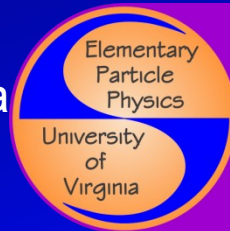


Beyond $E = mc^2$

**Using Rare Particle Decays to
Probe the Energy Frontier**

**Exploiting Heisenberg's Uncertainty
Principle to trump Einstein's $E=mc^2$**

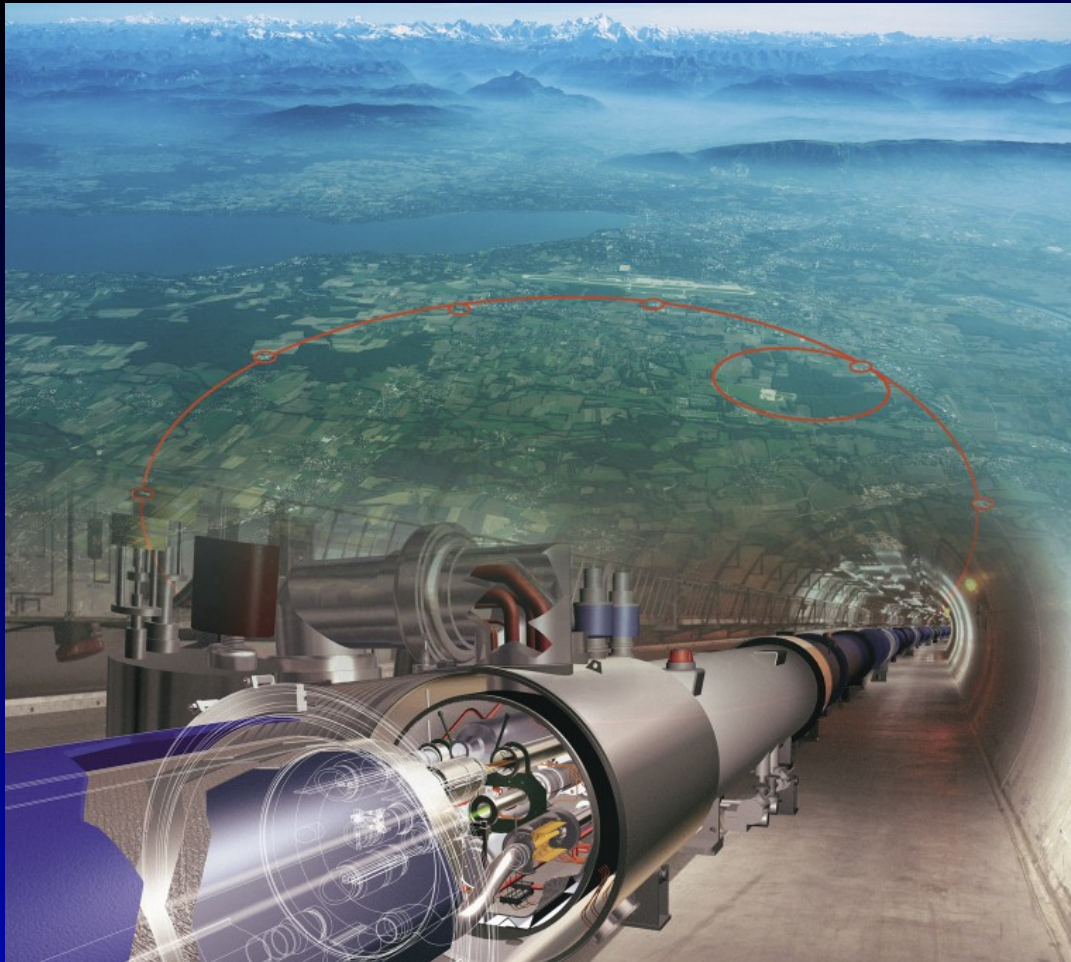
E. Craig Dukes
University of Virginia
BCVSPIN 2011
July 27, 2011



*Antimatter Asymmetry Group
at the University of Virginia*

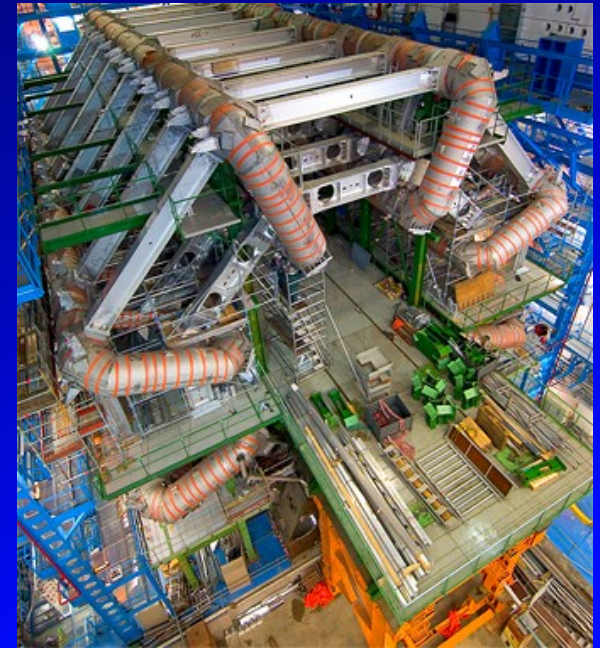


The World Eagerly Awaits Physics from the LHC

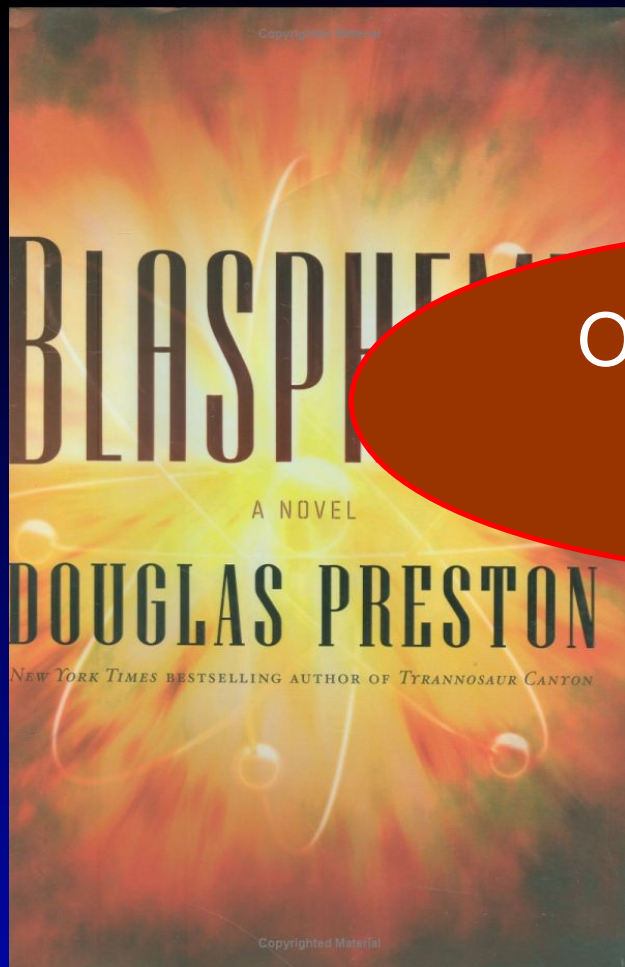


7 trillion (10^{12}) eV proton
energy:

7X Fermilab Tevatron



LHC Has Already Had a Profound Impact



On bad literature!

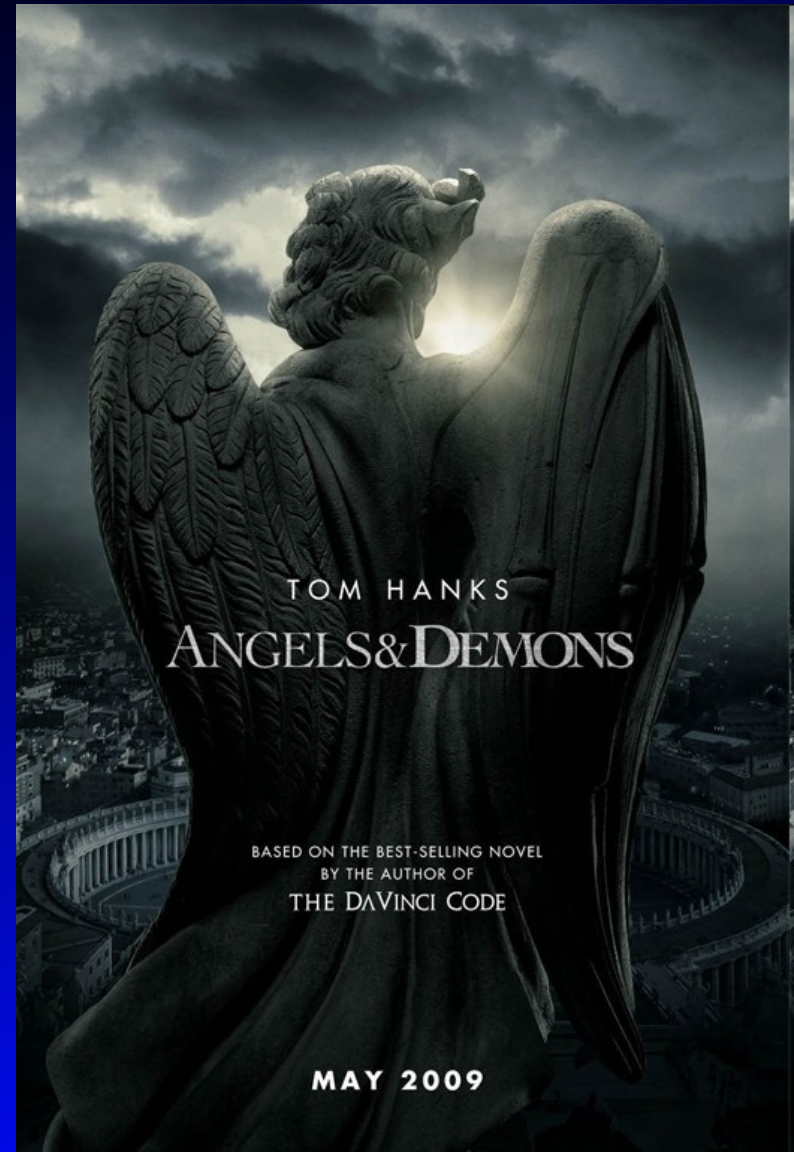
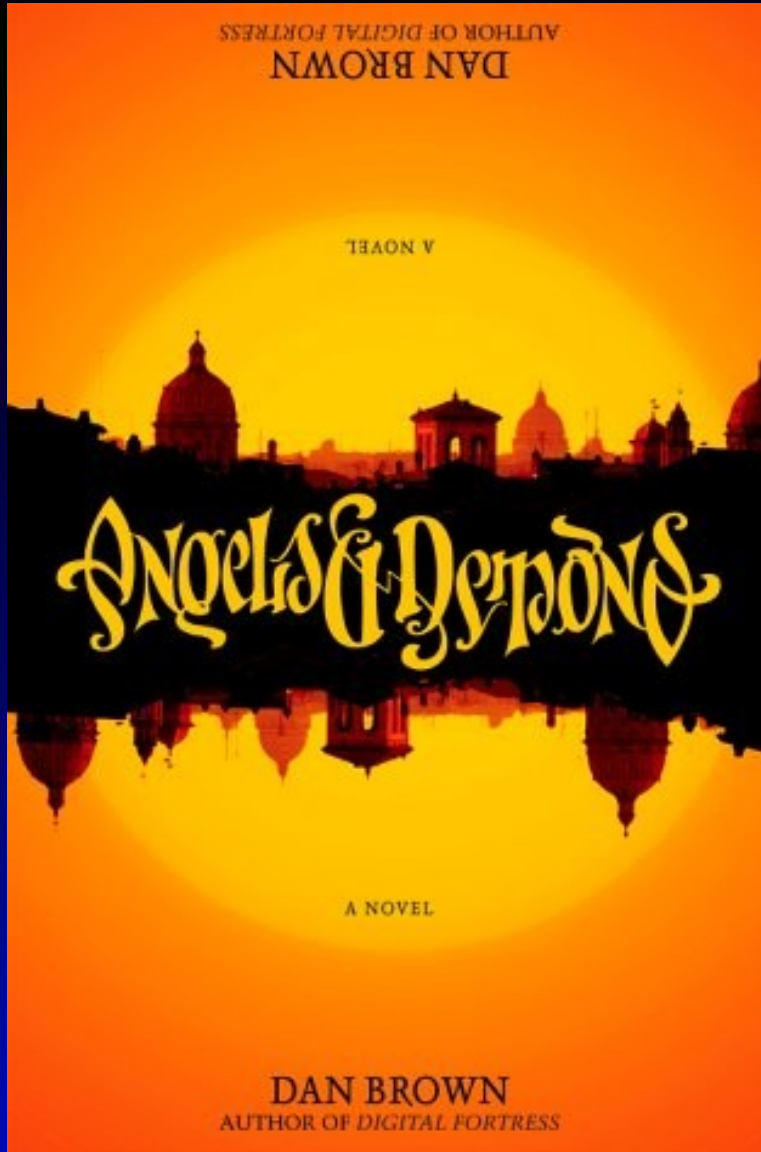
JULY

KEN DOLBY STOOD BEFORE HIS WORKSTATION, his smooth, polished fingers caressing the controls of Isabella. He waited, savoring the moment, and then he unlocked a cage on the panel and pulled down a small red bar.

... something to indicate that the most expensive
... on. Except that, two hundred
... slightly.

... the fine vibration of her
... a woman, and in his more
... what she looked like—tall and
... the desert night, beaded with sweat. Isabella. He had shared these feelings with no one—no point in attracting ridicule. To the rest of the scientists on the project, Isabella was an “it,” a dead machine built for a specific purpose. But Dolby had always felt a deep affection for the machines he created—from when he was ten years old and constructed his first radio from a kit. Fred. That was the radio’s name. And when he thought of Fred, he saw a fat carrot-haired white man. The first computer he had built was Betty—who looked in his head like a brisk and efficient secretary. He couldn’t explain why his machines took on the personalities they did—it just happened.

Angels and Demons



The Cause of our Excitement

JOE H

Understanding of the electromagnetic and weak interactions \hat{c} Higgs?

Origin of neutrino mass hierarchy

Solution to hierarchy problem
supersymmetry?



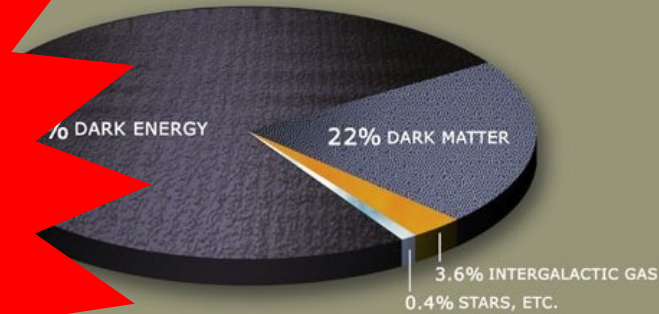
JOHNSON

Matter-antimatter
asymmetry in universe

Dark matter

Dark energy

Some of this
New Physics
may appear
at energies too
high
for the LHC

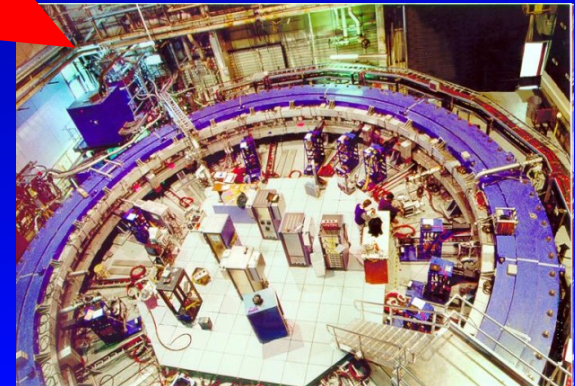


EMIR

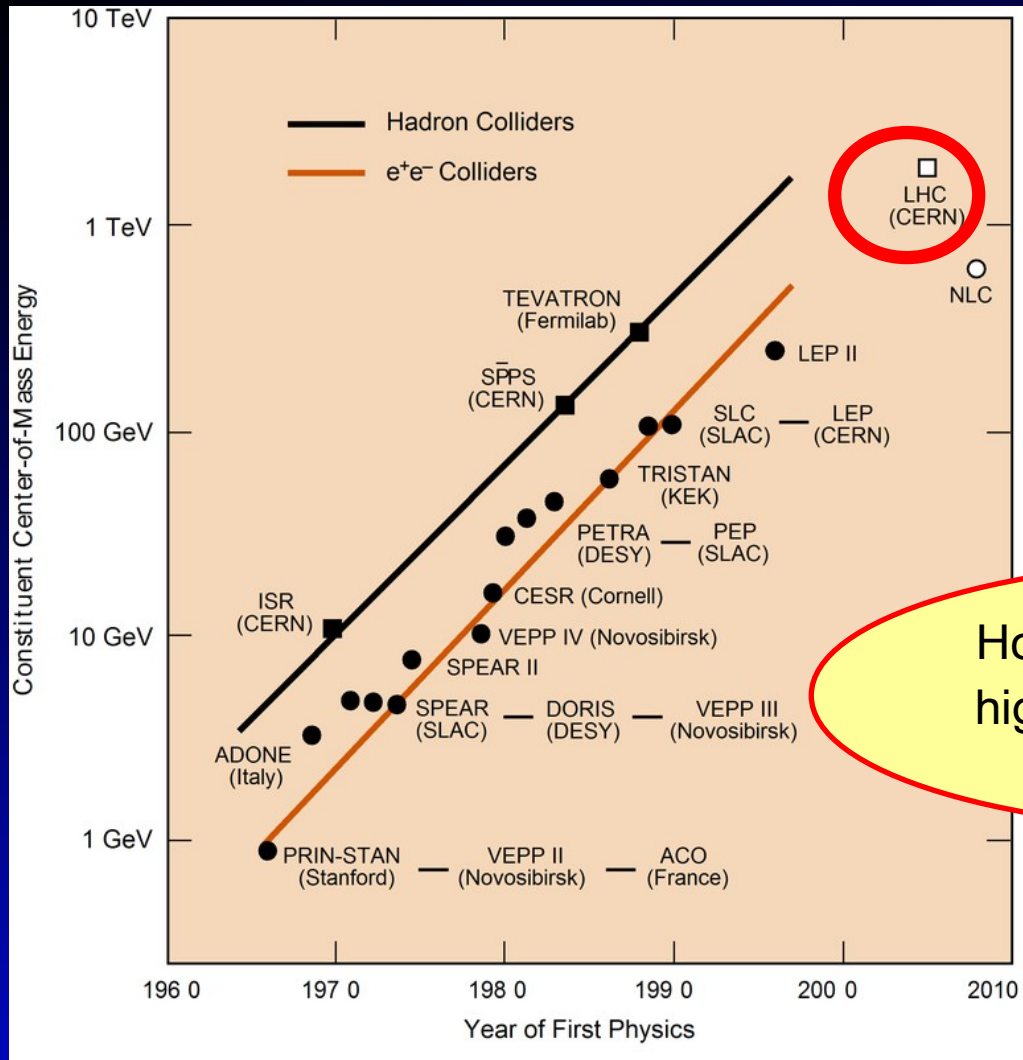
Neutrino mass
physics beyond

Weak interaction studies \hat{c} Higgs?

Other hints, muon $g-2$, NuTeV, CP
phases in Bs mixing, Ds decay rates,



Is the Large Hadron Collider the Last in a Long Line of ever Higher Energy Particle Accelerators?



Presently there are no concrete plans for an accelerator to probe the next energy regime

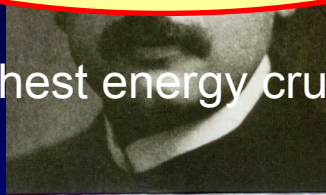
How are we going to probe higher mass scales without higher energies?

Rather than exploiting Einstein's mass-energy relation, $E=mc^2$, exploit Heisenberg's uncertainty principle, $\Delta E \Delta t \gtrsim \hbar/2$

$$E=mc^2$$

appearance of **real** new particles

Highest energy crucial



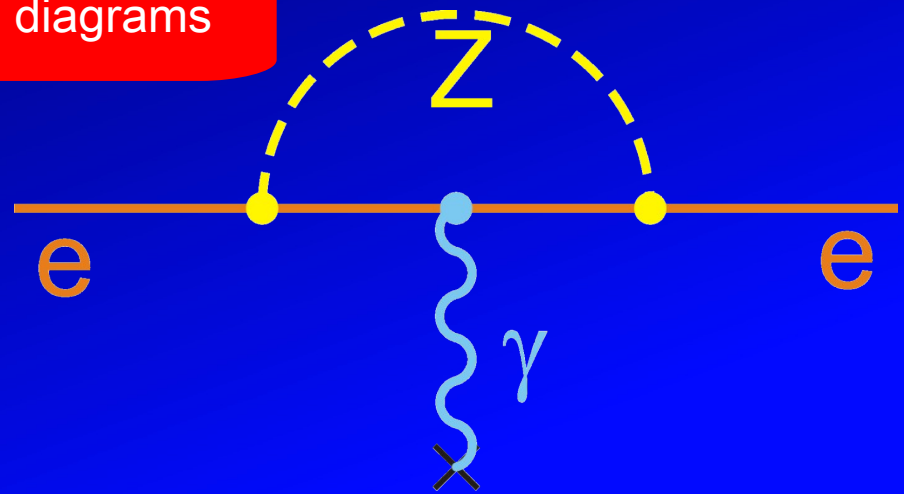
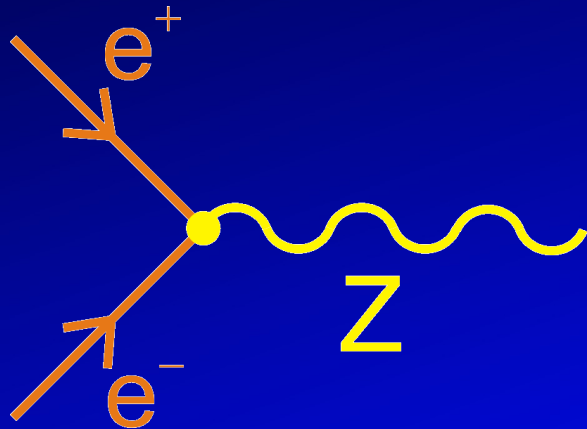
$$\Delta E \Delta t \gtrsim \hbar$$

appearance of **virtual** new particles

Highest intensities crucial



Loop diagrams

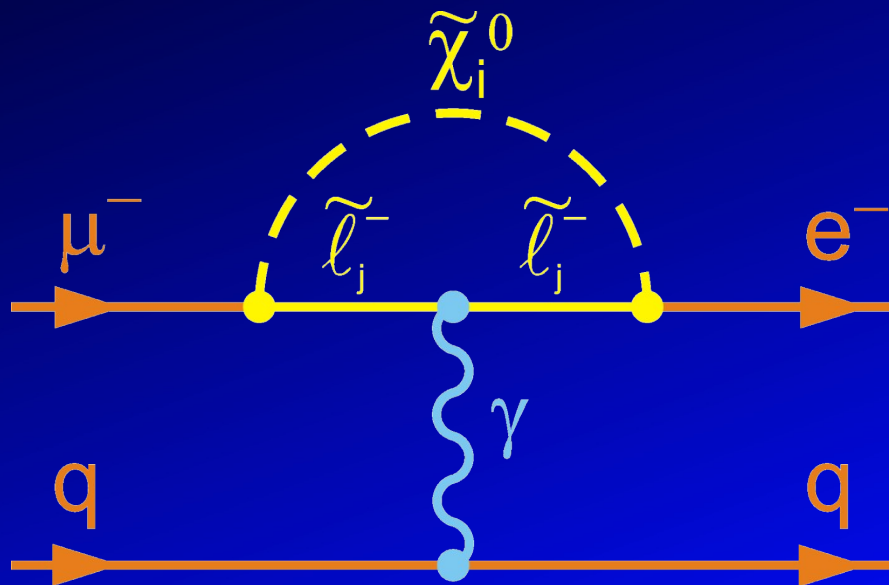


Virtual Particles Can Have Two Effects

They can produce slight deviations from expected properties

Difficult experiments:

- incredible precisions often needed
- theoretical precision needed too
- precision $\propto 1/\sqrt{N}$



\$50,000 reward from Cornell lab of ornithology!

