Phys 531 Lecture 21 3 November 2005
Coherence Theory: Temporal
Last time, discussed interferometers Michelson Fabry-Perot
Mostly considered monochromatic light makes interference easy
Can also use incoherent light sometimes necessary
Today, discuss incoherent sources be more quantitative than before

Outline:

- Random waves
- Temporal coherence function
- Applications to interference
- Power spectral density

More Fourier transforms today!

Developing toward Hecht Chapter 12 Material from Ch 7, 9, 11

In Ch 12, Hecht focuses on spatial coherence: We'll cover in next lecture

Today, apply same ideas to temporal coherence Easier way to start Random Waves (Hecht 7.4.3)

Most light sources produce wave that fluctuates

Then E(t) varies \sim randomly in time

Example:



Why?

Source composed of many atoms

= many radiators

Quantum mechanics:

Each atom excited in discrete steps Radiates briefly, then stops (until excited again)



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Typical atomic decay time = 10 ns

Makes field that is *coherent* over 10 ns timescale

Recall coherent = oscillating with constant phase

Define coherence time τ_c = time over which wave is coherent

Question: What would be the coherence time of a pulsed laser?

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Note 10 ns $\approx 10^7$ optical periods Atomic radiation is rather coherent

Most thermal sources not that good: Examples: light bulb, candle, sunlight
Atoms constantly collide with neighbors

interrupt phase of oscillation

Typical coherence time = 2-3 fs

1-2 optical periods

Want to understand effect on interference

Random Waves (Hecht 12.3)

How to treat mathematically?

Don't work with E(t) directly

Assume we have sample of "possible" E(t)'s

Work with averages

 $\langle .. \rangle$ = average over sample

Imagine running experiment many times, collecting data $E_1(t)$, $E_2(t)$, $E_3(t)$...

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For instance: $\langle E(t) \rangle =$ average possible values of E at time t= 0 for truly random wave

But might have $E(t) = \overline{E}(t) + \delta E(t)$ $\overline{E}(t) = \text{non-random} = deterministic$ $\delta E(t) = \text{random noise}$

Then $\langle E(t) \rangle = \overline{E}(t)$ = deterministic part

Already know how deterministic part works For today, assume $\langle E(t) \rangle = 0$

Also assume that averages are independent of time \Rightarrow Fluctuations in E have constant character Say that E(t) is *stationary*

For stationary wave,

can record samples E(t) sequentially in time Sample length T, get N samples in time NT

Then $\langle .. \rangle$ equivalent to time average

$$\langle f \rangle = \lim_{T \to \infty} \frac{1}{T} \int_{-T/2}^{T/2} f(t) dt$$

What can we measure?

Know irradiance \neq 0:

$$I = \frac{1}{2\eta_0} \left\langle |E|^2 \right\rangle$$

For stationary wave, I is constant

Same whether light is coherent or not

Really want to know how E(t) compares to $E(t+\tau)$ Tells how correlations decay in time

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Define temporal coherence function

 $\Gamma(\tau) = \langle E(t+\tau) E^*(t) \rangle$

Have $\Gamma(0) \equiv 2\eta_0 I$

- has irradiance information

Also has coherence information: If $\tau \gg \tau_c$, then $E(t + \tau)$ independent of E(t) $\Rightarrow \Gamma(\tau) = 0$ If $\tau \ll \tau_c$, then $E(t + \tau)$ determined by E(t)

 $\Rightarrow \Gamma(\tau)$ large

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 $\Gamma(\tau)$ gives precise measure of coherence Examples:

• Monochromatic wave $E(t) = E_0 e^{-i\omega_0 t}$

$$\Gamma(\tau) = \left\langle E_0 e^{-i\omega_0(t+\tau)} E_0^* e^{i\omega_0 t} \right\rangle$$
$$= |E_0|^2 e^{-i\omega_0 \tau}$$

Oscillates at ω_0

Magnitude = $|E_0|^2$ for all τ

- perfectly coherent wave

• Atomic radiation, frequency ω_0 :

 $\Gamma(\tau) = |E_0|^2 e^{-|\tau|/\tau_c} e^{-i\omega_0 \tau}$

exponential decay, time constant au_c

Question: Is it possible to have $|\Gamma(\tau)| > \Gamma(0)$?

