Phys 531 Lecture 9 30 September 2004 Ray Optics II

Last time, developed idea of ray optics approximation to wave theory

Introduced paraxial approximation: rays with $\theta \ll 1$ Will continue to use

Started disussing imaging and lenses:

Thin lens equation $\frac{1}{s_o} + \frac{1}{s_i} = \frac{1}{f}$ Basic equation of paraxial optics

Example: Suppose a thin lens of focal length f = -100 cm is placed 50 cm in front of a small light bulb. Where will the image of the bulb be located?



Solution:

Real object upstream of lens, so $s_o = +50$ cm.

Then $\frac{1}{s_i} = \frac{1}{f} - \frac{1}{s_o} = \frac{1}{(-100 \text{ cm})} - \frac{1}{50 \text{ cm}} = -0.03 \text{ cm}^{-1}$ So $s_i = -33.3$ cm

Negative, so image located 33.3 cm before lens

= 16.7 cm from object

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Today: continue with imaging

- Imaging extended objects
- Multiple lenses
- Mirrors
- Apertures

Next time: Techniques for dealing with multi-lens systems

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Finite Imaging (Hecht 5.2.3)

Before, all pictures showed point-to-point imaging



Points are on *optic axis* = symmetry axis of lens

Finite object = collection of points must deal with points that are off axis Picture:



Construct image using ray diagram

Three simple rays:

- ray through lens center: undeviated
- ray through front focal point: becomes horizontal
- horizontal ray: hits back focal point



Image where rays intersect

Each point in object plane \rightarrow point in image plane

Here, image inverted: object height $y_o > 0$ maps to image height $y_i < 0$ Define magnification $m = \frac{y_i}{y_o}$ 5

Get magnification from diagram:



Triangles are similar, so $\frac{s_i}{y_i} = -\frac{s_o}{y_o}$

Then $m = \frac{y_i}{y_o} = -\frac{s_i}{s_o}$ with s_i determined by thin lens equation

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For negative parameters, follow sign conventions:



Since f < 0, "front" and "back" focal points reversed

See that $s_i < 0$ and m > 0

Ray diagrams good tool for understanding

Lens systems

If more than one lens:

apply thin lens law in succession

Consider two lenses f_1 , f_2 , separated by d



Find image distance s_i

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First lens: image distance s_{i1}

$$\frac{1}{s_{i1}} = \frac{1}{f_1} - \frac{1}{s_o}$$

Second lens: object distance $s_{o2} = d - s_{i1}$

Then final image at s_i : $\frac{1}{s_i} = \frac{1}{f_2} - \frac{1}{s_{o2}}$

Combine, get

$$s_i = \frac{f_2 ds_o - f_1 f_2 d - f_1 f_2 s_o}{s_o d - s_o f_1 - s_o f_2 - df_1 + f_1 f_2}$$

Not very enlightening

Simple case $d \rightarrow 0$: $s_{o2} = -s_{i1}$ Then $\frac{1}{s_i} = \frac{1}{f_2} + \frac{1}{s_{i1}} = \frac{1}{f_2} + \left(\frac{1}{f_1} - \frac{1}{s_o}\right)$ So $\frac{1}{s_o} + \frac{1}{s_i} = \frac{1}{f_1} + \frac{1}{f_2}$ like single lens with $\frac{1}{f_{tot}} = \frac{1}{f_1} + \frac{1}{f_2}$

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т.	ж.

Define $\frac{1}{f} = power of lens$ units: diopters ($\equiv m^{-1}$)

Lens powers add for adjacent lenses

Question: My eyeglass prescription is -4 diopters. What is the focal length of my lenses?