The

es,

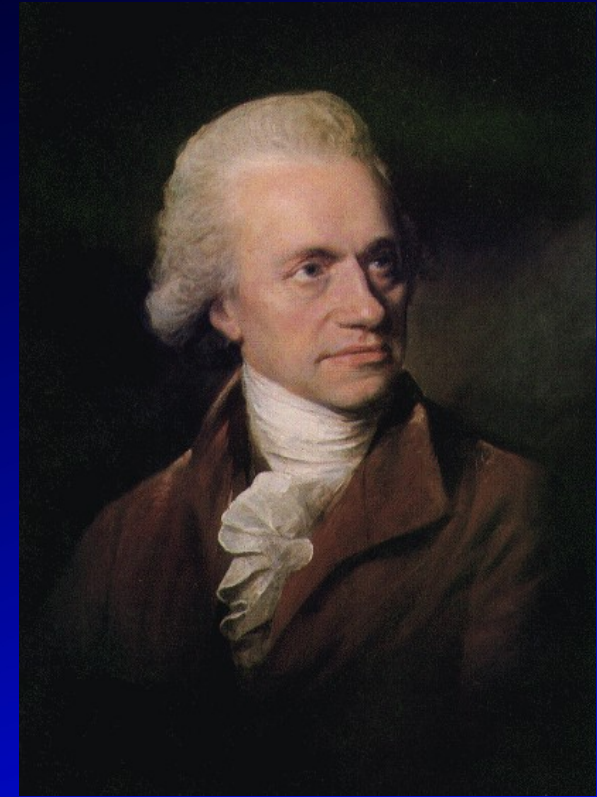
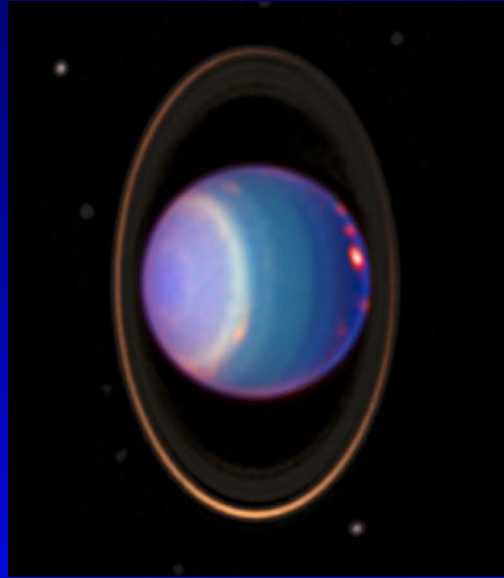
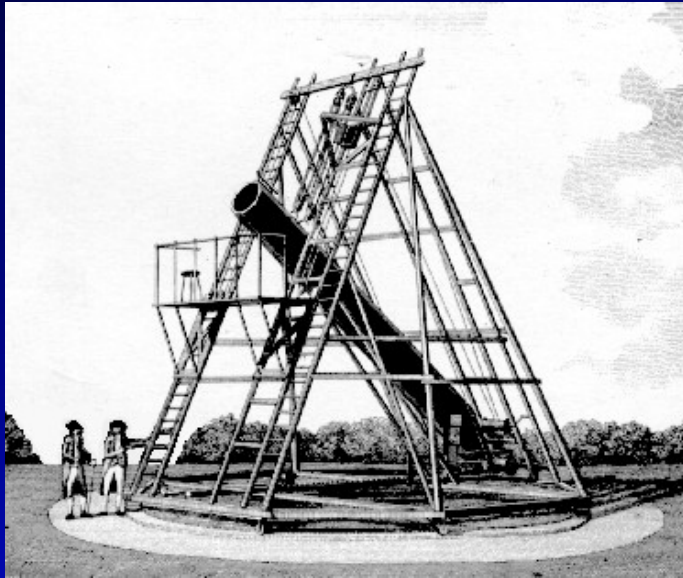
Ivory-billed woodpecker

“Easy” search experiments:

- theory \propto process is “forbidden”
- one event can be enough
- sensitivity $\propto 1/N$

This Indirect Approach has a Distinguished Past

- Starts with William Herschel's discovery Uranus in 1781
- Uranus was the first planet to be discovered – all others were known to the ancients
- Soon a problem appeared ζ orbit of Uranus was found to deviate from predictions from Newton's Laws

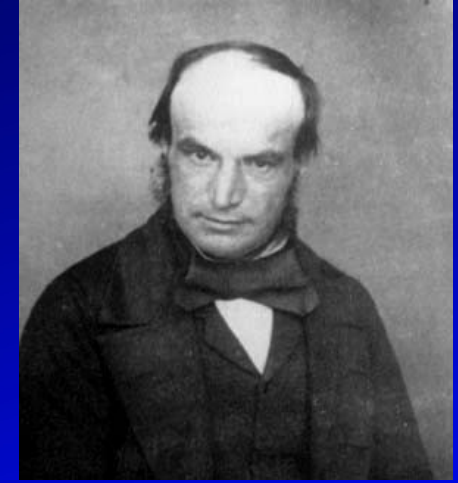
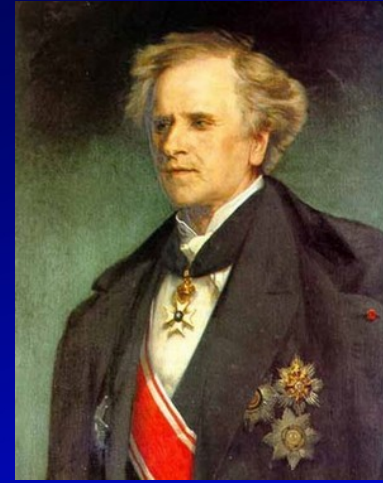


Race to find explanation began

George Airy thought that explanation lied with deviations from Newton's law of gravitation



Le Verrier (France) and John Adams (UK) thought that the deviations were due to a hitherto unknown planet

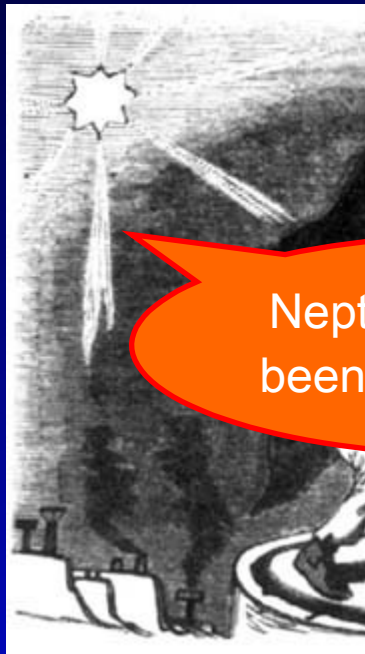


Newton Vindicated

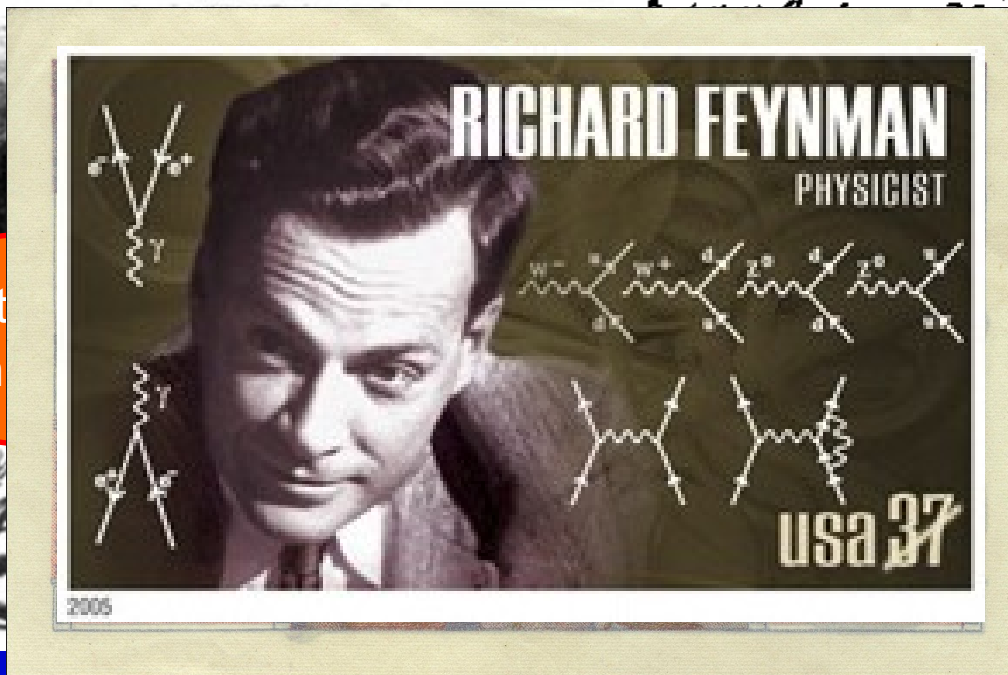
- Le Verrier won: September 18, 1846 he wrote a letter asking Johann Galle at the Berlin Observatory if he might have a look
- On September 23 Galle in Berlin found Le Verrier's planet

Le Verrier had discovered a planet with the tip of his pen, without other instrument than the strength of his calculations alone"

Flammarion

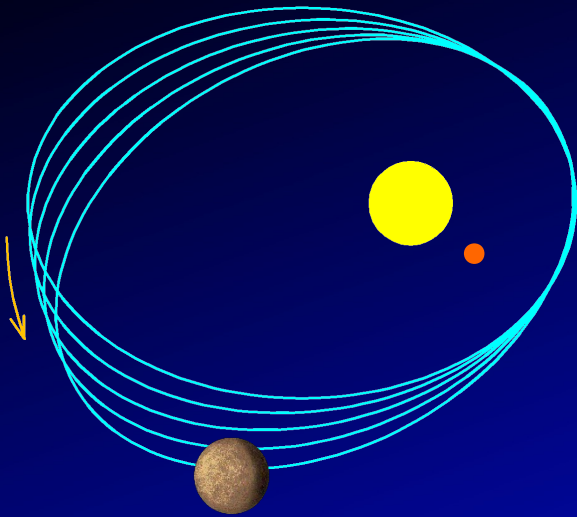


Nept
been



What about Mercury?

Le Verrier turned his formidable talents to the orbits of the planets and found a discrepancy in the precession of the perihelion of Mercury



The race to find planet Vulcan had begun!

In 1859 he found a discrepancy of 1/10,000 of a degree per year!

Effect of the planets on the precession of Mercury's perihelion

Venus	280.6"
Earth	83.6"
Mars	2.6"
Jupiter	152.6"
Saturn	7.2"
Uranus	0.1"
Total per century:	527"
Observed:	527"+38"

ne pouvait être vue que le soir ou le matin dans les vapeurs de l'horizon; en sorte qu'avant l'invention et le perfectionnement des lunettes, il était impossible de l'observer hors de ses elongations. Copernic, empêché par les brouillards de la durée des crépuscules en été, ne put jamais parvenir à

III et IV.

Race to Find Vulcan

- Vulcan was seen many times, but never confirmed
- “Dark matter” near the sun was invoked as a solution
- A modification of Newton’s law of gravity was another: $1/r^{2.00000016}$
- The problem remained outstanding at the beginning of the 20th century

VULCAN AND THE CORONA.

RESULTS OF THE RECENT ECLIPSE.
THE MYSTERIOUS PLANET — THE GASEOUS
AND NON-GASEOUS ELEMENTS OF THE
CORONA—A REVIEW OF THE LABORS OF
THE OBSERVERS BY PROF. YOUNG, OF
PRINCETON.

It is early, as yet, to estimate the full scientific meaning and value of the observations made during the recent solar eclipse, but it is already evident that, though, in some respects disappointing, they are yet, on the whole, of an importance quite equal to those obtained on any similar occasion.

One brilliant discovery will probably date from this occasion, and hold a conspicuous place in the annals of science. The planet Vulcan, after so long eluding the hunters, showing them from time to time only uncertain tracks and signs, appears at last to have been fairly run down and captured. At least it seems to us that the observations of Prof. Watson at Rawlings, and Swift at Denver, must for the present be taken as conclusive, though perhaps not settling the question beyond the possibility of reopening or dispute. The gentlemen are both astronomers of repute, accustomed to sweep for faint objects, and provided with excellent instruments. The negative results of Profs. Newcomb, Wheeler, Holden and others, who, with similar instruments, went over the same ground and found nothing, are, indeed, unsatisfactory and puzzling; but they can hardly outweigh the positive evidence on the other side, though they certainly justify a certain reserve in accepting the conclusion.

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Vulcan not Found but General Relativity is

Einstein calculates the perihelion of Mercury using his new theory of General Relativity and recovers the missing 38"/century found by le Verrier 56 years earlier!

Erklärung der Perihelbewegung des Merkur aus der allgemeinen Relativitätstheorie.

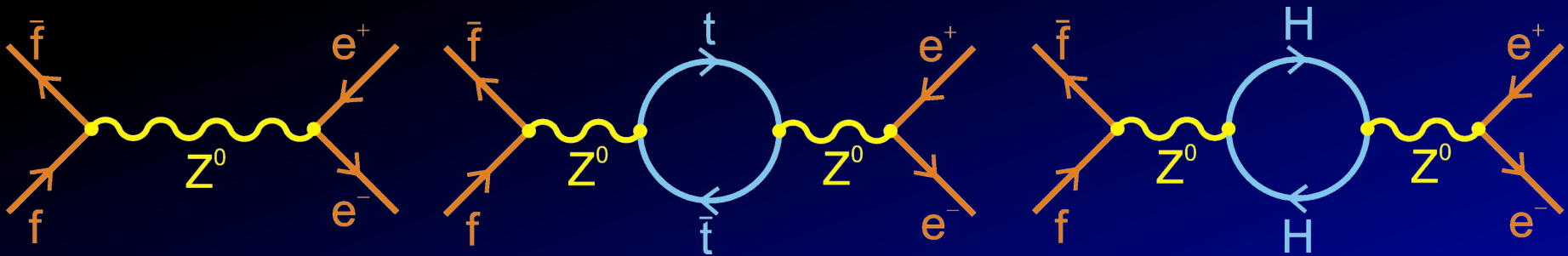
VON A. EINSTEIN.

Königlich Preussische Akademie der Wissenschaften (Berlin).
Sitzungsberichte (1915): 831-839.



Three weeks later, a critic wrote in the *Prussian Academy of Sciences* journal, "Einstein's theory was a better toy for children than a science." The critic had been critical of the "pedantic account of Einstein's theory" in the *Prussian Academy of Sciences* journal, *Sitzungsberichte*, November 1915.

Example: Top and Higgs Masses

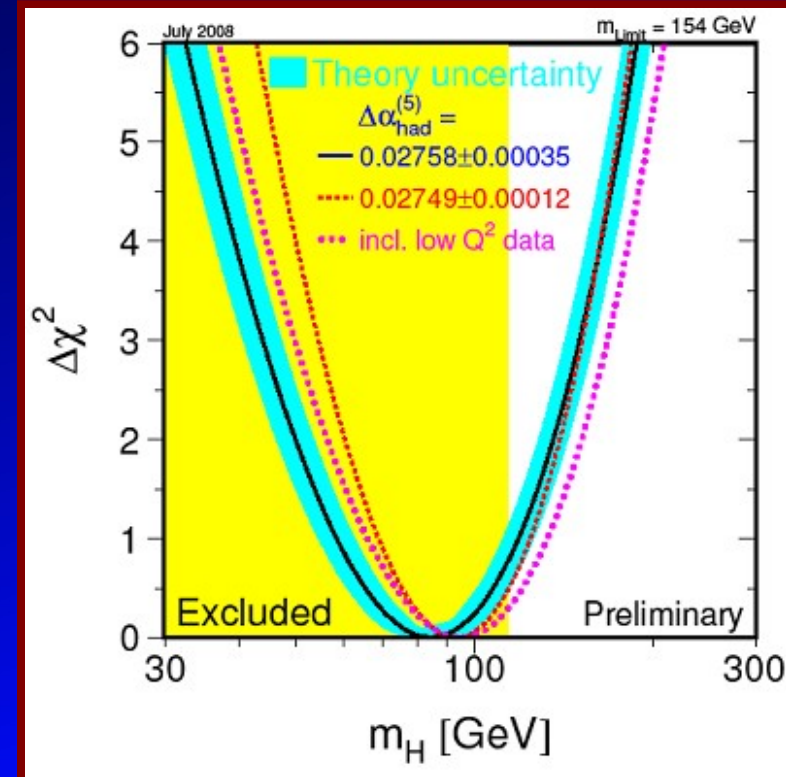
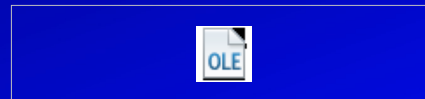


Virtual particles relate the properties of the weak force carriers (W, Z), and the masses of the top quark and Higgs

Top quark mass “predicted” by precision electroweak measurements before it

was directly discovered by CDF and DØ at Fermilab: $m_t = 162 \pm 9$ GeV (Ellis, Fogli, Lisi, 1994).
Now that top quark mass is known the same game is being played to predict the Higgs mass

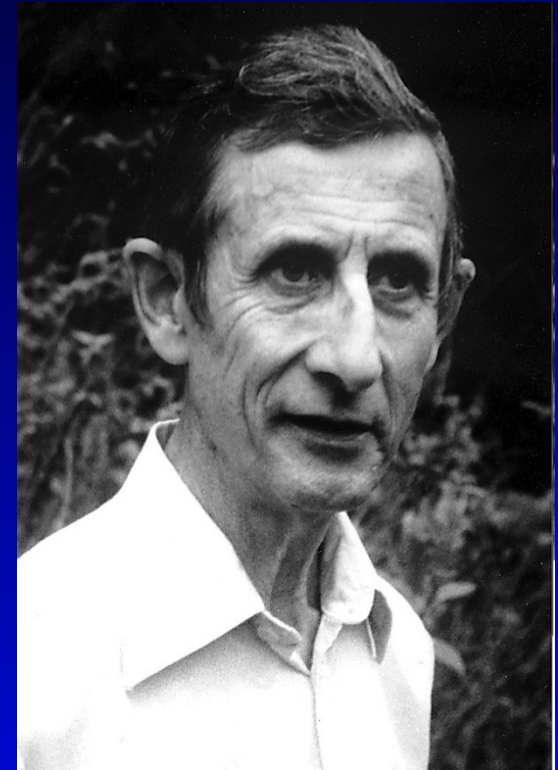
most probable value:



Rare and Precision Physics has a Distinguished Past

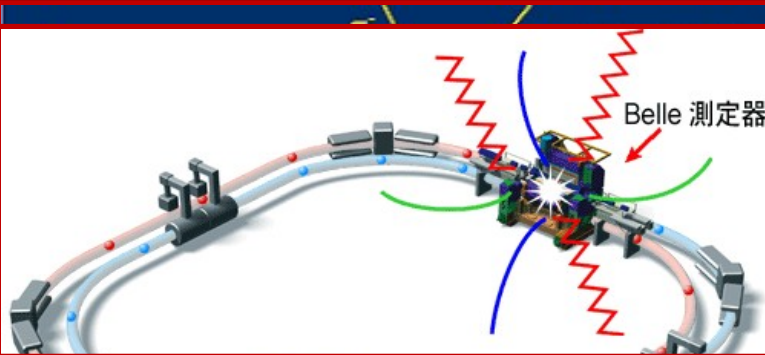
“The results of my survey are then as follows: four discoveries on the energy frontier, four on the rarity frontier, eight on the accuracy frontier. Only a quarter of the discoveries were made on the energy frontier, while half of them were made on the accuracy frontier. For making important discoveries, high accuracy was more useful than high energy.”

