frac{1}{s_{i}}=-\frac{2}{R}
$$

Just like thin lens: $f \rightarrow-R / 2$


Here $R<0$ : like $f>0$, get real image But optic axis also folded

Lens equivalent:


Trick to remember factor of 2 :
if object at center of curvature, so is image


Then $s_{o}=s_{i}=-R \quad$ (since $R<0$ here)
and

$$
\frac{1}{f}=\frac{1}{s_{o}}+\frac{1}{s_{i}}=-\frac{2}{R}
$$

Question: Could you tilt a curved mirror and use it to both fold axis and focus rays?


Stops
stop $=$ aperture $=$ hole
Formally:
stop $=$ object that limits rays reaching image


Lens edge always a stop
often introduce additional ones

Stops important for two questions:

- How bright will image be?
$=$ how much light is collected
- What is field of view?
$=$ what area of object is imaged

Both important for real system design

Assume points close to optic axis off-axis points: vignetting (Hecht pgs. 1723)

Image Brightness:
Generally many stops in system
Define aperture stop ( $\equiv \mathrm{AS}$ )
$=$ the limiting stop


In complicated system, trace rays to determine

Aperture stop sets amount of light collected

Suppose AS = first element
then only rays hitting element are imaged
Defines acceptance angle $\theta_{\text {max }}$


Characterize source by brightness $B$

$$
B=\frac{\text { power }}{(\text { source area)(solid angle) }}
$$

units $W /\left(m^{2} \mathrm{sr}\right)$
Question: What's a steradian (sr)?

Typical light bulb:
$P=100 \mathrm{~W}$
surface area $=4 \pi \times(3 \mathrm{~cm})^{2}$
radiates into $4 \pi \mathrm{sr}$
So $B \approx 700 \mathrm{~W} /\left(\mathrm{m}^{2} \mathrm{sr}\right)$

System accepts solid angle $\pi \theta_{\text {max }}^{2}$
Power into system $=B \times($ source area $) \times \pi \theta_{\text {max }}^{2}$
Gives image intensity

What if first element $\neq$ AS?
Define entrance pupil
$=$ image of AS from object
System acts like first element is entrance pupil

Example:


Say $f=10 \mathrm{~cm}$, stop radius $=0.5 \mathrm{~cm}$

Find image of stop through lens:

$$
s_{o}=5 \mathrm{~cm} \Rightarrow s_{i}=-10 \mathrm{~cm}
$$

Magnification $m=-s_{i} / s_{o}=2$

So entrance pupil located 10 cm behind lens
radius $=1 \mathrm{~cm}$


Defines rays accepted by system here $\theta_{\max }=\tan ^{-1}(1 / 30) \approx 0.033 \mathrm{rad}$

Similarly, define exit pupil
$=$ image of AS seen from image of system


Exit pupil acts like "window" all rays from object pass through window

Exit pupil important for systems viewed by eye microscope, telescope, binoculars, etc.

Exit pupil like window

- if pupil small, far away:
observe small disk surrounded by darkness
See in cheap binoculars
- if pupil large, close to eye: many rays don't enter eye ... wasted money

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Ideal system: exit pupil aligned to pupil of eye

- 3-5 mm diameter
- located about 1 cm behind eyepiece
$\binom{$ Eyepiece $=$ last element before eye }{ Distance eyepiece to exit pupil $=$ eye relief }
Good binoculars:
- image fills view
- comfortable to look through

Generally: when combining two systems, make exit pupil of first $=$ entrance pupil of second

Still have second question: what is field of view?

Define field stop (FS) $=$
stop that limits which object points are imaged


Often, field stop $=$ edge of detector (film, CCD sensor, retina of eye, etc.)

But not always:


Good field stop always in an image plane (final or intermediate)

Field stops useful for non-imaging detectors

- photodiodes, PMT, bolometers Use field stop to eliminate background light

Question: Photographers use a stop to set the exposure level of a camera. Does this refer to the field stop or the aperture stop?

Also, both AS and FS impact aberrations (imaging errors)

Idea: use stops to block non-paraxial rays

Last element: prisms comment briefly

Two uses: reflection and dispersion
Reflecting: either TIR or mirror coatings nice way to hold mirrors close together


See Hecht 5.5.2 for more

Dispersing prisms: familiar rainbow effect Use index $n=n(\omega)$

Snell's law $\Rightarrow \theta=\theta(\omega)$


Analysis straighforward (Hecht 5.5.1)
More often use diffraction gratings

Summary:

- Ray diagrams help analyze off-axis imaging
- Analyze multiple lenses in sequence
- Curved mirrors act much like lenses
- Stop $=$ limiting edge
- Aperture stop sets brightness of image
- Field stop sets field of view