Generally, can't calculate $\Gamma(\tau)$ in optics Need to know about physics of source \rightarrow Usually quantum mechanics

Can measure for given source

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Often use normalized version of Γ

$$\gamma(\tau) = \frac{\Gamma(\tau)}{\Gamma(0)}$$

so γ independent of irradiance

Called *complex degree of temporal coherence* (pretty dumb name)

For atomic radiation

 $\gamma(\tau) = e^{-|\tau|/\tau_c} e^{-i\omega_0 \tau}$

Interference (Hecht 9.2, 9.3)

Use $\Gamma(\tau)$ to analyze interference



Arm length difference = h

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Suppose source = random wave $E_0(t)$

Let length of arm 1 = d

Then output $E_1(t) = \beta E_0(t - d/c)$

- transmission factor β
- time delay d/c

Length of arm 2 = d + h

And output $E_2(t) = \beta E_0(t - d/c - h/c)$

Output irradiance given by

$$|E(t)|^{2} = |E_{1}(t) + E_{2}(t)|^{2}$$
$$= |E_{1}(t)|^{2} + |E_{2}(t)|^{2}$$
$$+ E_{1}^{*}(t)E_{2}(t) + E_{1}(t)E_{2}^{*}(t)$$

But only want to look at averages:

$$\left\langle |E|^2 \right\rangle = \left\langle |E_1|^2 \right\rangle + \left\langle |E_2|^2 \right\rangle + \left\langle E_1^* E_2 \right\rangle + \left\langle E_1 E_2^* \right\rangle$$

Have $\left\langle |E_1|^2 \right\rangle = \left\langle |E_2|^2 \right\rangle = |\beta|^2 \left\langle |E_0|^2 \right\rangle$
$$= |\beta|^2 \Gamma(0)$$

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For interference terms, have

$$\langle E_1(t)E_2^*(t)\rangle = |\beta|^2 \langle E_0(t-d/c)E_0^*(t-d/c-h/c)\rangle$$

With stationary wave can rearrange times:

$$\langle E_0(t - d/c)E_0^*(t - d/c - h/c) \rangle = \langle E_0(t + h/c)E_0^*(t) \rangle$$
$$= \Gamma(\tau)$$

for $\tau = h/c =$ time delay

Also $\langle E_1^*(t)E_2(t)\rangle = |\beta|^2\Gamma^*(\tau)$

So total output is

$$\left\langle |E|^2 \right\rangle = |\beta|^2 \left[2\Gamma(0) + \Gamma(h/c) + \Gamma^*(h/c) \right]$$

Suppose $\Gamma(\tau) = \Gamma(0)e^{-|\tau|/\tau_c}e^{-i\omega_0 t}$

Define $\ell_c = c\tau_c = longitudinal$ coherence length

Then

$$\begin{split} \left\langle |E|^2 \right\rangle &= |\beta|^2 \Gamma(0) \Big[2 + e^{-|h/\ell_c|} e^{-i\omega_0 h/c} + e^{-|h/\ell_c|} e^{i\omega_0 h/c} \Big] \\ &= 2|\beta|^2 \Gamma(0) \left[1 + e^{-|h/\ell_c|} \cos\left(\frac{\omega_0 h}{c}\right) \right] \end{split}$$

See oscillation with h: interference

- but amplitude decays for $|h|\gtrsim\ell_c$

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Define visibility of interference pattern

$$\mathcal{V} = \frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}}$$

For perfect fringe, $I_{min} = 0 \Rightarrow \mathcal{V} = 1$

Generally good measure of fringe contrast In our example,

 $I_{\max} = \frac{|\beta|^2 \Gamma(0)}{\eta_0} \left(1 + e^{-|ch/\tau_c|} \right)$ $I_{\min} = \frac{|\beta|^2 \Gamma(0)}{\eta_0} \left(1 - e^{-|ch/\tau_c|} \right)$

So
$$V = \frac{2e^{-|h/\ell_c|}}{2} = e^{-|h/\ell_c|}$$

General result:

 $\mathcal{V} = |\gamma(\tau)|$

when interfering waves with time delay $\boldsymbol{\tau}$

If amplitudes of E_1 and E_2 are different, get

$$\mathcal{V} = \frac{2\sqrt{I_1 I_2}}{I_1 + I_2} |\gamma(\tau)|$$

Question: Is \mathcal{V} higher or lower if $I_1 \neq I_2$?