Freeman Dyson, review of *The Lightness of Being*, F. Wilczek



Rare Decays Require Flavor Factories

High
school
CERN
stud



Flavor
production
(flavor)



BES-III tau-charm
Factory at Beijing, China

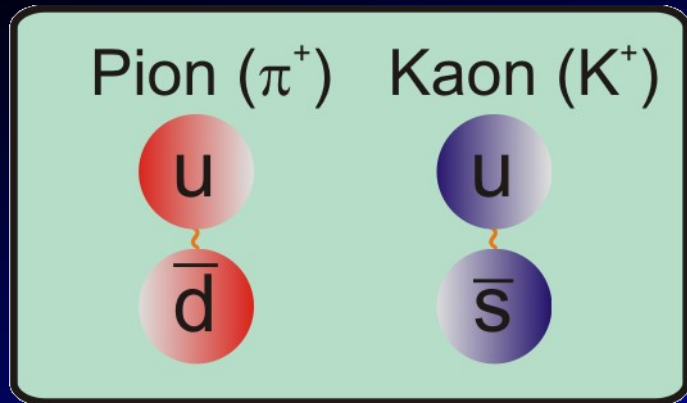
Forces	Strong	g	Gluons
	Electromagnetic	γ	Photon
	Weak	Z^0, W^\pm	Weak bosons
	Gravitational	G	Graviton

Constituents				Charge	} γ, Z^0, W^\pm G	
	1	2	3			
	Quarks	u	c	t		$2/3$
		d	s	b		$-1/3$
Leptons	ν_e	ν_μ	ν_τ	0		
	e	μ	τ	-1		

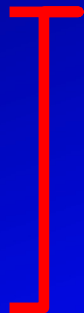
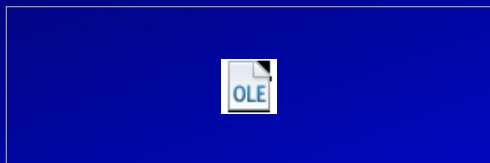
10 billion b/anti-b pairs/yr
10 billion τ /anti- τ pairs/yr

Proton Accelerators are General Purpose Flavor Factories

Produce beams of pions and kaons



Which in turn produce beams of muons and muons neutrinos



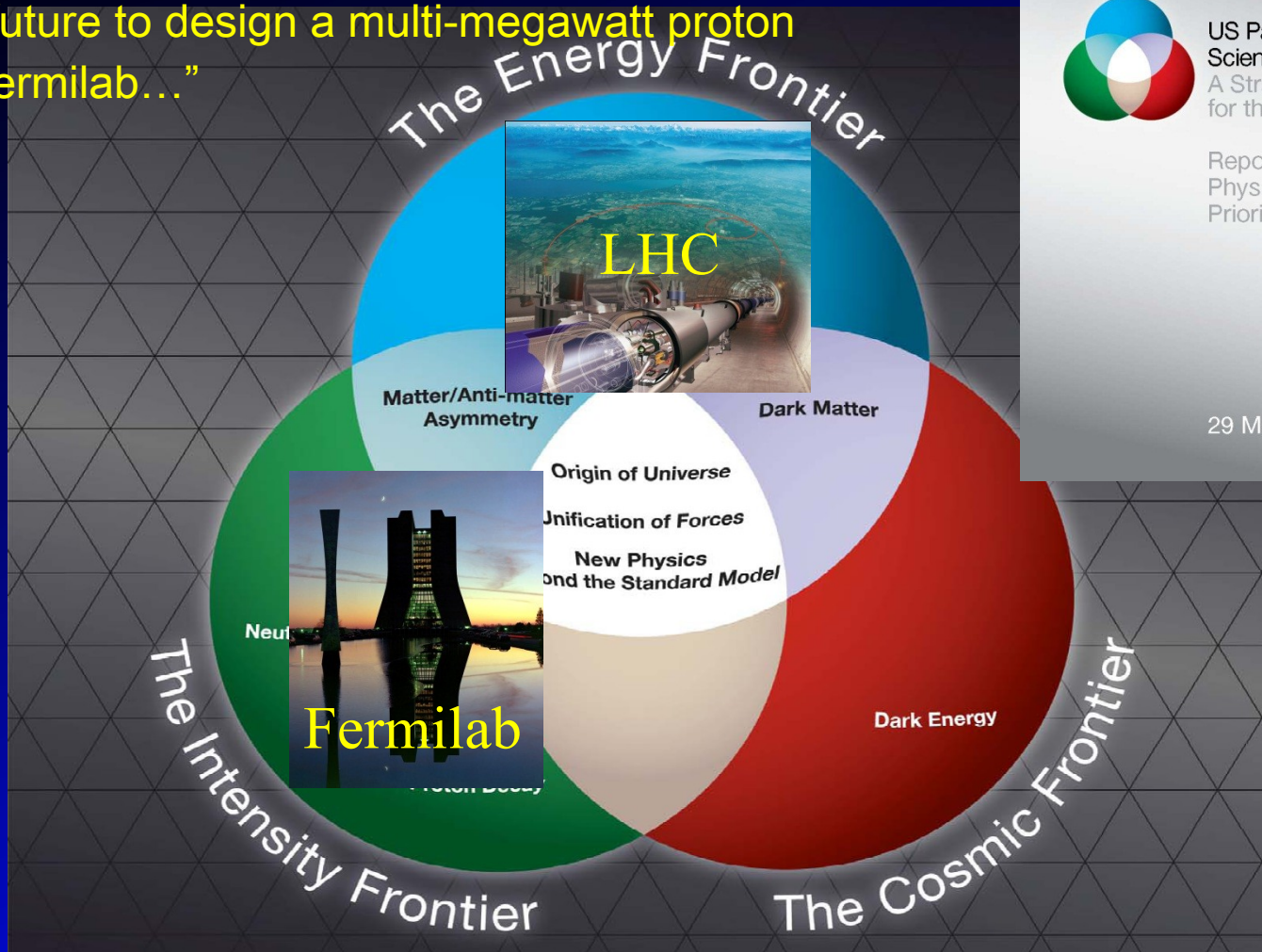
~1-100 billion/s can be produced

Forces	Strong	g	Gluons			
	Electromagnetic	γ	Photon			
	Weak	Z^0, W^\pm	Weak bosons			
	Gravitational	G	Graviton			
Constituents		1	2	3	Charge	
	Quarks	u	c	t	$2/3$	} g γ Z^0, W^\pm
		d	s	b	$-1/3$	
	Leptons	ν_e	ν_μ	ν_τ	0	} γ Z^0, W^\pm
		e	μ	τ	-1	

Fermilab Pushing Forward on Intensity Frontier

Strategic Plan for the Next Ten Years:

“The panel recommends an R&D program in the immediate future to design a multi-megawatt proton source at Fermilab...”



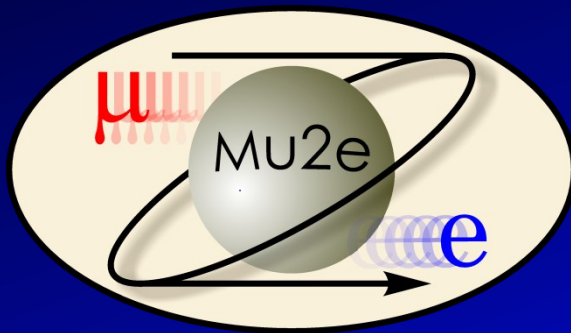
US Particle Physics:
Scientific Opportunities
A Strategic Plan
for the Next Ten Years

Report of the Particle
Physics Project
Prioritization Panel

29 May 2008

A Snapshot of Two Intensity Frontier Experiments

Rare muon decay experiment



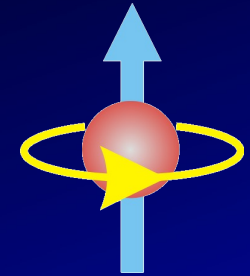
Precision muon experiment



Magnetic Moment of the Muon: Theory

Measuring magnetic moments of fundamental particles has long and productive history

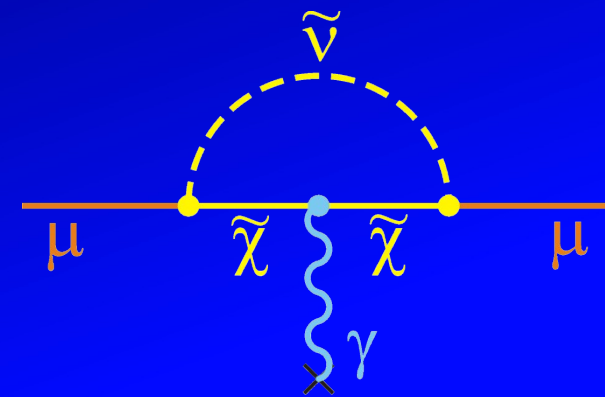
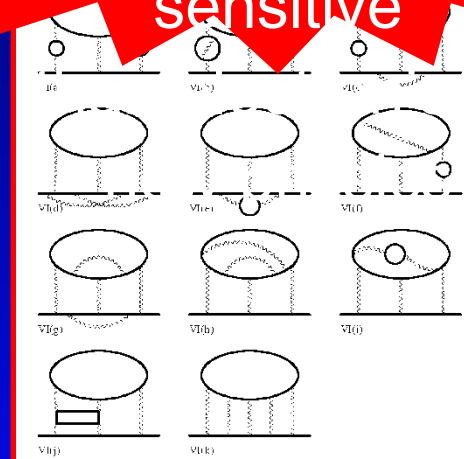
- ⇒ anomalous magnetic moments of protons and neutrons implied substructure \checkmark quarks
- ⇒ Most precise test of quantum electrodynamics (QED)



Muon's magnetic moment
40,000 times more sensitive

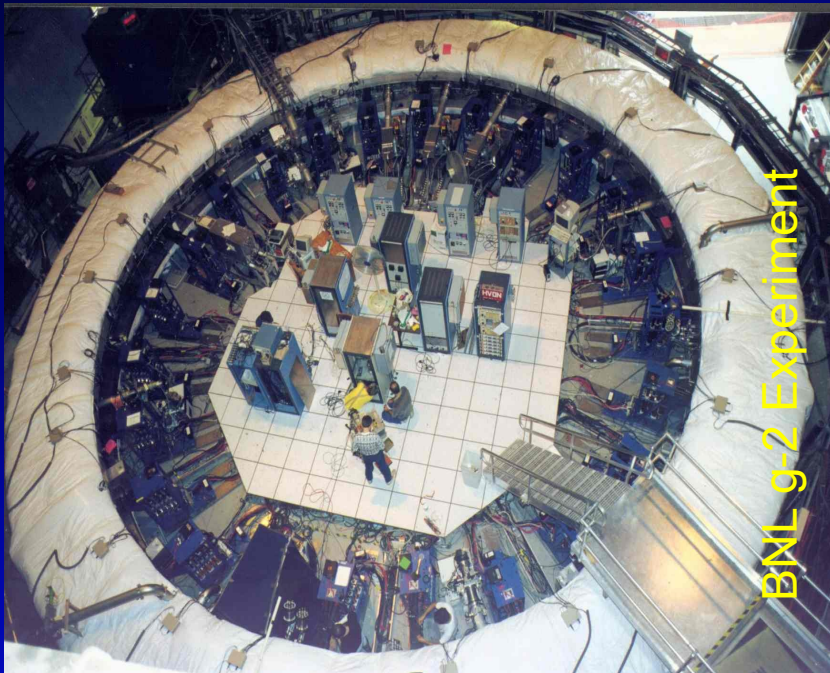
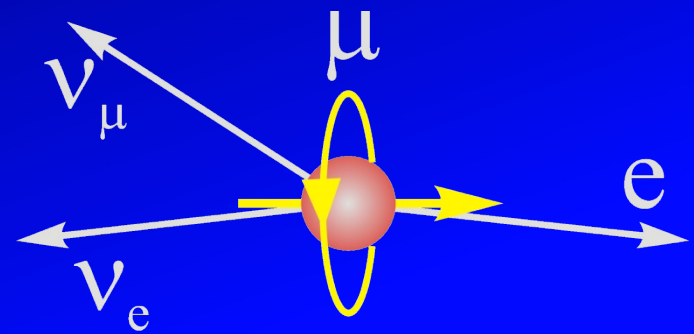
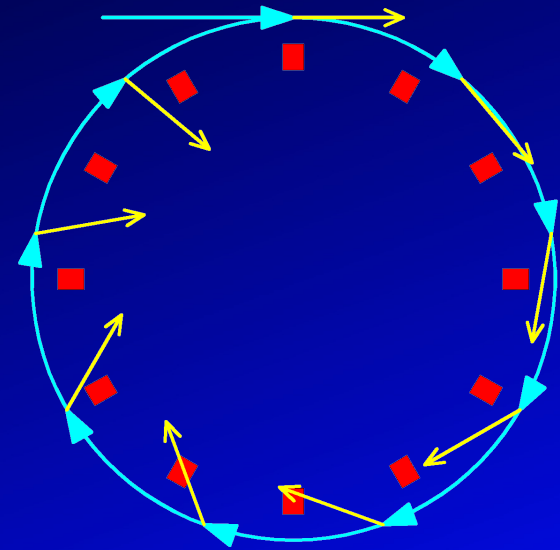
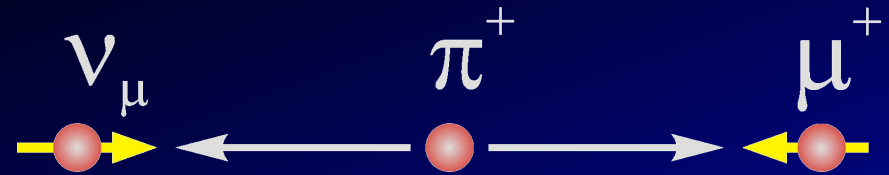
Dirac

New Physics



Magnetic Moment of the Muon: Experiment

1. Pion decay produces polarized muons
2. Precession in magnetic field proportional to anomalous magnetic moment
3. Parity-violating decay of muons reveals precessed magnetic moment direction




Theory Meets Experiment

Theory and experiment differ by $> 3\sigma$

$$\Delta = (296 \pm 81) \times 10^{-11}$$

New Physics?

WINNERS & LOSERS



ARIEL SHARON
Bags biggest landslide in Israeli history. Transformed from a hawk to a phoenix


ANTHONY HOPKINS
Hannibal star proves what many already know: beastly behavior gets results in Hollywood

THE MUON
Throws 30-year-old theory of the universe into doubt. Not bad for a subatomic particle

JOE CAROLLO
Police claim Miami mayor hit wife with a tea canister. Good thing he isn't seeking re-election

HELMUT KOHL
Former German Chancellor fined for taking illegal campaign donations. What, no furniture?

60 MINUTES
Ratings slipping after 23 straight years in Top 10. Lucky CBS has another *Survivor*



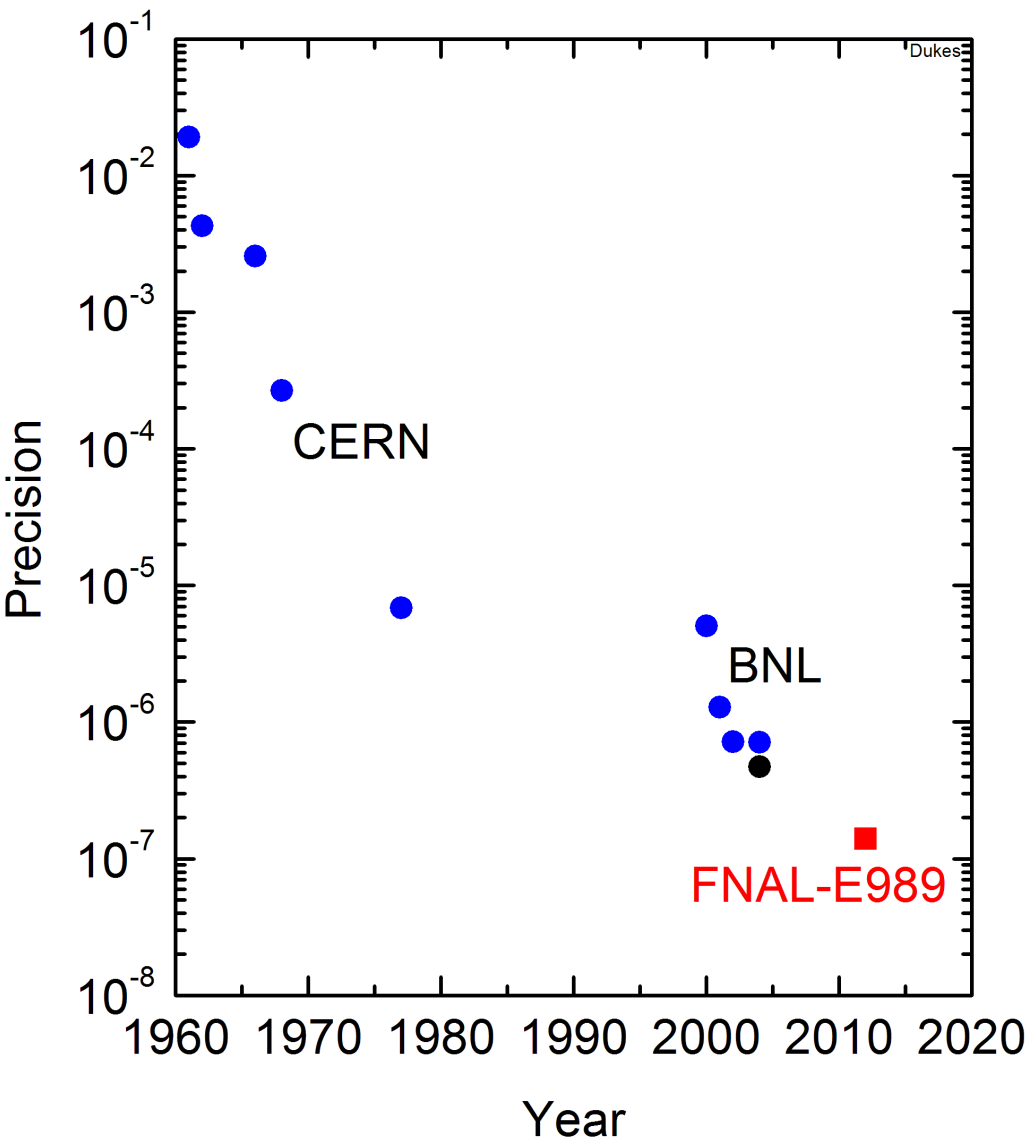
TIME, FEBRUARY 19, 2001

17

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116 591
The
116 592
 μ
116 593
116 594
116 595
 $\times 10^{-11}$

Pushing the Precision at Fermilab: E989

Approved
February 2011



Error Budget (ppm)			
	BNL		Fermilab
Statistical	0.46	→	0.10
Systematic	0.18	→	0.07
Total:	0.54	→	0.14

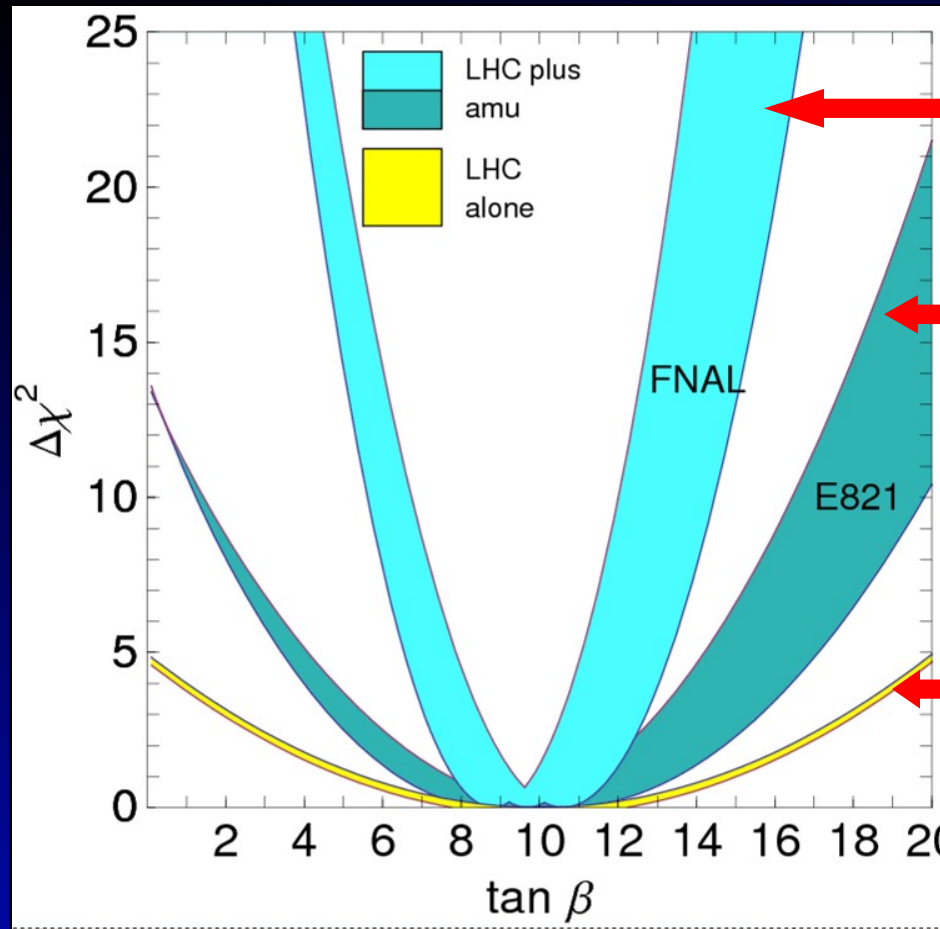
20% more beam
Better

- calorimeter/DAQ
- Longer beamline

Moving the BNL Magnet to Fermilab



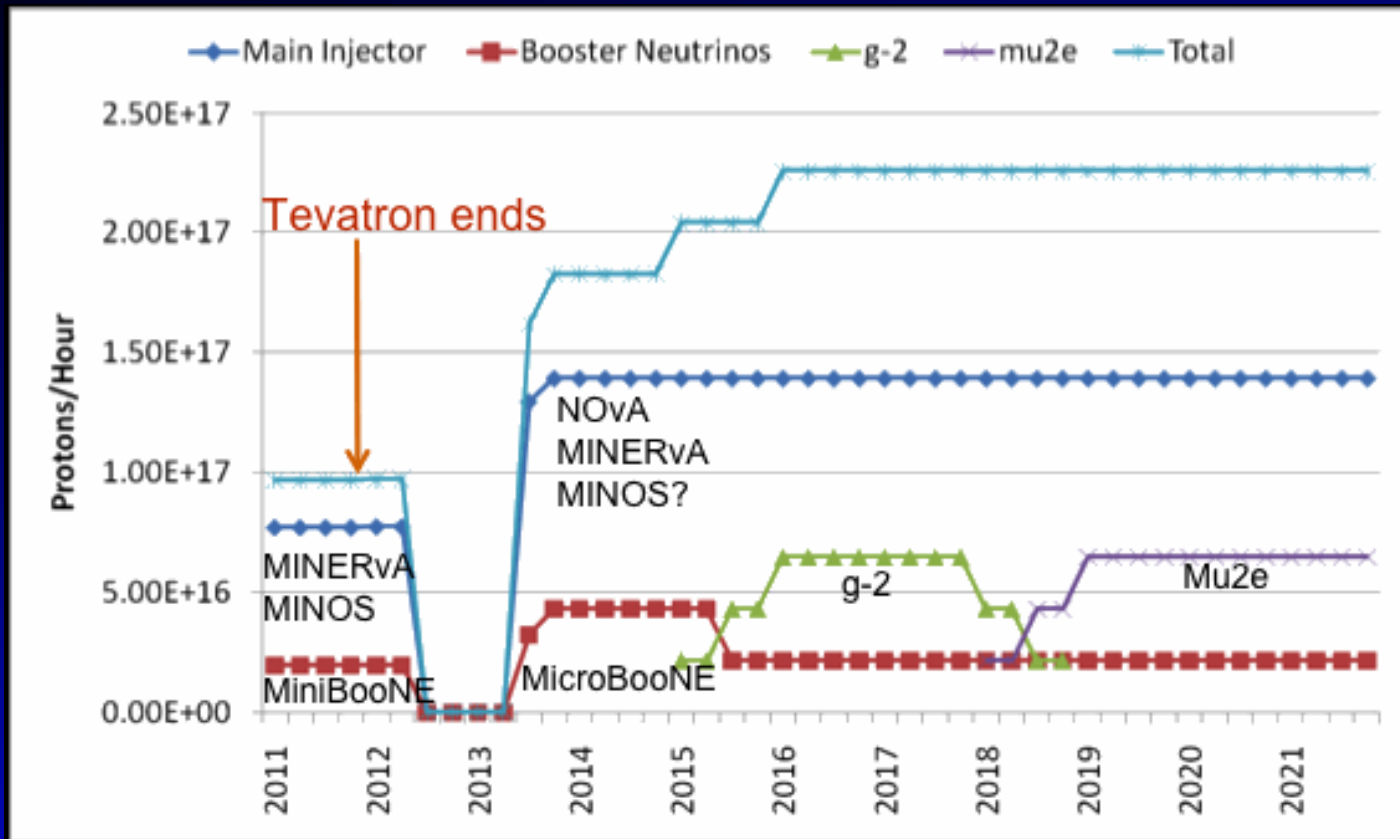
SUSY Meets Muon $g-2$



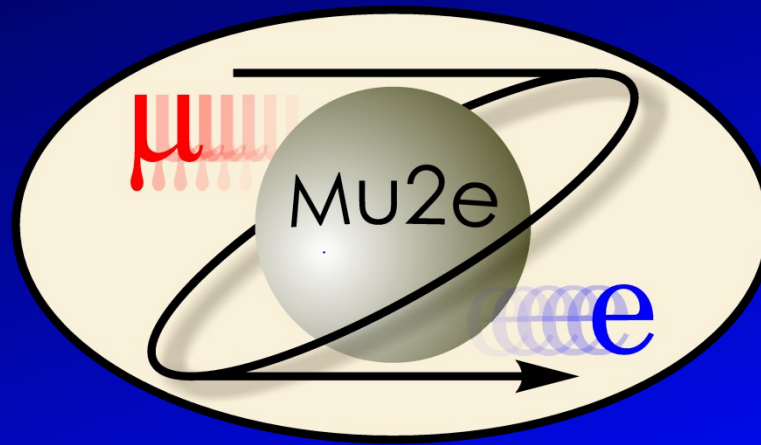
“ $g-2$ is the most important constraint (for SUSY), even more important than dark matter”

Proposed S schedule

- E989 approved in Feb. 2011
- Data taking in 2016



Search for Charged Lepton Flavor Violation

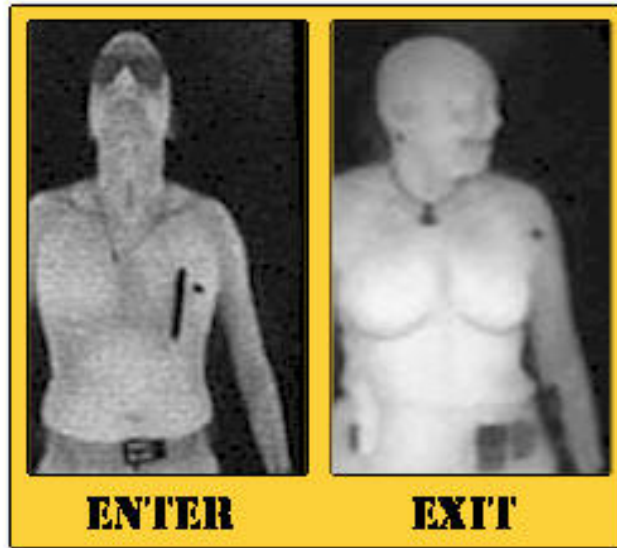


Quark and Lepton Alchemy

1 September 2005

Passenger sues TSA over "instant" sex change Agency blames malfunctioning X-ray scanner

DALLAS, Texas -- An American Airlines passenger who was passing through a newly installed X-ray machine here got more than scanned this past month. In a freak accident during a thunderstorm at the Dallas/Fort Worth International Airport, a lightning strike caused a surge in power just as the passenger entered the device. According to Dr. Sanjay McMurphy, a physician who happened to be at the scene during the incident, "the momentary increase in voltage apparently caused the instantaneous mutation of the passenger's hormonal structure, instantly changing him from a man into a woman."