Next time, see better ways to treat lens systems For now, move on to other elements Mirrors (Hecht 5.4)

Flat mirrors just deflect rays



Makes ray tracing hard



(3) Result correct relative to axis

Mirrors can be curved too

act like a lens

Get perfect imaging with aspheric mirror



Parabola = points equidistant from focus and line so \mathcal{S} = constant for all rays

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Aspheric surfaces still expensive: spherical mirrors more common work for paraxial rays

Analyze with law of reflection or Fermat's principle

Result:

$$\frac{1}{s_o} + \frac{1}{s_i} = -\frac{2}{R}$$

Just like thin lens: $f \rightarrow -R/2$



Here R < 0: like f > 0, get real image But optic axis also folded

Lens equivalent:



Trick to remember factor of 2:

if object at center of curvature, so is image



Then $s_o = s_i = -R$ (since R < 0 here)

and

$$\frac{1}{f} = \frac{1}{s_o} + \frac{1}{s_i} = -\frac{2}{R}$$

Question: Could you tilt a curved mirror and use it to both fold axis and focus rays?



Stops

stop = aperture = hole

Formally:

stop = object that limits rays reaching image



Lens edge always a stop often introduce additional ones Stops important for two questions:

- How bright will image be?
 how much light is collected
- What is field of view?

= what area of object is imaged

Both important for real system design

Assume points close to optic axis

off-axis points: vignetting (Hecht pgs. 172-3)

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Image Brightness:

Generally many stops in system

Define aperture stop (\equiv AS)

= the limiting stop



In complicated system, trace rays to determine

Aperture stop sets amount of light collected

Suppose AS = first element

then only rays hitting element are imaged

Defines acceptance angle θ_{max}



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Characterize source by brightness B

 $B = \frac{\text{power}}{(\text{source area})(\text{solid angle})}$ units W/(m² sr)

Question: What's a steradian (sr)?

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Typical light bulb:

P = 100 \text{ W}

surface area = 4\pi \times (3 \text{ cm})^2

radiates into 4\pi \text{ sr}

So B \approx 700 \text{ W/(m}^2 \text{ sr})
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System accepts solid angle $\pi \theta_{\max}^2$

Power into system = $B \times (\text{source area}) \times \pi \theta_{\text{max}}^2$

Gives image intensity

What if first element \neq AS?

Define *entrance pupil* = image of AS from object

System acts like first element is entrance pupil

Example:



Say f = 10 cm, stop radius = 0.5 cm

Find image of stop through lens: $s_o = 5 \text{ cm} \Rightarrow s_i = -10 \text{ cm}$

Magnification $m = -s_i/s_o = 2$

So entrance pupil located 10 cm behind lens radius = 1 cm



Defines rays accepted by system here $\theta_{\rm max} = {\rm tan}^{-1}(1/30) \approx 0.033$ rad

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Similarly, define exit pupil

= image of AS seen from image of system



Exit pupil acts like "window" all rays from object pass through window Exit pupil important for systems viewed by eye microscope, telescope, binoculars, etc.

Exit pupil like window

 if pupil small, far away: observe small disk surrounded by darkness
 See in cheap binoculars

• if pupil large, close to eye: many rays don't enter eye ... wasted money

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Ideal system: exit pupil aligned to pupil of eye

- 3-5 mm diameter
- located about 1 cm behind eyepiece

Eyepiece = last element before eye
 Distance eyepiece to exit pupil = eye relief

Good binoculars:

- image fills view
- comfortable to look through

Generally: when combining two systems, make exit pupil of first = entrance pupil of second

Still have second question: what is field of view?

Define field stop (FS) =stop that limits which object points are imaged





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But not always:



Good field stop always in an image plane (final or intermediate)

Field stops useful for non-imaging detectors

- photodiodes, PMT, bolometers Use field stop to eliminate background light

Question: Photographers use a stop to set the exposure level of a camera. Does this refer to the field stop or the aperture stop?

Also, both AS and FS impact aberrations (imaging errors)

Idea: use stops to block non-paraxial rays

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Last element: prisms comment briefly

Two uses: reflection and dispersion

Reflecting: either TIR or mirror coatings nice way to hold mirrors close together



See Hecht 5.5.2 for more

Dispersing prisms: familiar rainbow effect

Use index $n = n(\omega)$

Snell's law $\Rightarrow \theta = \theta(\omega)$



Analysis straighforward (Hecht 5.5.1)

More often use diffraction gratings

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Summary:

- Ray diagrams help analyze off-axis imaging
- Analyze multiple lenses in sequence
- Curved mirrors act much like lenses
- Stop = limiting edge
- Aperture stop sets brightness of image
- Field stop sets field of view