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Michelson interferometer

= good way to measure γ

Demo: Michelson with white light

Another example: Two slit interference



Interference at angle θ : path length difference = $a\theta$ Time delay = $a\theta/c$ So fringe visibility decays as $|\gamma(a\theta/c)|$ If coherence time τ_c , need $|a\theta/c| < \tau_c$ limits $|\theta| < \ell_c/a$

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For light bulb, $\tau_c = 2$ fs $\Rightarrow \ell_c = 0.6 \ \mu \text{m}$

If $a = 100 \ \mu\text{m}$, need $|\theta| < 6 \ \text{mrad} = 0.3^{\circ}$

But fringe spacing $\Delta \theta = \lambda/a = 5$ mrad Only observe about one fringe

Generally limited to $N \approx \frac{\ell_c}{\lambda}$ fringes

Hard to see interference with natural light

Power Spectral Density (Hecht 11.3.4)

 Γ also useful for characterizing spectrum of light

Idea of spectrum:

Polychromatic light has range of frequencies

Want to characterize by $I(\omega)$

- irradiance as function of frequency

For instance: Pass light through filter for freq ω_0 , frequency width $\Delta \omega$

Expect to transmit irradiance $I(\omega_0)\Delta\omega$

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Obvious approach: Fourier transform

Try to define

$$I(\omega) = \frac{1}{2\eta_0} |\mathcal{E}(\omega)|^2$$

for $\mathcal{E} = \text{transform of } E(t)$

Doesn't work - several reasons

• Units wrong:

 $\mathcal{E}(\omega)$ units Vs/m So $I(\omega)$ units (W s²)/m² Want units (W s)/m² so that $I(\omega)\Delta\omega = W/m^2$ • Bad for monochromatic light:

If
$$E(t) = E_0 e^{-i\omega_0 t}$$
 then
 $\mathcal{E}(\omega) = 2\pi E_0 \delta(\omega - \omega_0)$
 $I(\omega) = \frac{2\pi^2 |E_0|^2}{\eta_0} \delta(\omega - \omega_0)^2$
 $\delta()^2$ is nasty

• Bad for random light:

$$E(\omega) = \int_{-\infty}^{\infty} E(t) e^{i\omega t} dt$$

Need to average $\langle .. \rangle$, don't know how

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Attack averaging problem (solves others as well)

Define
$$\mathcal{E}_T(\omega) = \int_{-T/2}^{T/2} E(t) e^{i\omega t} dt$$

Then $|\mathcal{E}_T|^2 \Delta \omega / 2\eta_0 = \text{energy}/\text{m}^2$ in time T (Parseval's theorem)

Want power = average energy/time

Define
$$S(\omega) = \lim_{T \to \infty} \frac{|\mathcal{E}_T(\omega)|^2}{2\eta_0 T}$$

Call $S(\omega) = power spectral density$

Write out limit

$$\lim_{T \to \infty} \frac{1}{T} \int_{-T/2}^{T/2} E^*(t) e^{-i\omega t} dt \int_{-T/2}^{T/2} E(t') e^{i\omega t'} dt'$$

Change variables to (t,τ) with $t' = t + \tau$

 $\lim_{T \to \infty} \frac{1}{T} \int_{-T/2}^{T/2} \int_{-T/2-t}^{T/2-t} E(t+\tau) E^*(t) e^{-i\omega t} e^{i\omega(t+\tau)} d\tau dt$

$$= \lim_{T \to \infty} \frac{1}{T} \int_{-T/2}^{T/2} \int_{-T/2-t}^{T/2-t} E(t+\tau) E^*(t) e^{i\omega\tau} d\tau dt$$