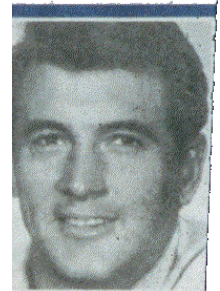
In response to the incident, Rhoda Prindable, a spokesperson for the Transportation Security Administration (TSA) said, "While we have had reports of people aging significantly while passing through the body scanners, this is the first we've heard of the machines contributing to a sex change. The TSA has opened an investigation, the results of which should be known before the devices are widely deployed."

Quarks

Leptons

1
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W

e^-



NS

AN

ES

vert to one
ton flavor

Why Search for Charged Lepton Flavor Violation?

In Standard Model not there $\mu \rightarrow e \gamma$ neutrino mass discovery implies an unobservable 10^{-52} rate

Hence, any signal unambiguous evidence of new physics

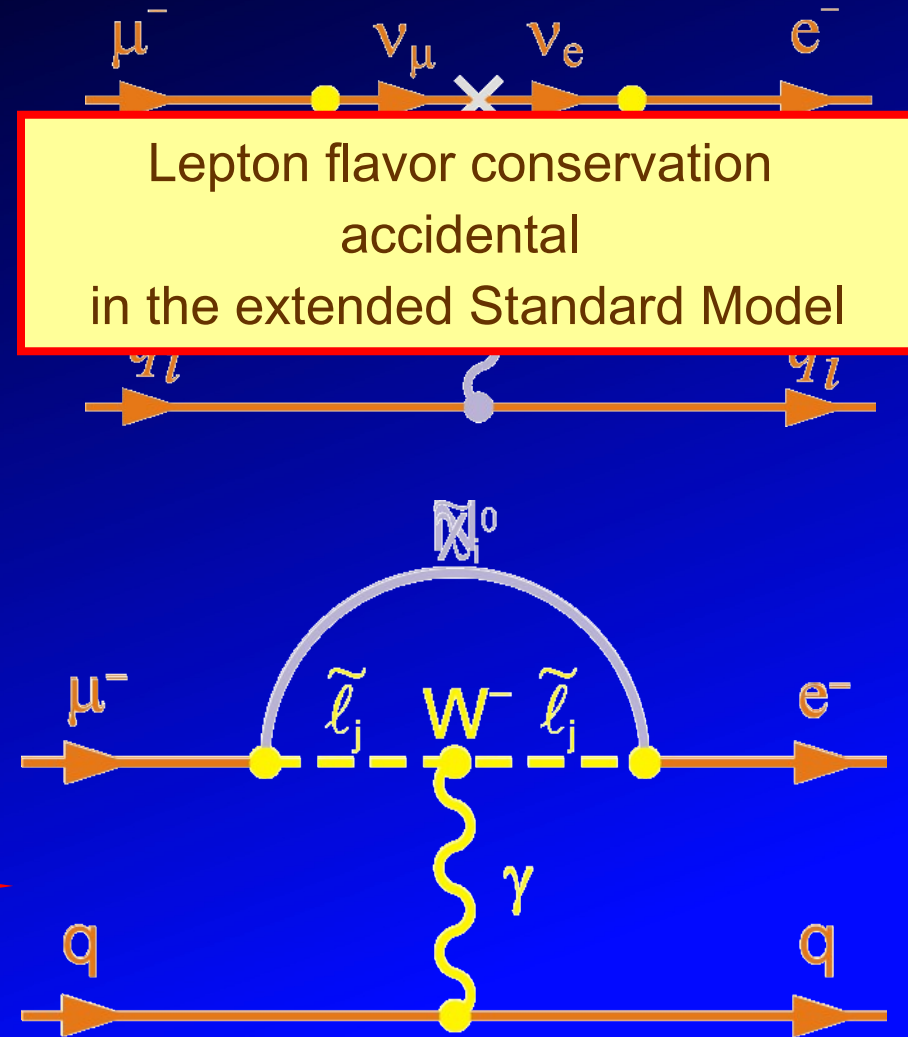
Exquisite sensitivities can be obtained experimentally

⇒ sensitivities that allow favored beyond-the-standard-model

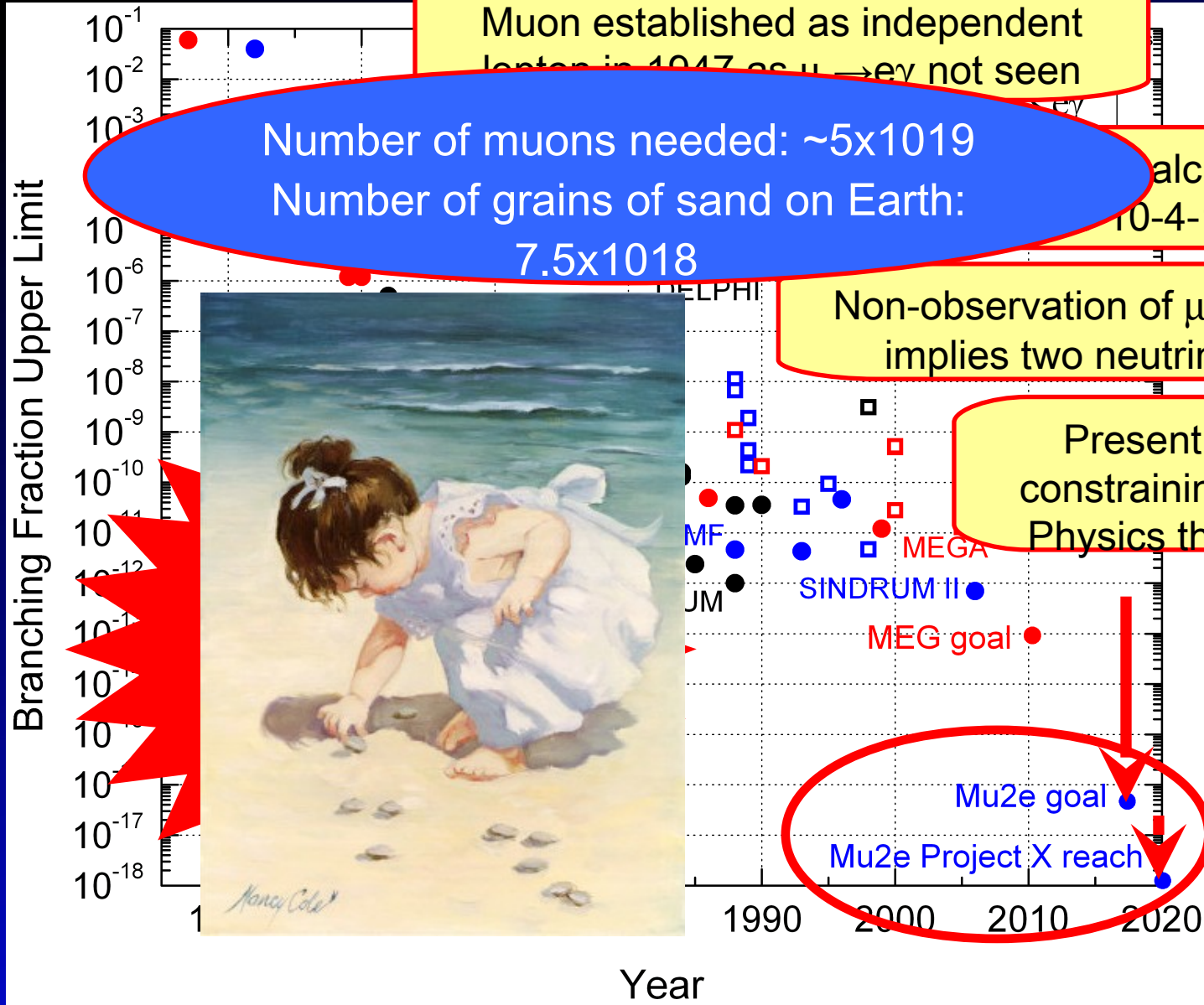
Almost all models explaining the

Supersymmetry

$\mu \rightarrow e \gamma$ at levels that will be probed by Mu2e



Incomplete History of Lepton Flavor Violation Searches



Muon established as independent lepton in 1947 as $\mu \rightarrow e\gamma$ not seen

Number of muons needed: $\sim 5 \times 10^{19}$
 Number of grains of sand on Earth: 7.5×10^{18}

Calculation: $10^{-4} - 10^{-5}$

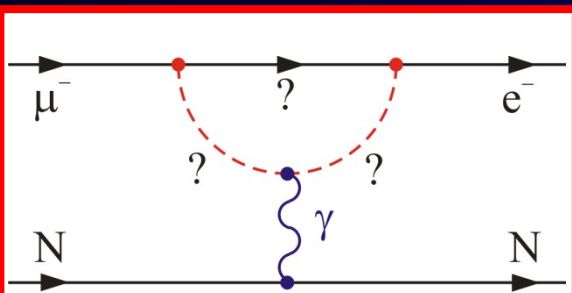
Non-observation of $\mu \rightarrow e\gamma$ implies two neutrinos

Present limit constraining New Physics theories



CLFV in $\mu \rightarrow e + \gamma$ and $\mu - N \rightarrow e - N$

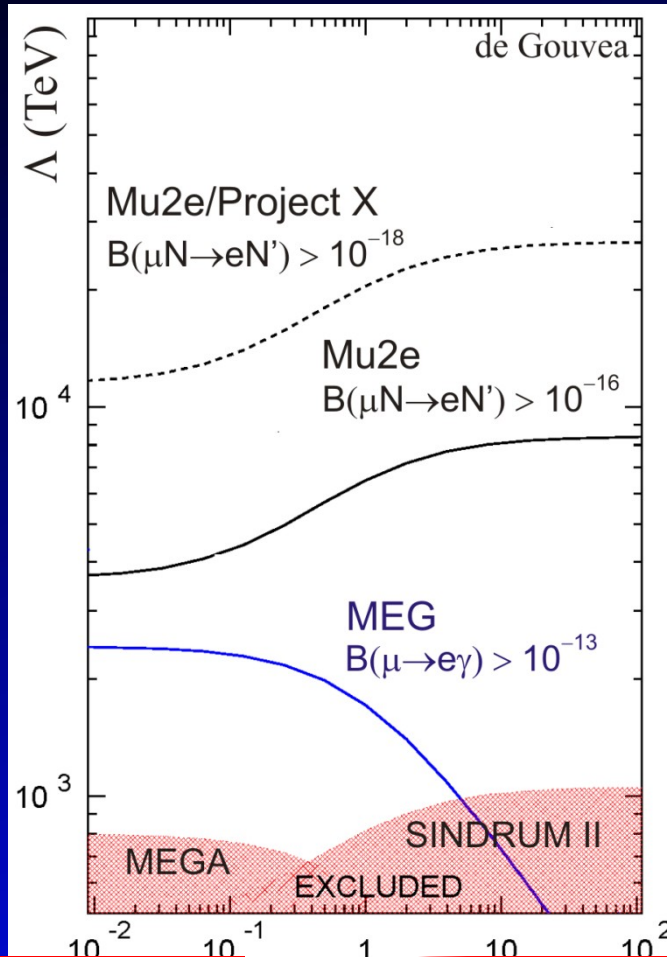
Model independent effective CLFV Lagrangian



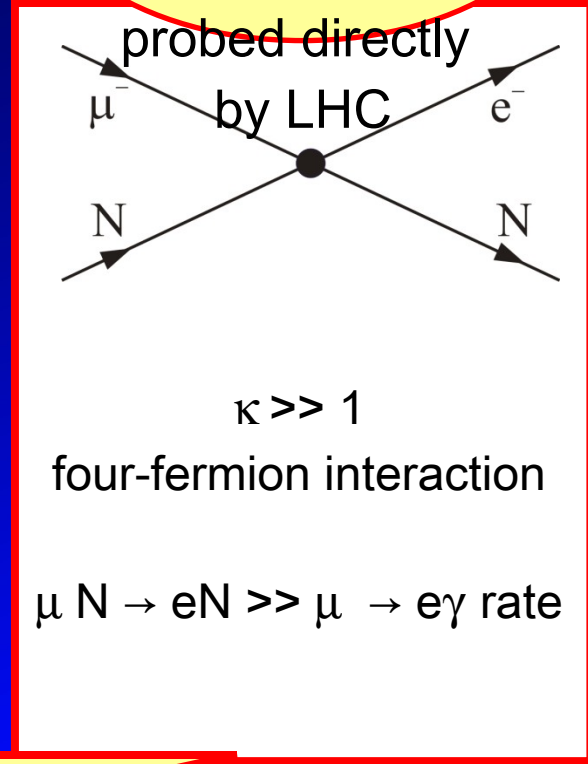
$$\kappa \ll 1$$

magnetic moment type operator

$\mu \rightarrow e\gamma$ rate $\sim 300X$
 $\mu N \rightarrow eN$ rate

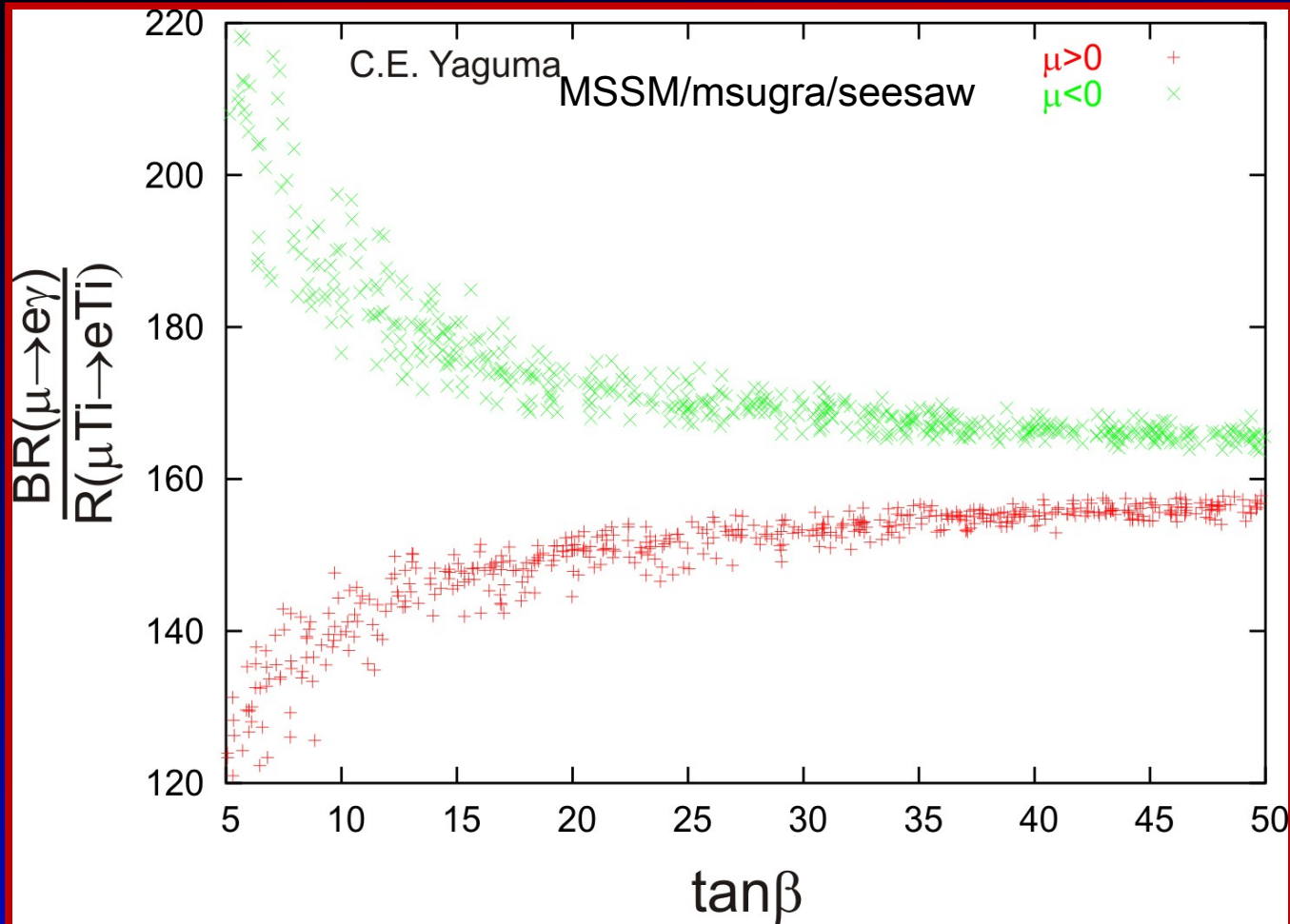


Mass scales probed $\sim 10,000$ times that



Two Methods are Complementary

Observation of CLFV in both $\mu^- \rightarrow e^- N$ and $\mu^+ \rightarrow e^+ \gamma$ could elucidate SUSY parameters

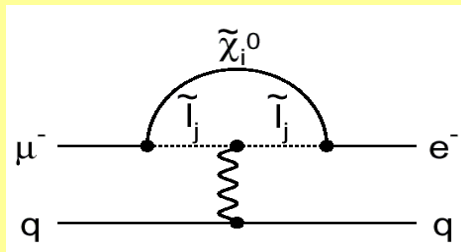


μ = mass term
mixing 2 Higgs
doublets

$\tan\beta$ = ratio of 2
vacuum values of
2 neutral Higgs

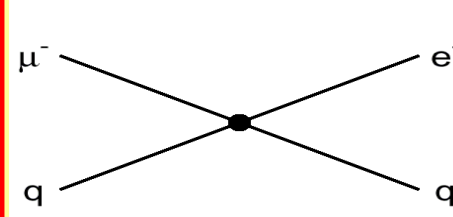
CLFV Sensitive to Many Sources of New Physics

Supersymmetry

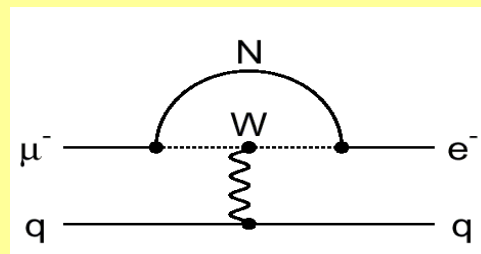


Predictions at 10-13

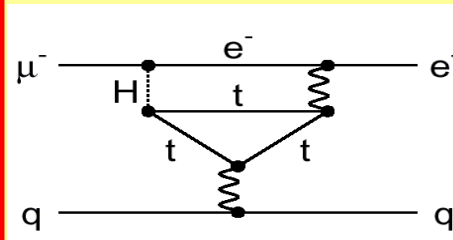
Compositeness



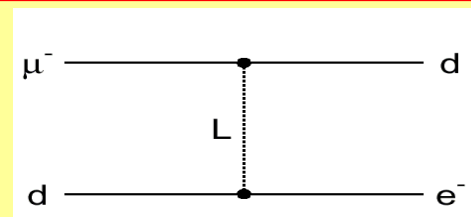
Heavy Neutrinos



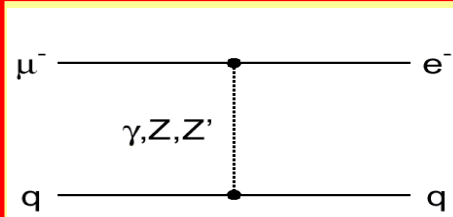
Second Higgs doublet



Leptoquarks



Heavy Z'
Anomalous Z
coupling



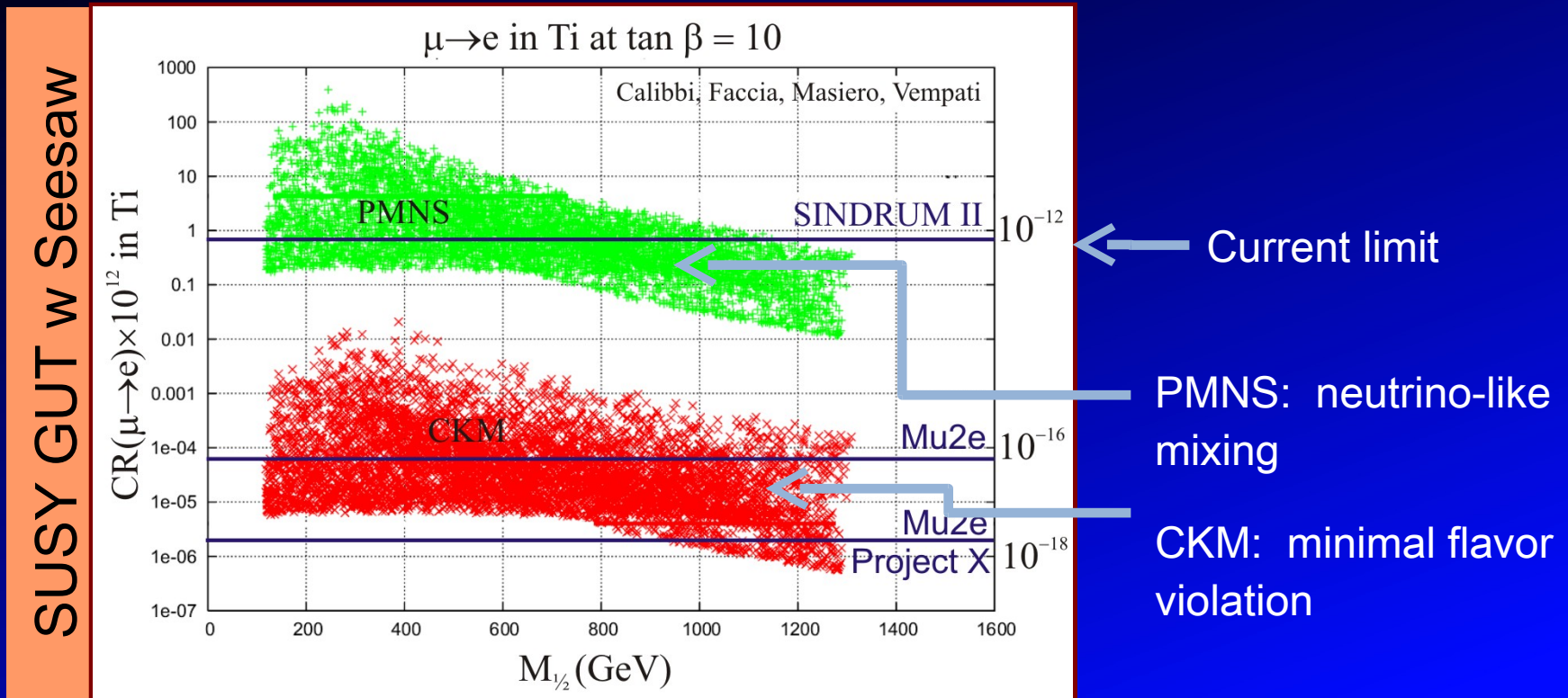
New physics probed by $\mu N \rightarrow e N$ (Marciano)

What Sensitivity is Needed?

Present sensitivity already interesting and constraining!

~ 10^{-16} removes many models

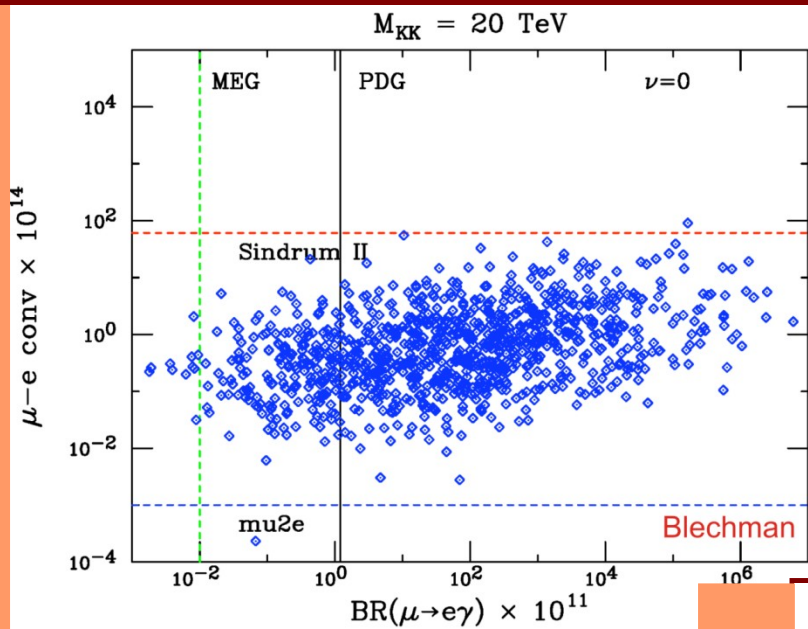
~ 10^{-18} extremely difficult for theorists to deal with



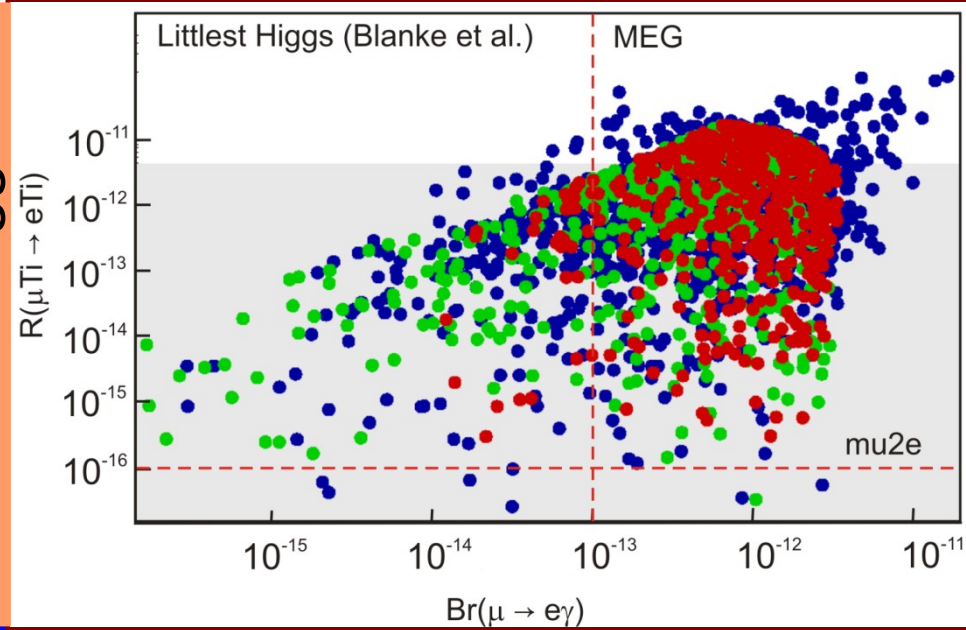
$M_{1/2}$ = gaugino mass

What Sensitivity is Needed?

Randall-Sundrum

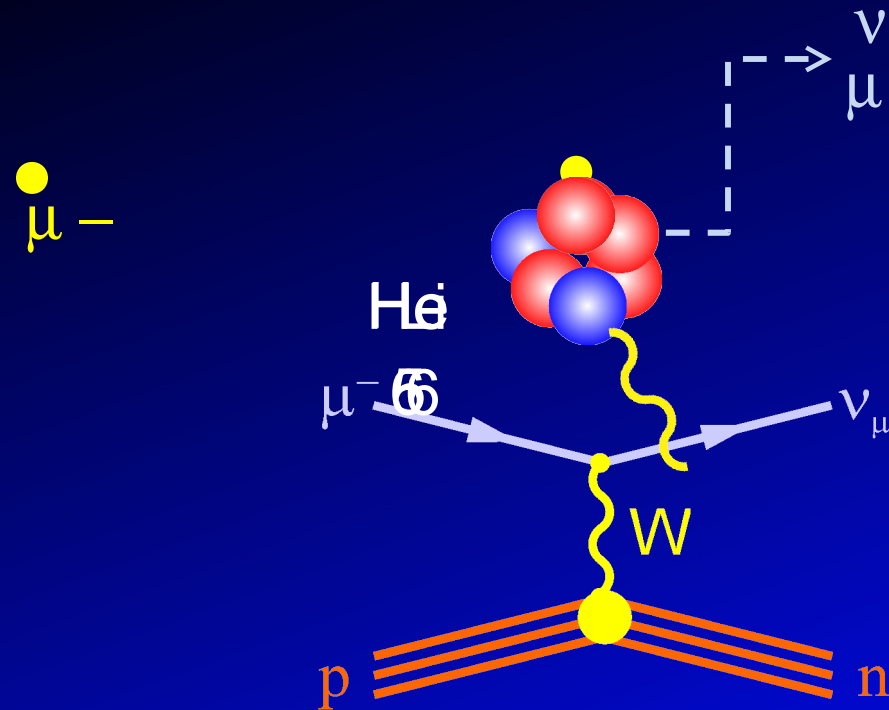
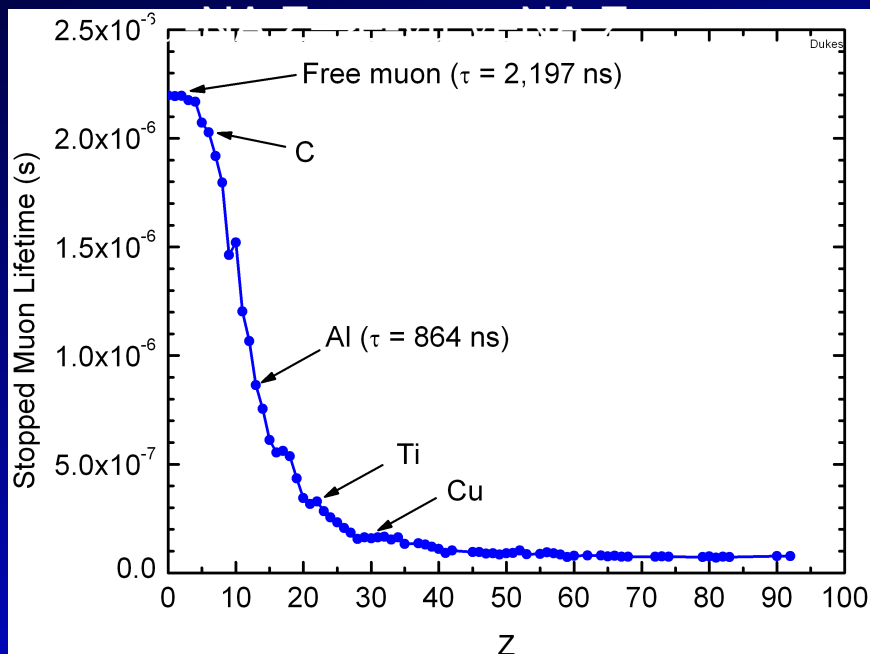


Littlest Higgs



How to Search for $\mu^- N \rightarrow e^- N$

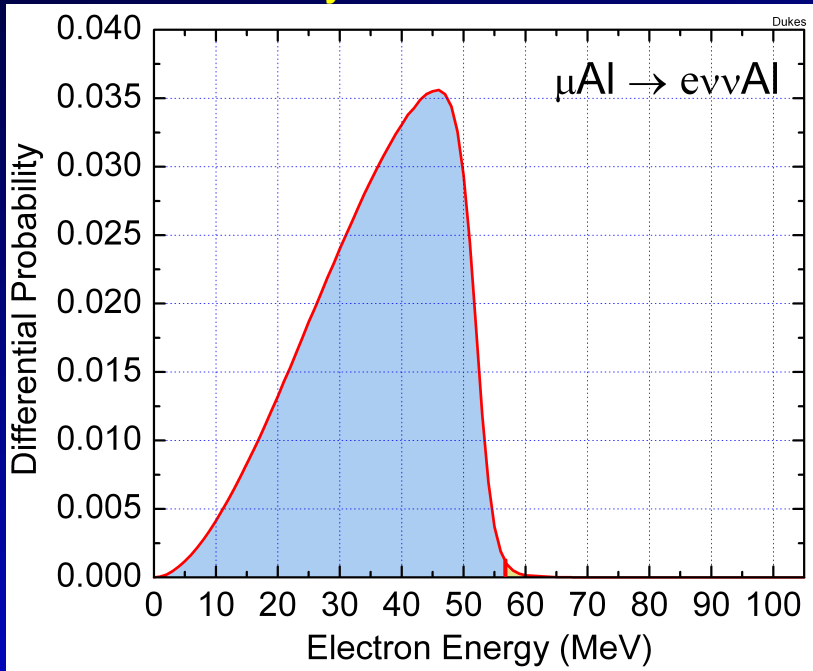
- Stop muon in atom
- Muon rapidly (10-16s) cascades to 1S state
- Circles the nucleus for up to $\sim 2 \mu\text{s}$
- Two things most likely happen:
 1. muon is captured by the nucleus:
 $\mu^- NA, Z \rightarrow \nu_\mu NA, Z-1$
 2. muon decays in orbit:



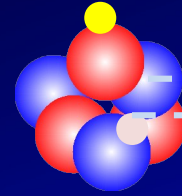
- New nucleus (N, Z-1) “de-excites” emitting neutrons, protons, and gammas
- **Al: $1.3 n + 0.1 p + 2\gamma$**
- These particles can produce “noise” in the detector

How to Search for μ -N \rightarrow e-N

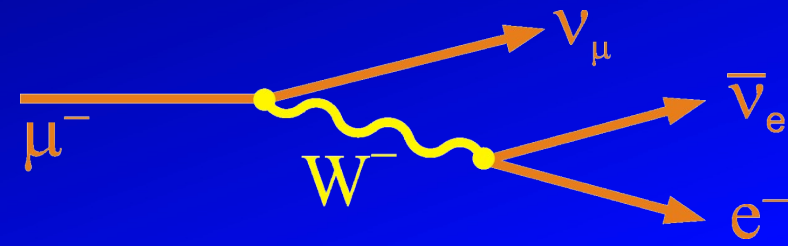
- Stop muon in atom
- Muon rapidly (10-16s) cascades to 1S state
- Circles the nucleus for up to $\sim 2 \mu$ s
- Two things most likely happen:
 1. muon is captured by the nucleus:
 $\mu^- - NA, Z \rightarrow \nu_{\mu} NA, Z-1$
 2. **muon decays in orbit:**



Li
6

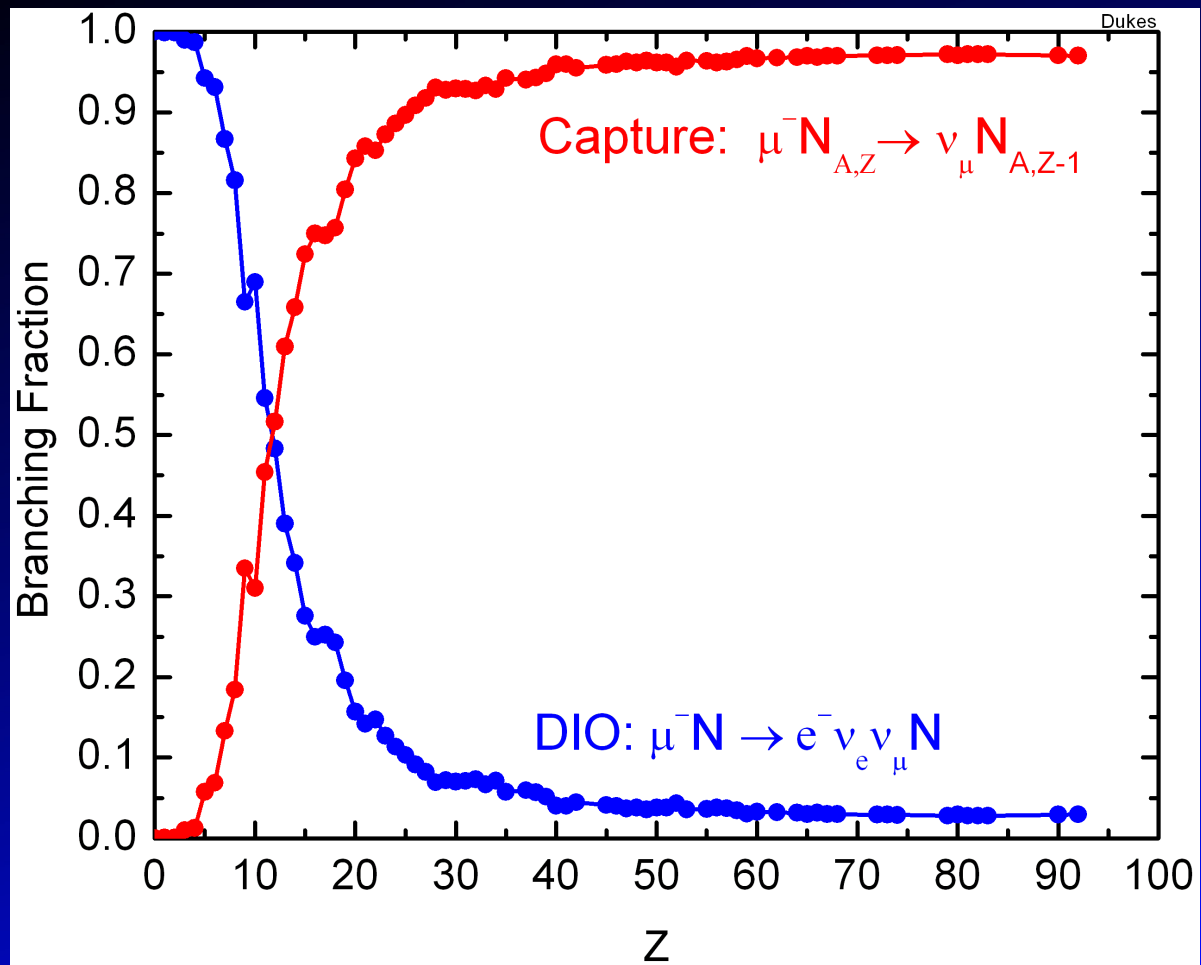


Note: tail not present in free decay due to recoil of nucleus



Muon Decay-in-Orbit

Relative Rate of Capture to DIO Depends on Target Z

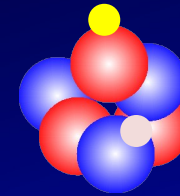


MUZE Searching for a Third Process: $\mu - N \rightarrow e - N$

N

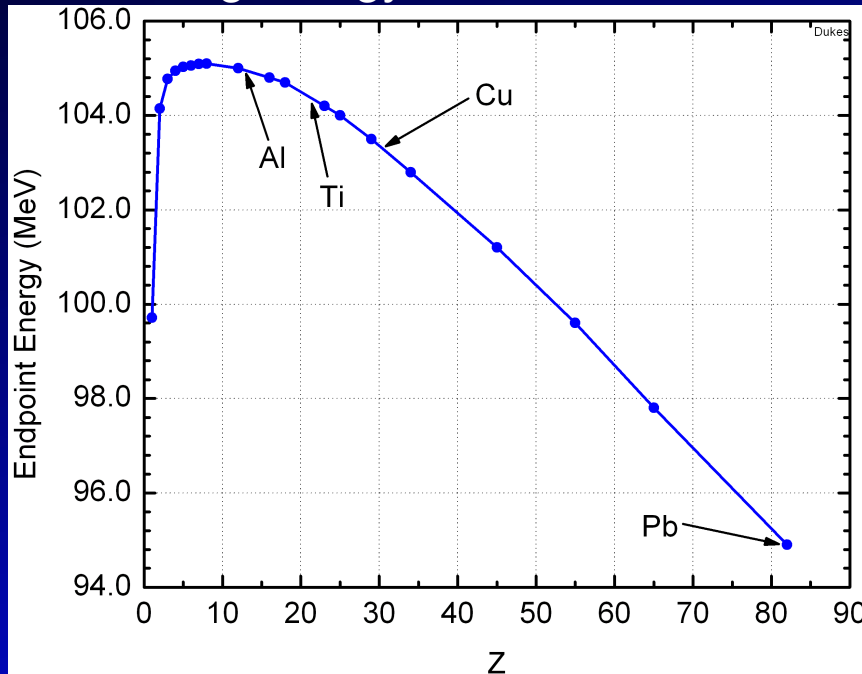
In $\mu - N \rightarrow e - N$ the muon coherently interacts with nucleus leaving it in ground state

- signature single isolated electron
- Electron energy given by the rest mass of the muon minus the nucleus recoil energy and the binding energy:



Li
6

e-



Measure ratio of conversion rate to capture rate

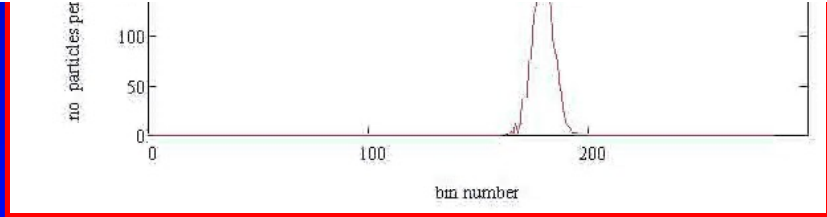
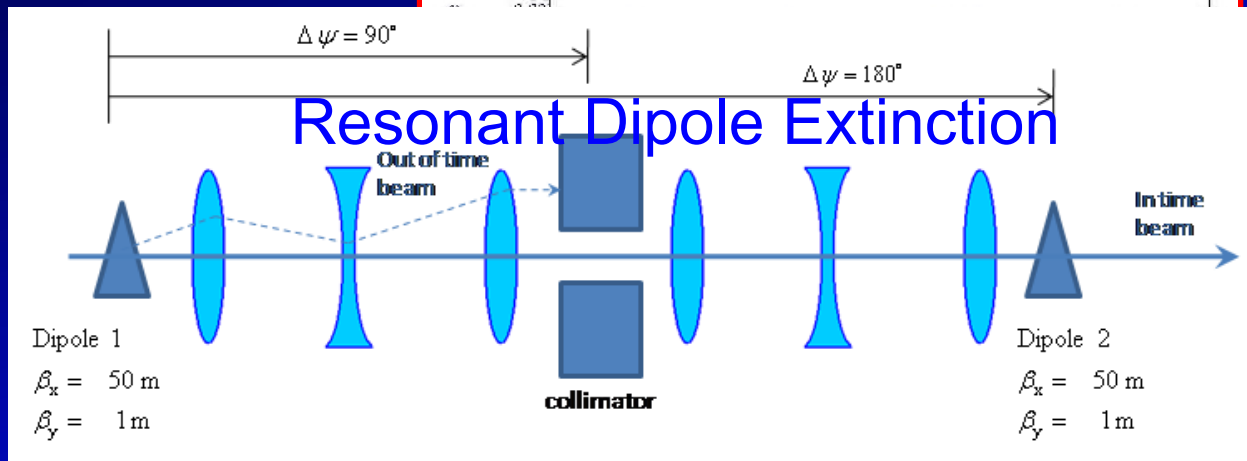
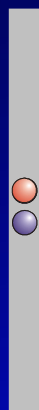
Bunched Beam Technique Needed

Need $\sim 10^{18}$ stopped muons

Signal: 105 MeV electron coming from the target, $\sim 1 \mu\text{s}$ after the μ is stopped in the foils

- Rate too high for continuous beam: need bunched muon beam: $50 \times 10^9 \mu / \text{s}$
- Need gate off detector for $\sim \text{tmN}$ ($\sim 800 \text{ ns}$) while bad stuff (pions, electrons) is around
- Ne

Bunched beam arrives with μ 's, π 's, and electrons



A huge amount of stuff comes off the target from scatters, captures etc.