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Move τ integral outside limit

$$= \int_{-\infty}^{\infty} e^{i\omega\tau} \left[\lim_{T \to \infty} \frac{1}{T} \int_{-T/2}^{T/2} E(t+\tau) E^*(t) dt \right] d\tau$$

Recognize expression in brackets:

$$\lim_{T \to \infty} \frac{1}{T} \int_{-T/2}^{T/2} E(t+\tau) E^*(t) dt = \langle E(t+\tau) E^*(t) \rangle$$
$$= \Gamma(\tau)$$

So
$$S(\omega) = \frac{1}{2\eta_0} \int_{-\infty}^{\infty} \Gamma(\tau) e^{i\omega\tau} d\tau$$

Power spectral density =Fourier transform of temporal coherence function

Called the Wiener-Khintchine theorem

Extremely useful, not just in optics

For instance:

Calculate spectrum of electronic noise Or spectrum of stock market fluctuations

Good way to characterize any noisy signal

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Examples:

• Monochromatic field $E(t) = E_0 e^{-i\omega_0 t}$

Saw already $\Gamma(\tau) = |E_0|^2 e^{-i\omega_0 \tau}$

So
$$S(\omega) = \frac{|E_0|^2}{2\eta_0} \int_{-\infty}^{\infty} e^{i(\omega-\omega_0)t} dt$$
$$= \frac{\pi |E_0|^2}{\eta_0} \delta(\omega - \omega_0)$$

 δ -peak at $\omega = \omega_0$: makes sense, monochromatic

Total irradiance is

$$I = \frac{1}{2\pi} \int_{-\infty}^{\infty} S(\omega) d\omega$$

Here

$$I = \frac{|E_0|^2}{2\eta_0} \int_{-\infty}^{\infty} \delta(\omega - \omega_0) d\omega$$
$$= \frac{|E_0|^2}{2\eta_0}$$

using $\int \delta(\omega) d\omega = 1$

Get expected result

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• Atomic radiation

Have $\Gamma(\tau) = \Gamma(0)e^{-|\tau|/\tau_c}e^{-i\omega_0\tau}$

$$S(\omega) = \frac{\Gamma(0)}{2\eta_0} \int_{-\infty}^{\infty} e^{-|\tau|/\tau_c} e^{i(\omega-\omega_0)\tau} d\tau$$

Did this already, HW problem 6.4

Get
$$S(\omega) = \frac{\Gamma(0)}{\eta_0} \frac{\tau_c}{1 + \tau_c^2 (\omega - \omega_0)^2}$$

Called Lorentzian function

Peak at
$$\omega_0$$

FWHM $\Delta \omega = 2/\tau_c$



Normalized:
$$\frac{1}{2\pi} \int_{-\infty}^{\infty} L(u) \, du = 1$$

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General property of Fourier transforms:

 $\Delta\omega\Delta t\gtrsim\pi$

Here $\Delta \omega$ is spectral bandwidth

= range of frequencies present

So $\Delta \omega$ is related to coherence time τ_c :

$$\tau_c \approx \frac{\pi}{\Delta \omega}$$

Incoherent source has broad bandwidth

Monochromatic source is very coherent

Gives another way to look at loss of visibility

Consider two slit interference

Incoherent source: bandwidth $\Delta \omega$

Imagine we have red and blue light

Red light makes red interference pattern w/ high visibility Blue light makes blue pattern

w/ high visibility

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Observe sum of red and blue:



Pattern washed out at high angles peaks of blue cancel troughs of red Find that interference goes away at $\theta \approx \ell_c/a$ as before Summary

- Characterize random waves with averages
- Describe coherence with $\Gamma(\tau)$ $\Gamma = \langle E(t+\tau)E^*(t) \rangle$
- Describe spectrum with $S(\omega)$ = Fourier transform of $\Gamma(\tau)$

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