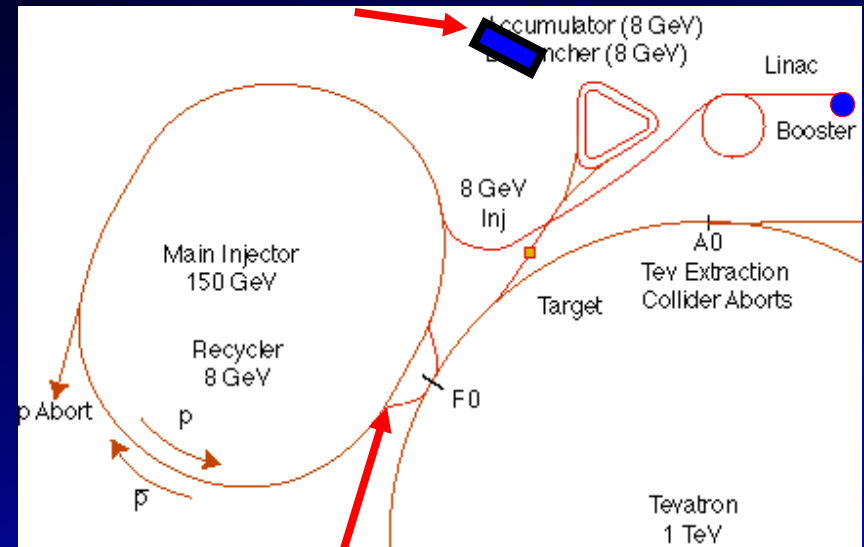
Stopping target

Need to be sure it scattered electron electron

Producing ~10¹⁸ Bunched Muons

- Energy: 8 GeV Booster beam optimal
- Structure: need bunch spacing on order of muon lifetime $\sim 1 \mu s$; orbit period of Accumulator/Debuncher optimal at 1.7 μs
- Use Recycler as a transfer line
- Stack in Accumulator; bunch in Debuncher
- Slow spill extraction: 90% duty factor

new detector hall and beamline

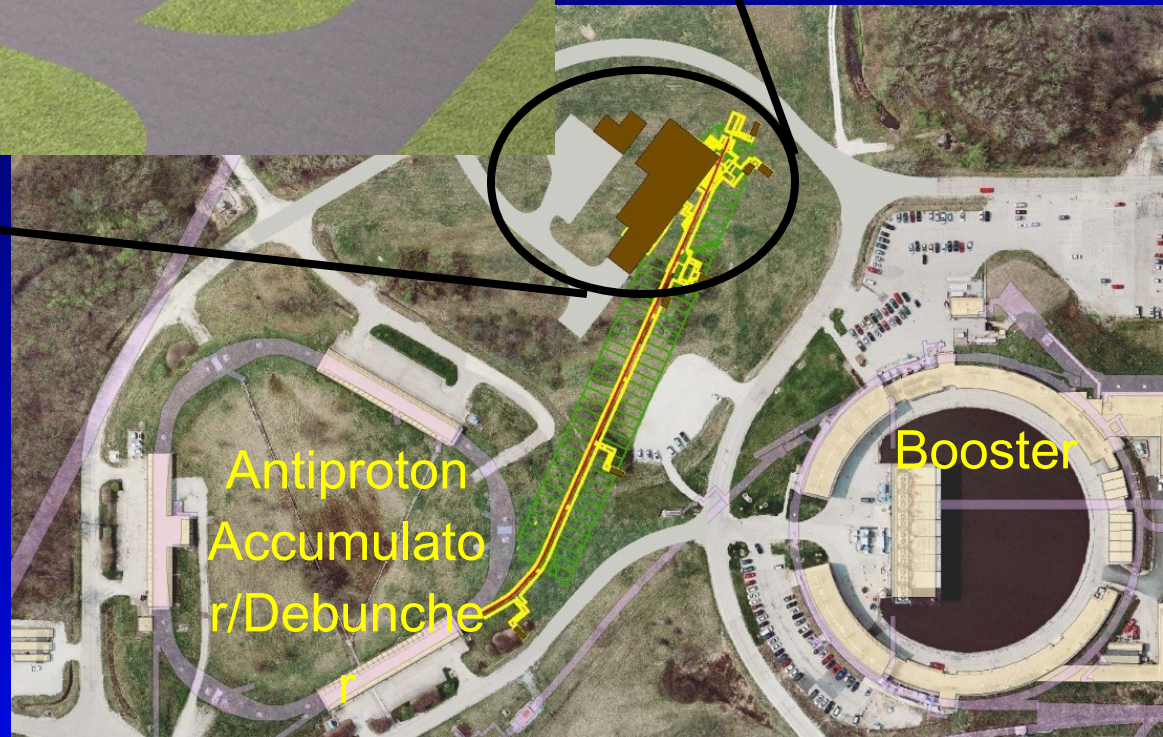


One of several possible schemes

Mu2e cannot take any protons away from the neutrino program

Mu2e cannot take any protons away from the neutrino program
 NOvA cannot use during Main Injector cycle

New Beamline and Detector Hall to be Built



Backgrounds, Backgrounds, Backgrounds . . .

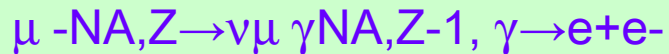
1. Stopped Muon Backgrounds

Muon decay in orbit (DIO):



- defeated by good energy resolution

Radiative muon capture (RMC):



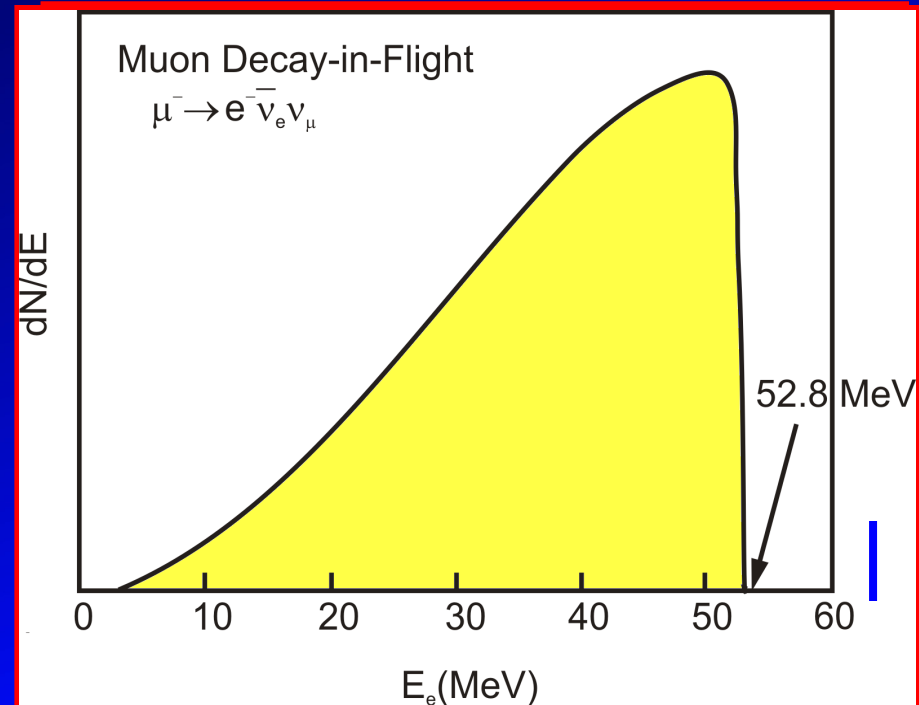
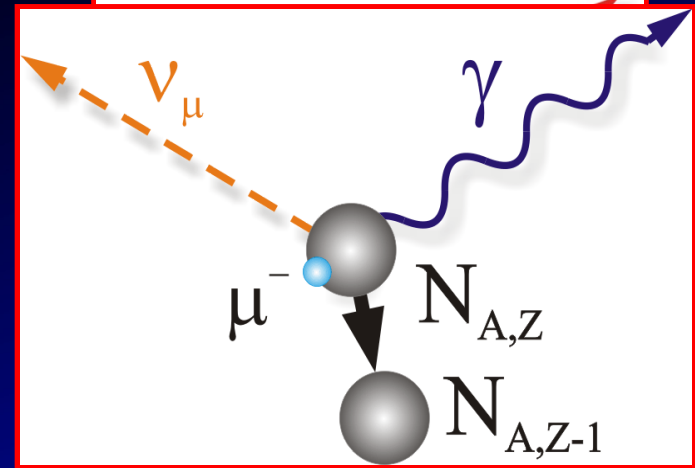
- defeated by $mZ < mZ-1$
- defeated by good energy resolution

Muon capture:



2

- Note: does not produce a physics background, but detector “noise”
- defeated by proton absorber



Backgrounds, Backgrounds, Backgrounds . . .

2. Prompt Beam Related Backgrounds

Radiative pion capture (RPC):

$$\pi^- N(A,Z) \rightarrow \gamma N(A,Z-1), \quad \gamma \rightarrow e^+ e^-$$

Note: 1.2% have $E_\gamma > 105 \text{ MeV}$

Muon decay in flight:

$$\mu^- \rightarrow e^- \nu \bar{\nu}$$

Note $p_\mu > 77 \text{ MeV}/c$

Pion decay in flight:

$$\pi^- \rightarrow e^- \bar{\nu}_e$$

Beam electrons scattering in target

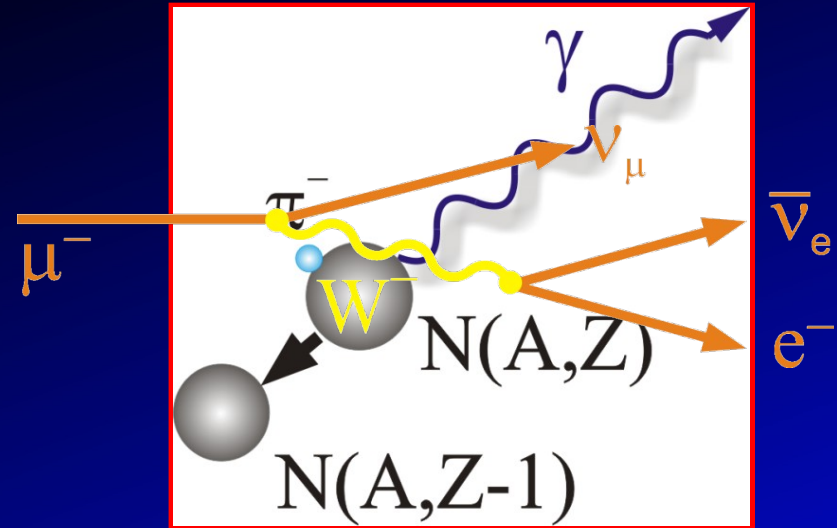
- > Defeated by 10⁻¹⁰ interbunch extinction
- > Defeated by hard cuts on momentum

Antiprotons annihilating

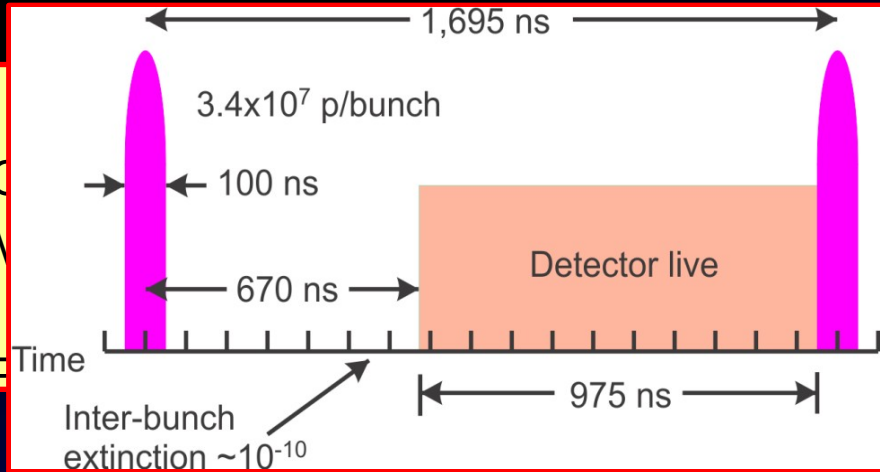
3. Time Dependent Backgrounds

Cosmic Rays

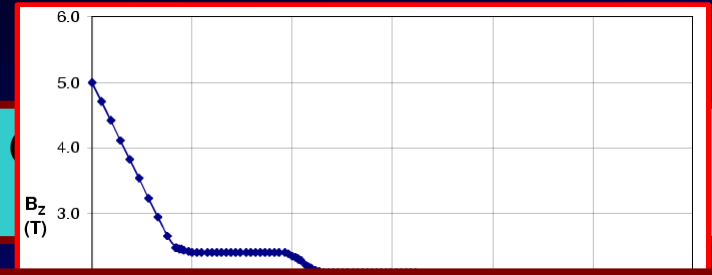
- > Defeated by active shield



Mu2e Apparatus



8



Look for 105 MeV electron spiraling through detector

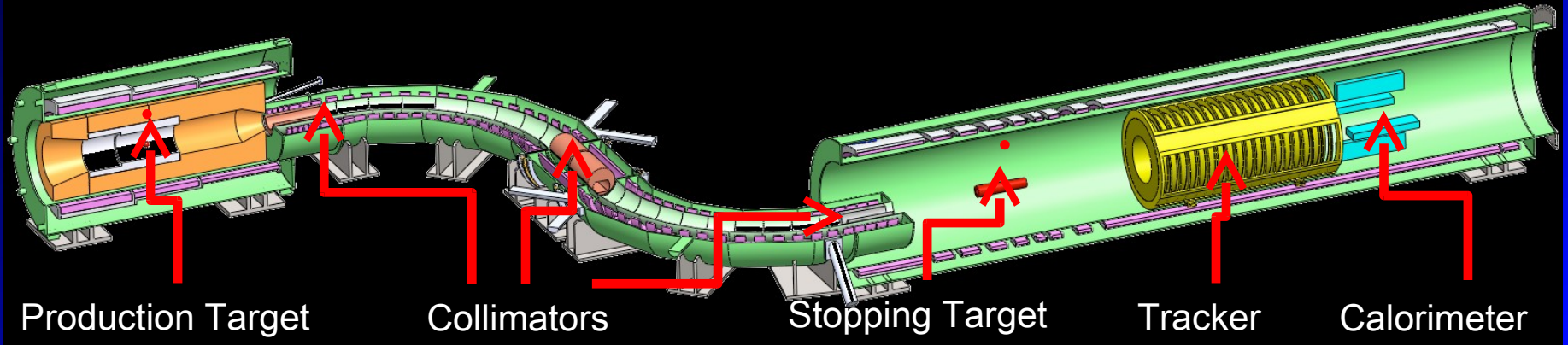
Muon Beam

Spectrometer

Production Solenoid

Transport Solenoid

Detector Solenoid



Production Target

Collimators

Stopping Target

Tracker

Calorimeter



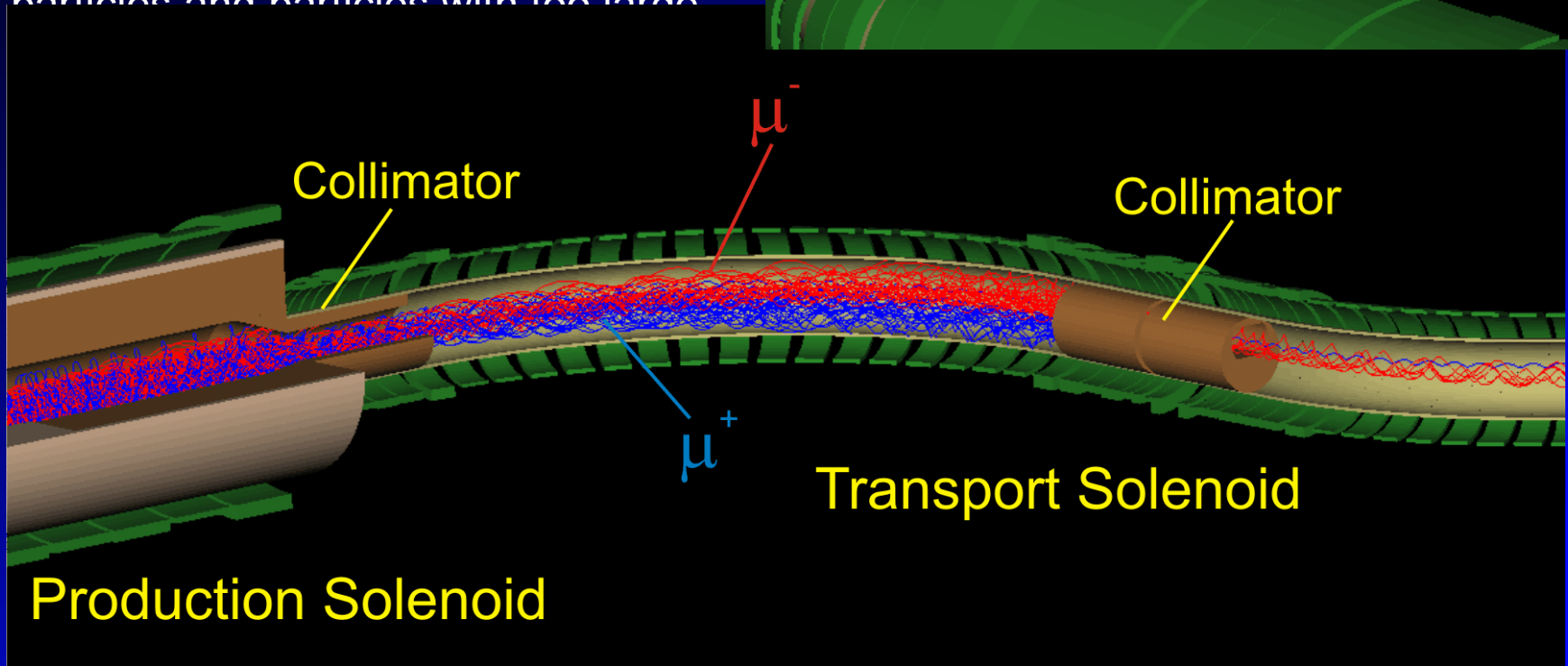
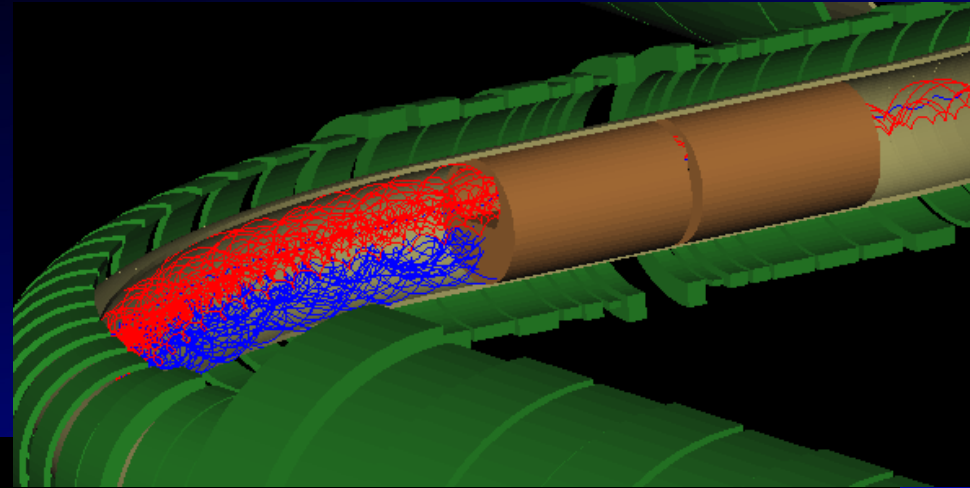
85,000 μ – stop every 1.7 μ s
50 billion μ – stops/second

Transport Solenoidal Magnet

·Curved solenoid:

1. separates charges by charge sign
2. reduces line-of-sight transport of neutrals

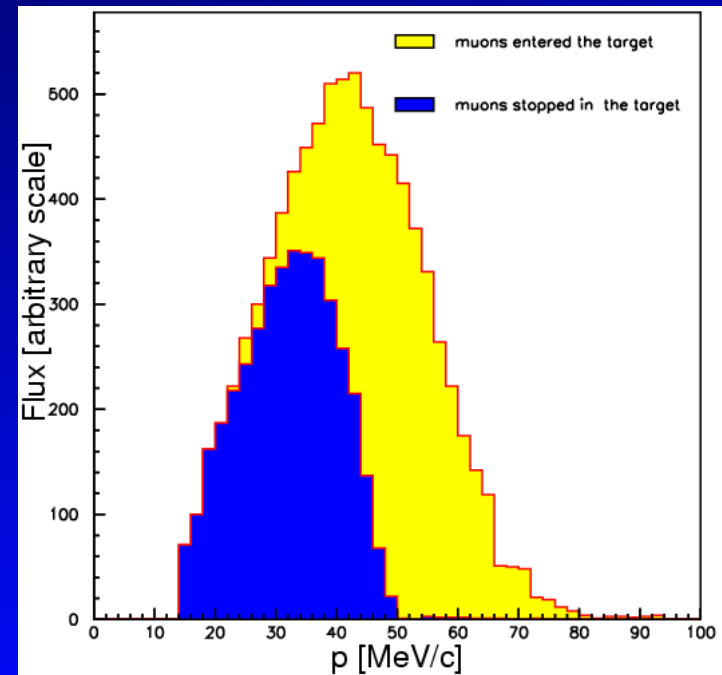
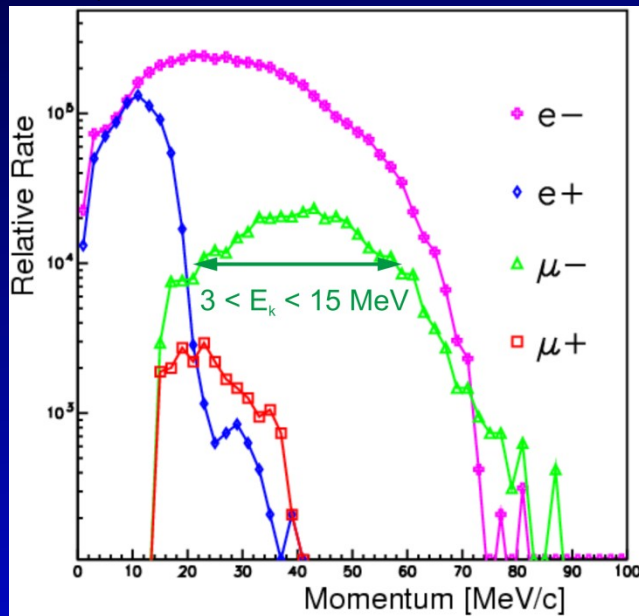
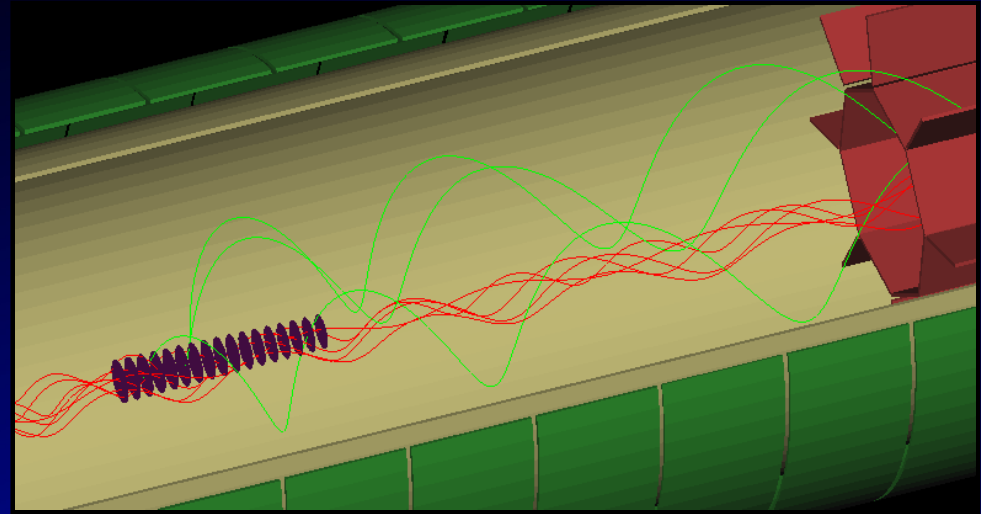
·Collimators eliminate wrong-sign particles and particles with too large



What we get at the Stopping Target

17 Al disks
each 200 μ m thick
83 mm to 65 mm radius
in graded magnetic field

1/230 incident protons produce a muon at the stopping target
58% of muons stop in target
50x10⁹ μ stops per spill second
85,000 μ stops per microbunch



Choice of Stopping Target Material

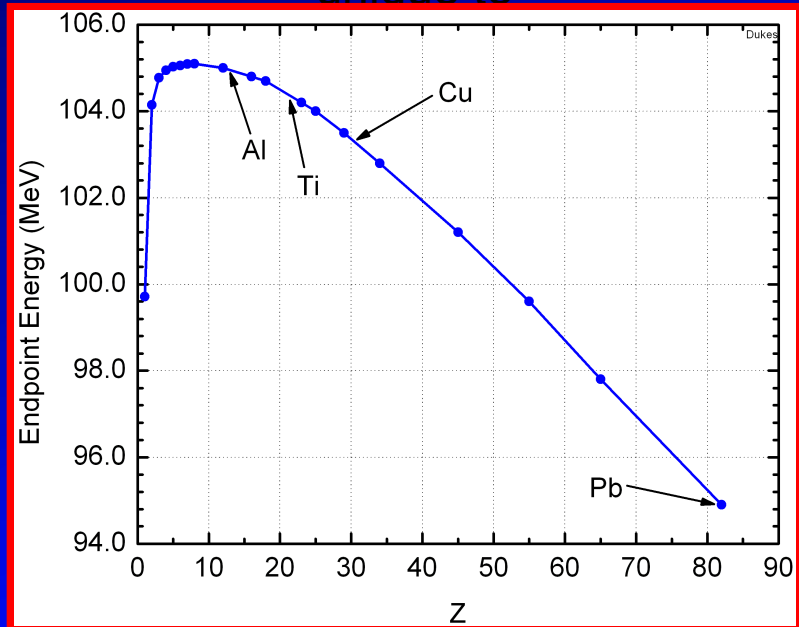
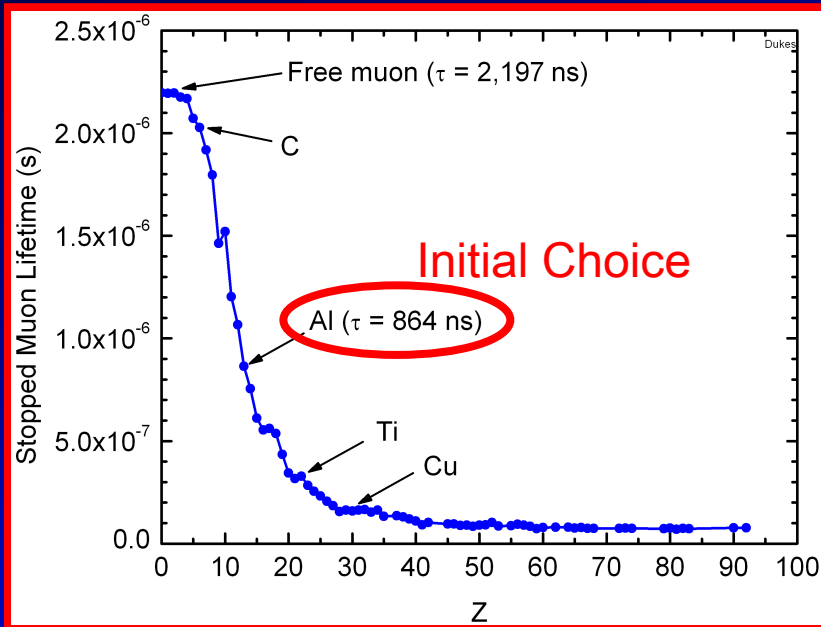
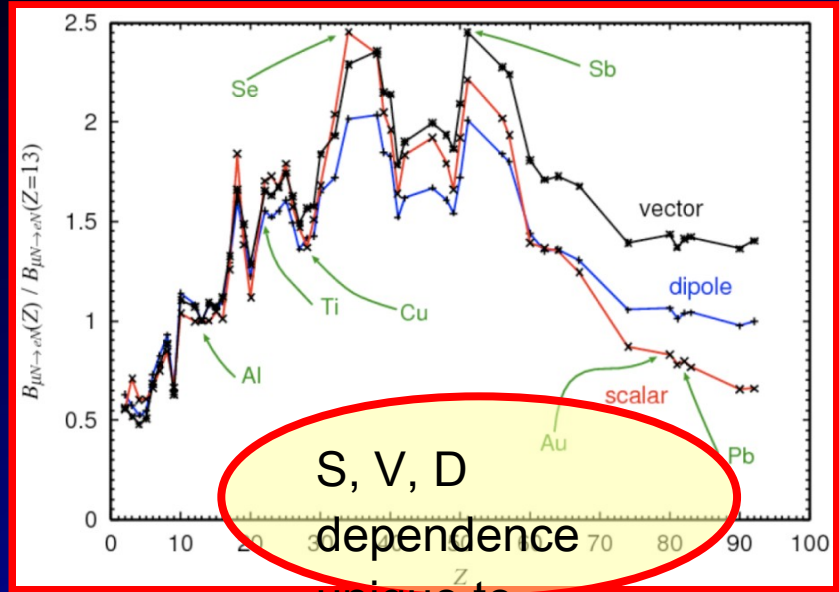
Large Z:

- rate $\propto Z|F_n|^2$ (F_n is the form factor)
- can reveal nature of interaction

Small Z:

- longer lifetime
- higher endpoint energy

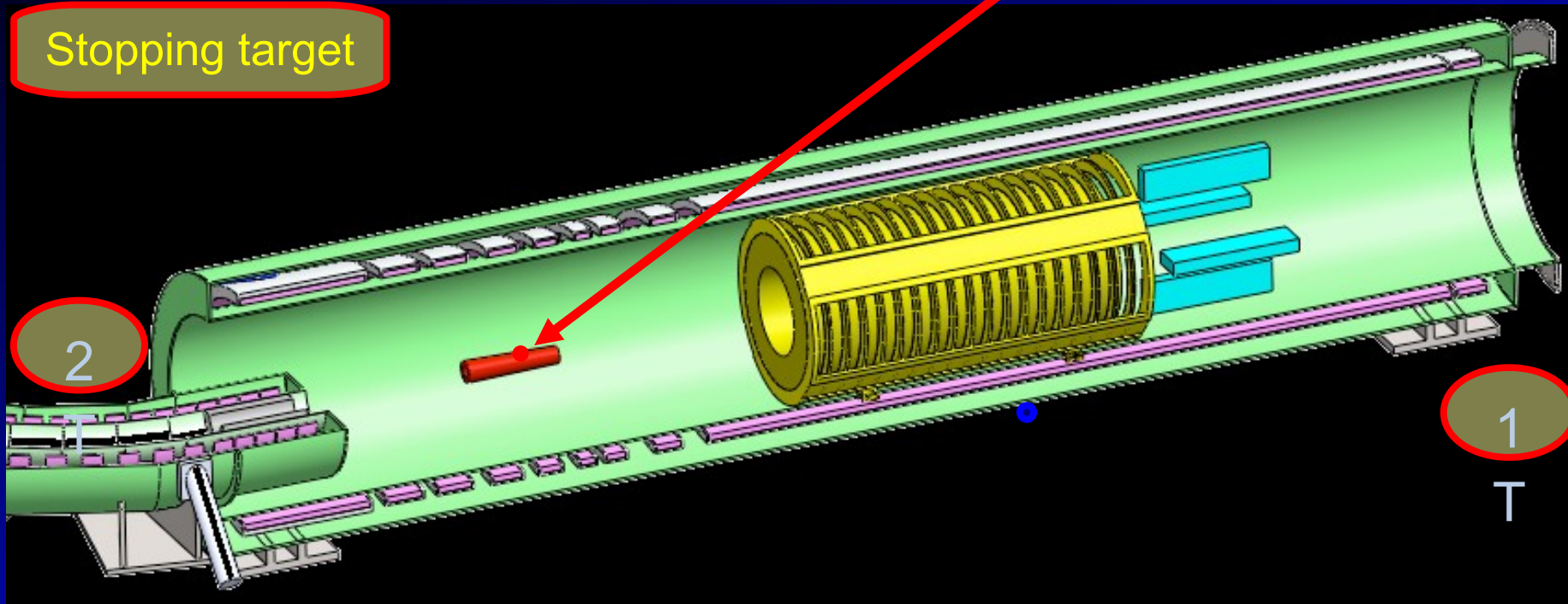
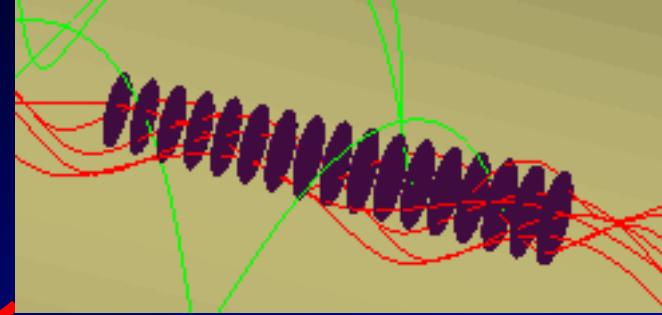
Note: Need $mZ-1 > mZ$ to place max. energy of radiative capture muons below signal electrons



Mu2e Spectrometer

Salient Features

- No detector element in region of transported beam
- Small acceptance for DIO electrons
- Minimal amount of material ζ detector elements in vacuum



Stopping target

2

1

T

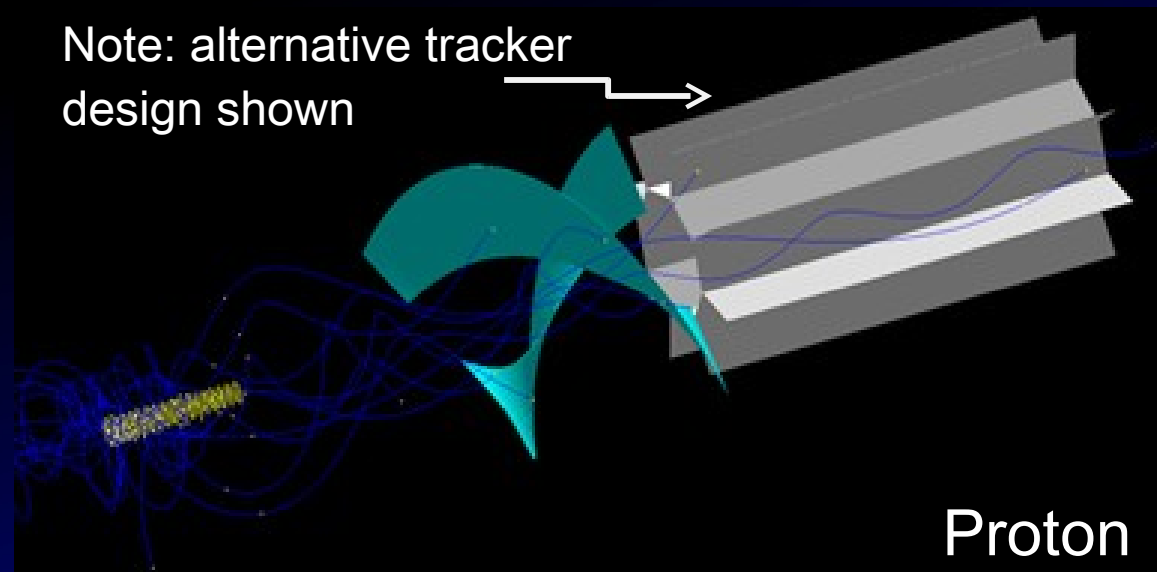
Proton
absorber

Straw tracker

Electromagnetic
calorimeter

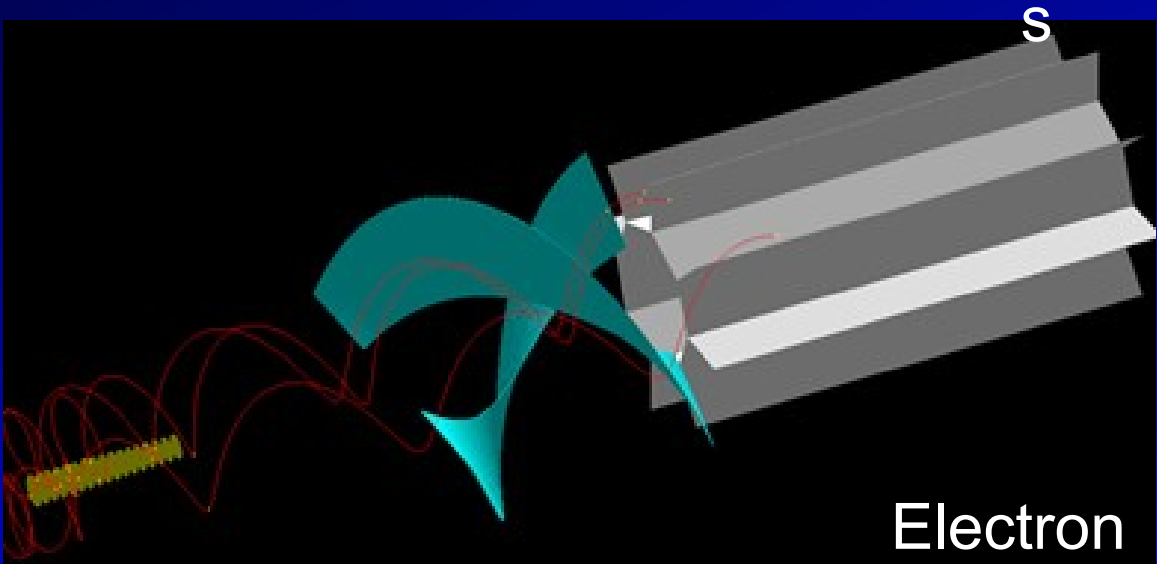
Helical Proton Absorber

Note: alternative tracker design shown



Proton

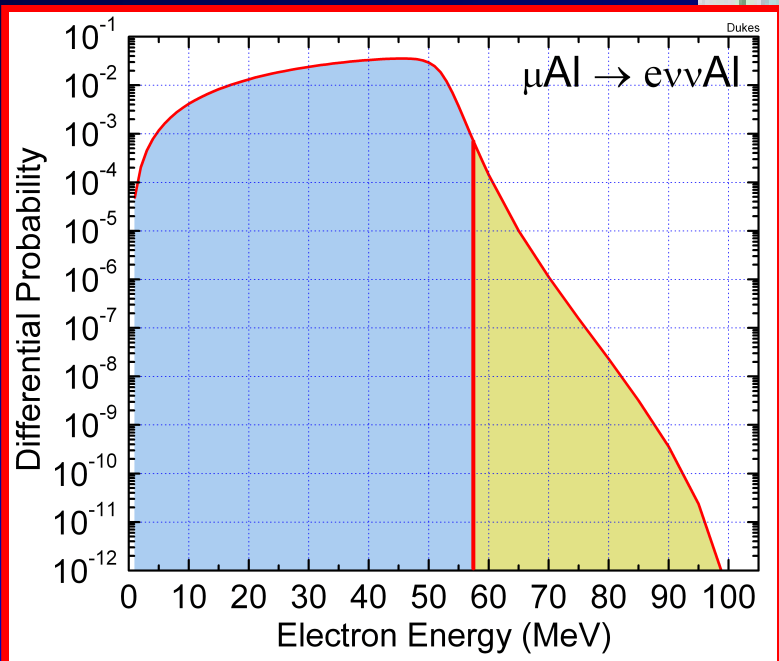
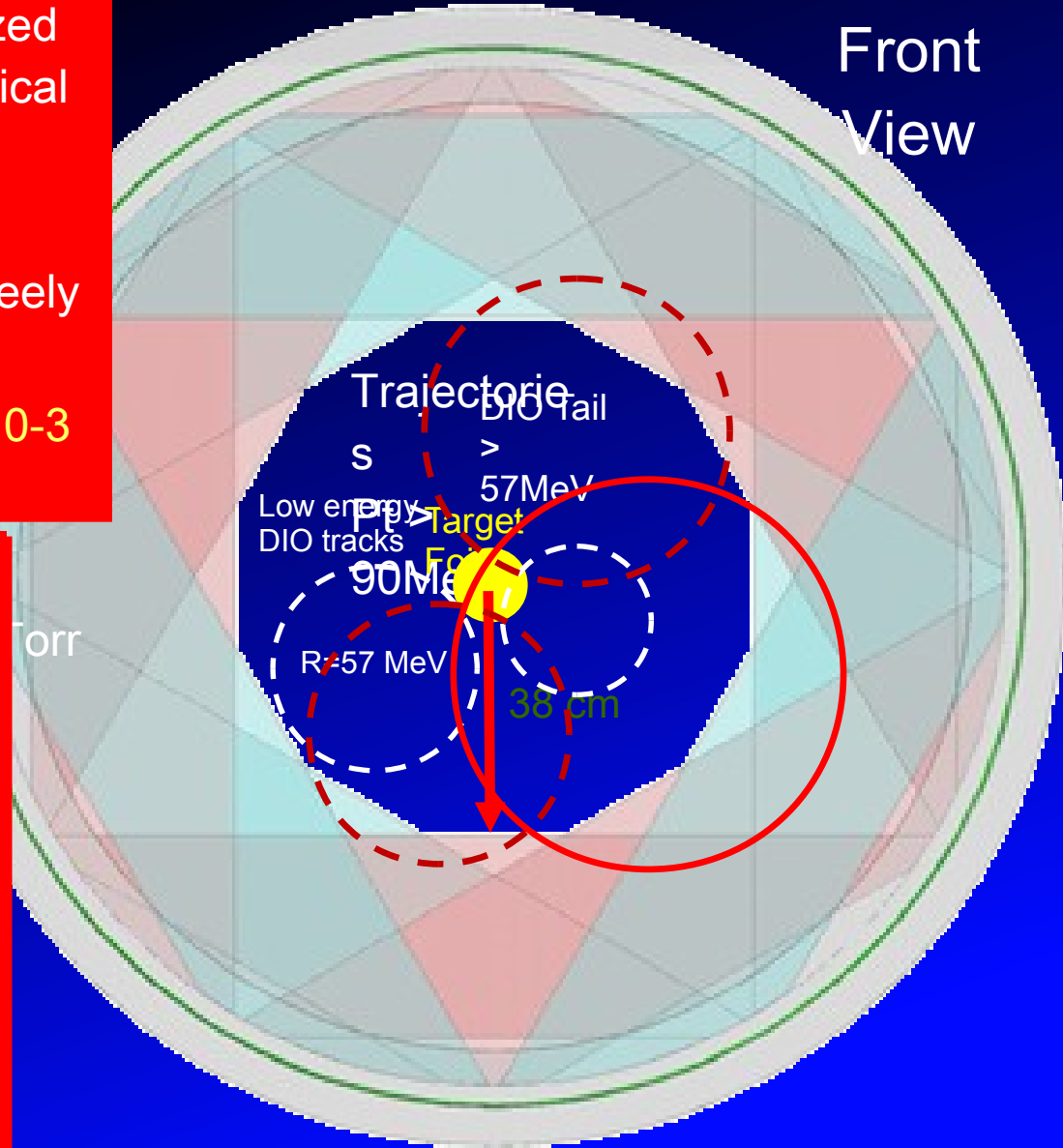
- Needed to beat down rate of muon capture protons: ~ 0.1 per capture: $\sim 3 \times 10^9/s$
- Note: capture protons not relativistic: KE and momentum not the same!
- Geometry exploits different helical trajectories of protons and electrons
- Material: 0.5 mm polyethylene
- Absorbers 93% of protons from muon capture
- Only 13% of the electrons impact absorber
- FWHM energy loss: 0.33 MeV



Electron

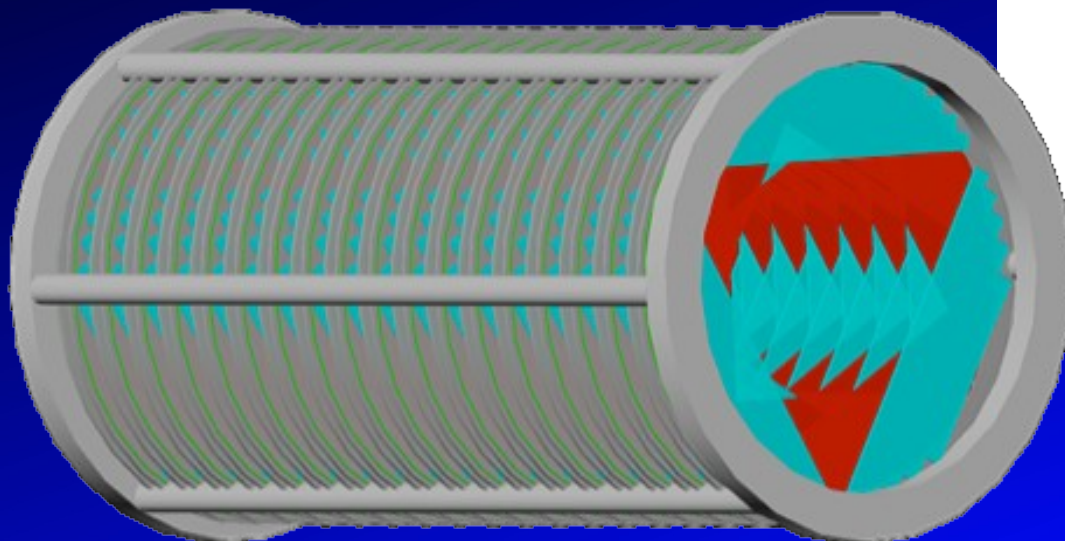
Heart of the Spectrometer: Straw Tracker

- Octagonal vane geometry optimized for reconstruction of 105 MeV helical trajectories
- Center beam region empty:
 - Intense muon beam passes freely through
 - Acceptance for DIO tracks $< 10^{-3}$
- Severe operating environment:

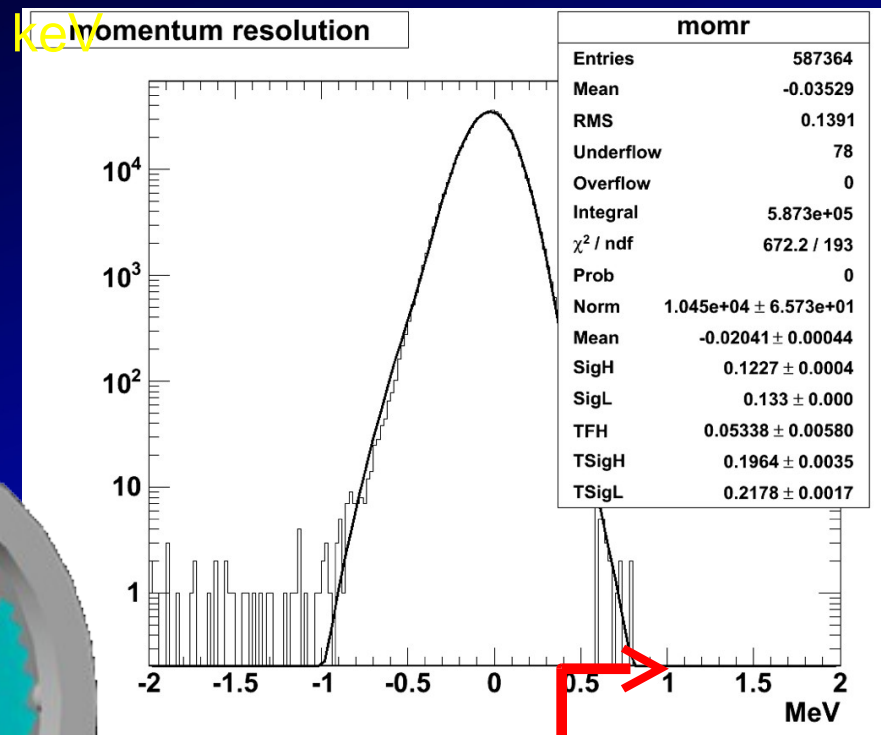


Tracker Performance

- 5 mm diam. Straws
- 18 stations:
 - 24 planes each
 - 50 straws/plane
- 21,600 straws total
- ArCO₂ 80:20
- ADC/TDC readout: 100 μ m resolution



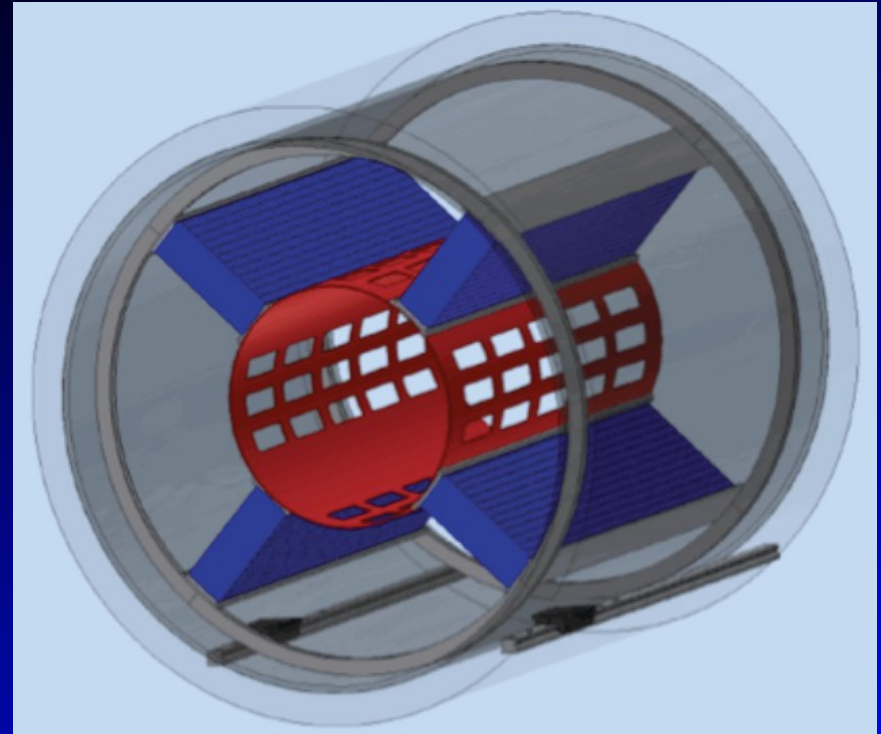
Large acceptance: 50% ($90^\circ \pm 30^\circ$)
Intrinsic energy resolution: $\sigma = 140$



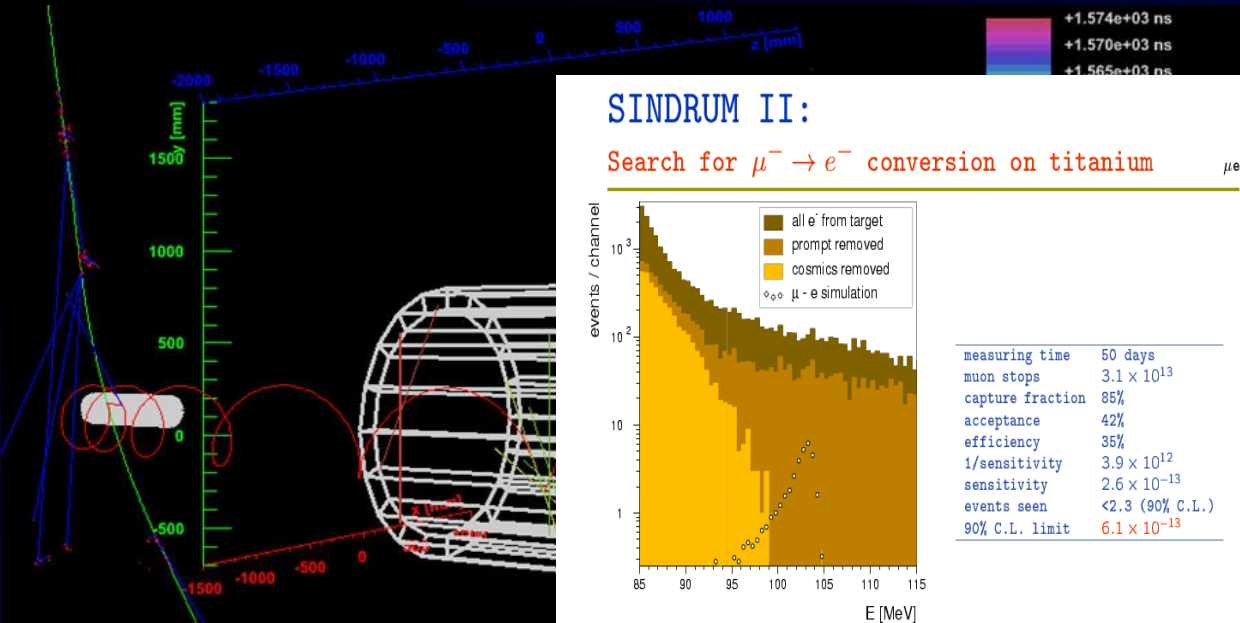
No high-side tail!

Electromagnetic Calorimeter

- Needed for:
 - trigger: 5% energy resolution → 1,000 triggers/s
 - Timing
 - Confirmation of the electron energy measurements of the straws
- Note: must operate in 160 Gy/yr
- 2112 3x3x13cm³ LYSO crystals
- Dual APD readout @ 100 MeV
- $\sigma E \sim 2-3\%$
- $\sigma \sim 0.1$ ns

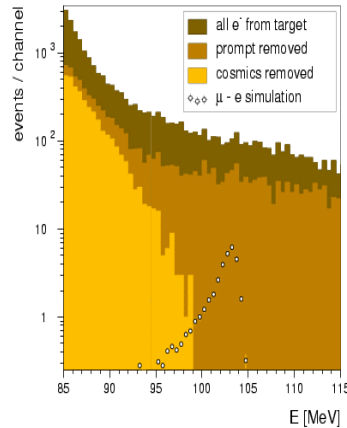


Cosmic Ray Veto



SINDRUM II:

Search for $\mu^- \rightarrow e^-$ conversion on titanium μe

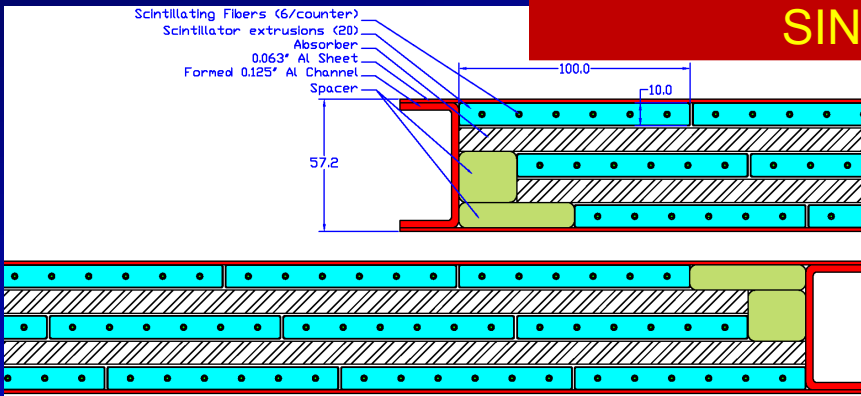


measuring time	50 days
muon stops	3.1×10^{13}
capture fraction	85%
acceptance	42%
efficiency	35%
1/sensitivity	3.9×10^{12}
sensitivity	2.6×10^{-13}
events seen	<2.3 (90% C.L.)
90% C.L. limit	6.1×10^{-13}

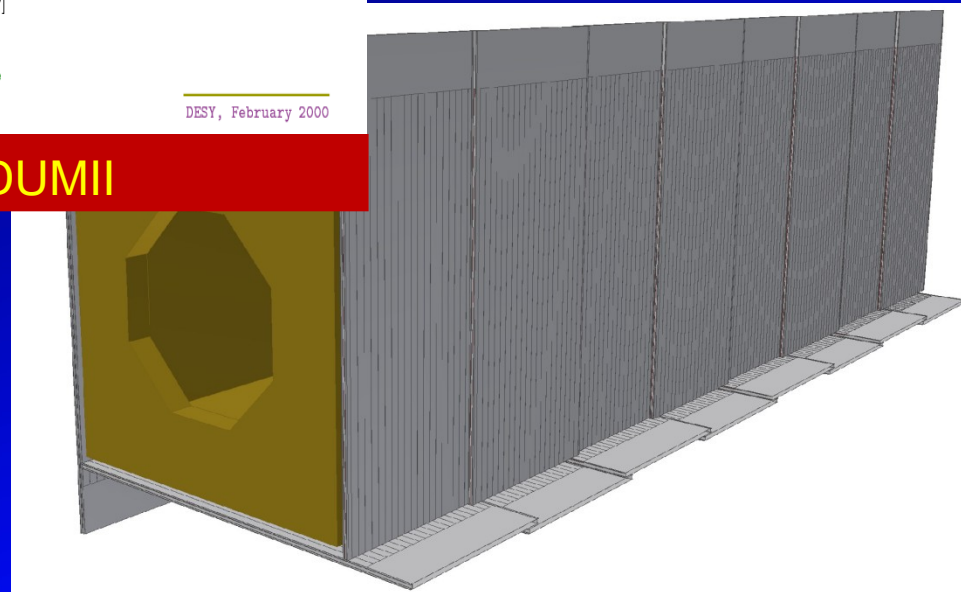
Electron energy distribution at three stages of the event selection and as predicted by a GEANT simulation of μe conversion at $B_{\mu e} = 4 \cdot 10^{-12}$.

DESY, February 2000

MC studies show 416 fake 105 MeV conversion electrons from cosmic-ray muons
Hence need 10^{-4} Inefficiency Surround Detector Solenoid by 3 layers of



SINDUMII



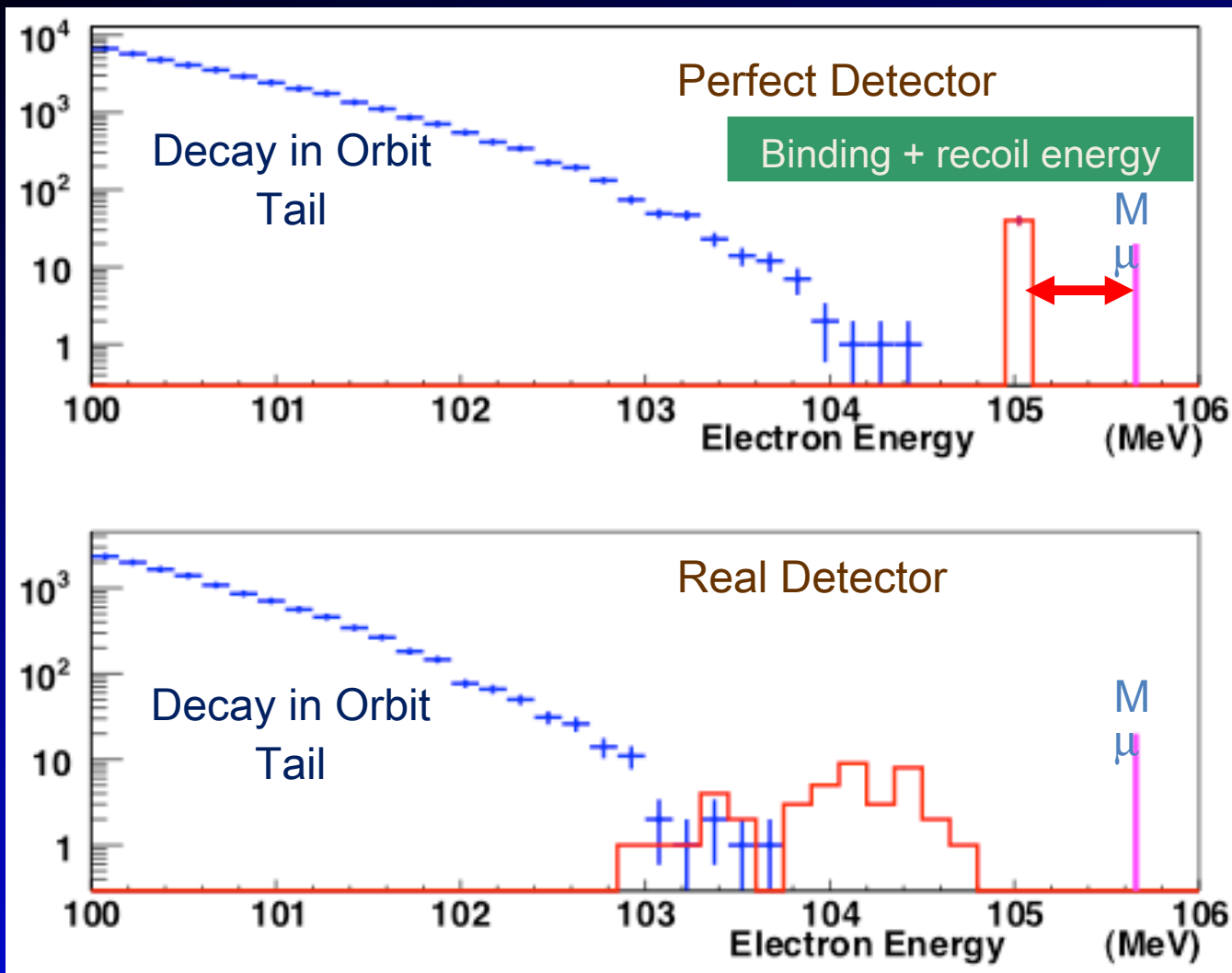
What we Get

Intrinsic tracker
energy resolution:

$$\sigma(E) \lesssim 150 \text{ keV}$$

Average energy loss
due to spectrometer
material:

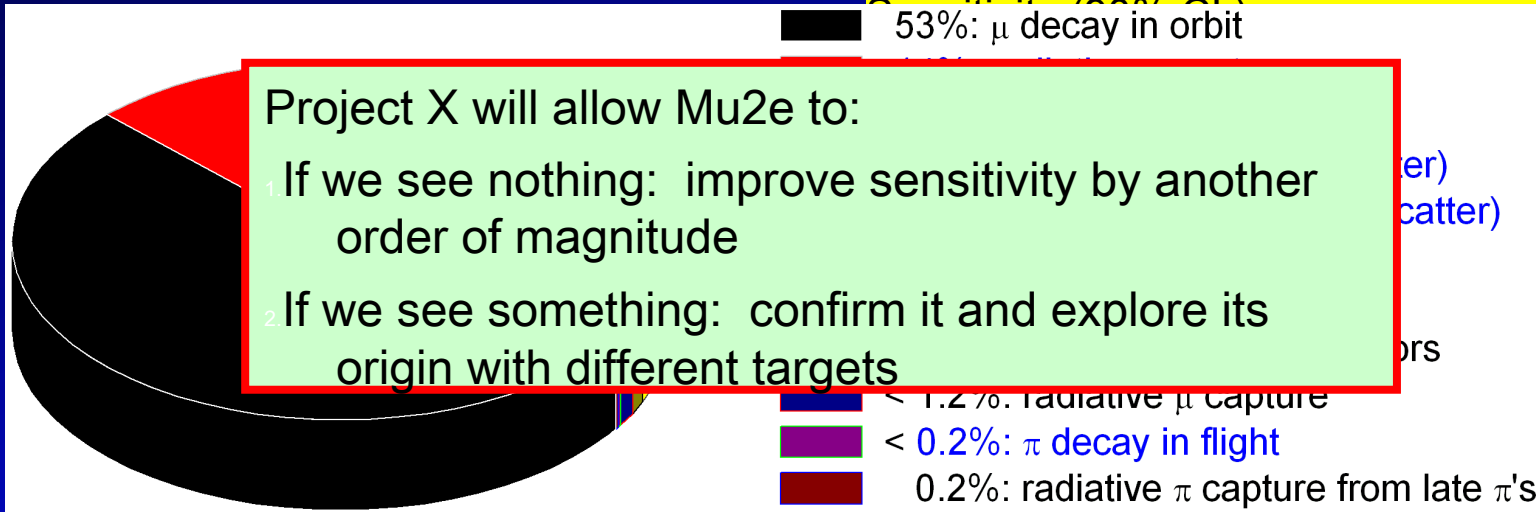
$$E(\text{shift}) \lesssim 1 \text{ MeV}$$



Mu2e Sensitivity

Roughly half of background is interbunch contamination related

Proton flux	1.8x10 ¹³ p/s
Running time	2x10 ⁷ s
Total protons	3.6x10 ²⁰ p
μ - stops/incident proton	0.0025
μ - capture probability	0.61
Time window fraction	0.49
Electron trigger eff.	0.80
Reconstruction and selection eff.	0.19
Overall sensitivity (90% CL)	6x10⁻¹⁷
	4
	0.4



Project X will allow Mu2e to:

- 1 If we see nothing: improve sensitivity by another order of magnitude
- 2 If we see something: confirm it and explore its origin with different targets

Mu2e Collaboration

Boston University R. Carey, V. Jorjadze, V. Khalatian, J. Miller*, B. Roberts

Brookhaven National Laboratory W. Marciano, V. Polychronakos, Y. Semertzidis, P. Yamin

Caltech D. Hitlin, F. Porter, R.-Y. Zhu

Currently:

134 scientists

28 institutions

University of Californ
University of Cal
City University

Fermi National A

W

DZERO

CDF

Institute for Nucle

DELPHI

CM

ATLAS

Univers



0

500

1,000

1,500

2,000

2,500

3,000

Size of collaboration

Jongh, N.
petz, R.
ee, S.
ebys, V.
White, K.

S. Rella, G.

hi

shikawa

Northern Illinois University D. Hedin, A. Dychkant

Northwestern University A. deGouvea

INFN Pisa, Università Di Pisa F. Cervelli, R. Carosi, M. Incagli, T. Lomtadze, L. Ristori, F. Scuri, C. Vannini

Rice University M. Corcoran

Syracuse University R.S. Holmes, P.A. Souder

Università di Udine D. Cauz, G. Pauletta


University of Virginia M. Bychkov, E.C. Dukes, R. Ehrlich, M. Frank, E. Frlez, C. Group, R. Hirosky, P.Q. Hung, Y.

Craig Dukes / Virginia

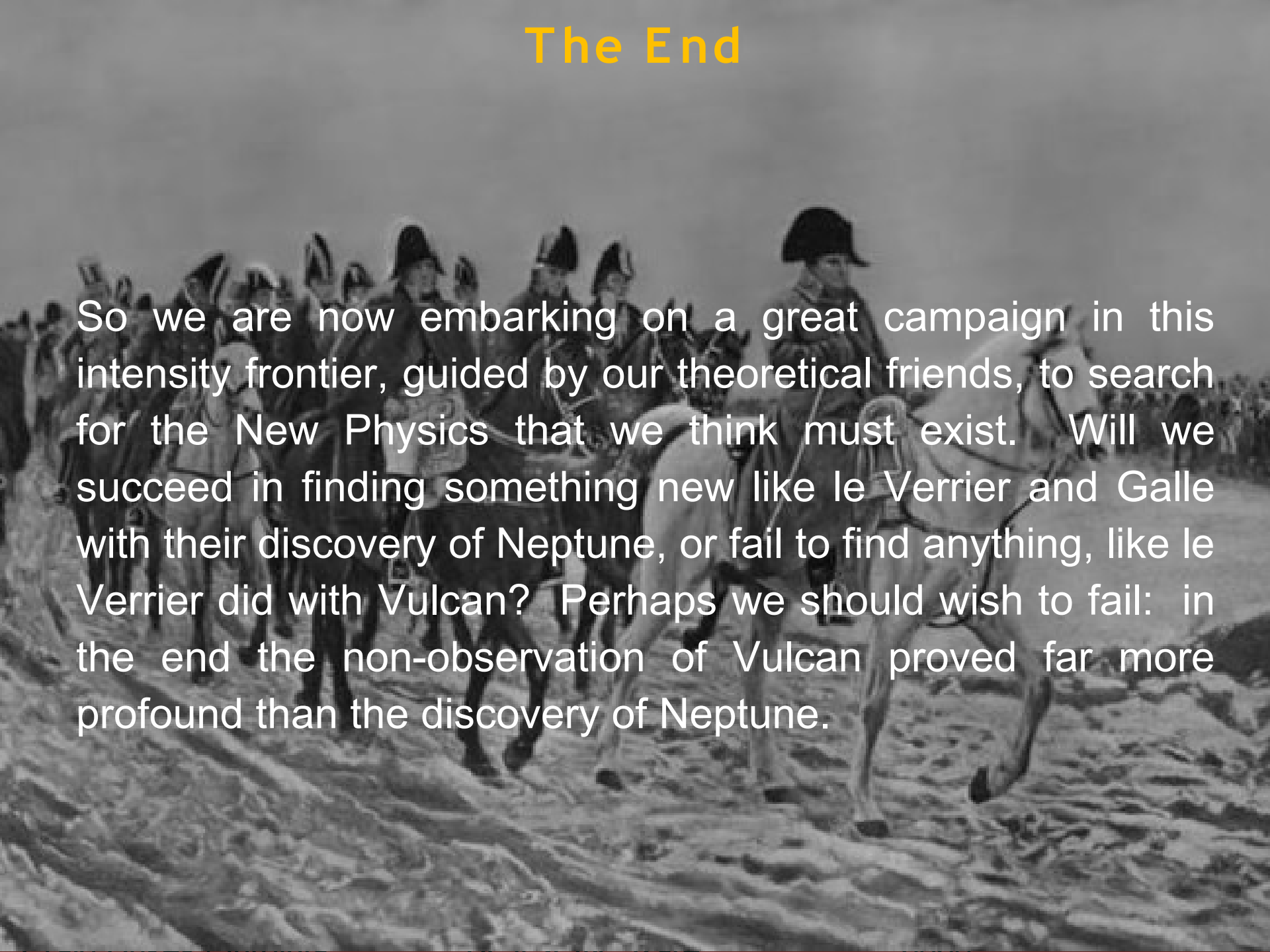
Oksuzian, K. Paschke, D. Pocanic

5858

Mu2e Status

1992	Solenoidal collection scheme first proposed for Moscow Meson Factory	 <p>US Particle Physics: Scientific Opportunities A Strategic Plan for the Next Ten Years</p> <p>Report of the Particle Physics Project Prioritization Panel</p> <p>29 May 2008</p>
1997	MECO proposed for the AGS at Brookhaven	
1998-2005	Intensive work on MECO technical design. \$58M, detector at \$27M	
July 2005	RSVP cancelled for financial reasons not met	
2006	Steering committee set up to work out meeting with Fermilab, keeping detector the same	
June 2007	Mu2e EOI submitted to Fermilab	
October 2007	LOI submitted to Fermilab	
May 2008	P5 "recommends pursuing the muon-to-electron conversion experiment subject to approval by the Fermilab PAC, and is to be considered by the panel."	
Fall 2008	Proposal submitted to Fermilab and receives Stage I approval. Total project cost estimated at \$180M	
November 2009	Approval of mission need by DOE (CD-0) granted (\$145M - \$205M)	
FY2010	Mu2e receives \$4M in R&D funding (\$10M FY2011, \$20M FY2012)	
Fall 2011	Approval of preliminary baseline range by DOE (CD-1)	
2014	Start of Construction (CD-3)	

The End



So we are now embarking on a great campaign in this intensity frontier, guided by our theoretical friends, to search for the New Physics that we think must exist. Will we succeed in finding something new like le Verrier and Galle with their discovery of Neptune, or fail to find anything, like le Verrier did with Vulcan? Perhaps we should wish to fail: in the end the non-observation of Vulcan proved far more profound than the discovery of Neptune.