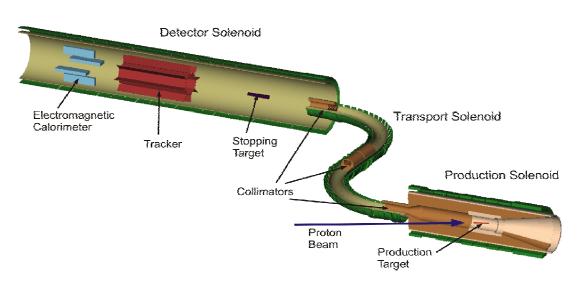
The Mu2e Experiment at FNAL

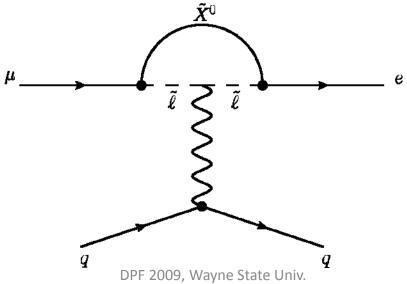
A precision window into physics beyond the standard model



Andrew Norman
University of Virginia
For the Mu2e Collaboration

Why Precision Measurements & Ultra-Rare Processes?

- We want to access physics beyond the standard model
 - This means access to High and Ultra-High Energy interactions
 - One way to get to these energies through loops
 - Getting at Loops means making precision measurements and looking for ultra-rare decays
- Ideally we start with processes that are forbidden or highly suppressed in the standard model
 - Any observation becomes proof of non-SM physics



Why Precision Measurements & Ultra-Rare Processes?

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- Ideally we start with processes that are forbidden or highly suppressed in the standard model
 - Any observation becomes proof of non-SM physics
- Flavor Changing Neutral Currents
 - FCNC in quark sector
 - $B_s \rightarrow \mu \; \mu$, $b \rightarrow s \; \gamma$, $K \rightarrow \pi \; \nu \; \nu$
 - Allowed but HIGHLY suppressed in Standard Model
 - Can receive LARGE enhancements in SUSY and other beyond-SM physics
 - FCNC in charged lepton sector
 - $\mu \to {\rm e} \; \gamma$, $\mu \to {\rm e} \; {\rm e} \; {\rm e} \; {\rm e} \; {\rm h} \; {\rm N} \to {\rm e} \; {\rm N}$ (Lepton Flavor Violating)
 - No SM amplitudes (except via ν loops)
 - Permitted in beyond-SM models, and have extreme reach in energy

Lepton Mixing in the Standard Model

We have three generations of leptons:

$$\left(\begin{array}{c} e \\ \nu_e \end{array} \right) \left(\begin{array}{c} \mu \\ \nu_\mu \end{array} \right) \left(\begin{array}{c} \tau \\ \nu_ au \end{array} \right)$$
 No SM couplings between generation!

- In the standard model Lagrangian there is no coupling to mixing between generations
- But we have explicitly observed neutrino oscillations
- Thus charged lepton flavor is **not** conserved.
- Charged leptons must mix through neutrino loops

$$Br(\mu \to e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{\ell} V_{\mu\ell}^{\star} V_{e\ell} \frac{m_{\nu_{\ell}}^2}{M_W^2} \right|^2$$

$$\leq 10^{-54}$$

But the mixing is so small, it's effectively forbidden

Charged Lepton Flavor Violation (CLFV) Processes with μ 's

• There are three basic channels to search for μ -CLFV in:

$$\mu^{+} \rightarrow e^{+} \gamma$$

$$\mu^{+} \rightarrow e^{+} e^{+} e^{-}$$

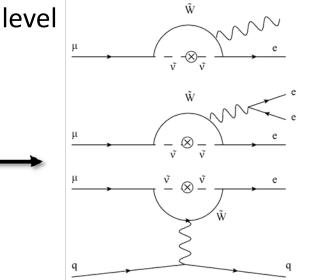
$$\mu^{-} N \rightarrow e^{-} N$$

 If loop like interactions dominate we expect a ratio of rates:

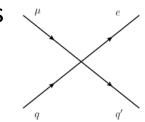
$$pprox$$
 400 to 2 to 1

• If contact terms dominate then μ N \rightarrow eN can have rates 200 times that of μ \rightarrow e γ

 New physics for these channels can come from loop



• For μ N \rightarrow e N and μ \rightarrow eee we also can have contact terms \downarrow^{μ}



Charged Lepton Flavor Violation (CLFV) Processes with μ 's

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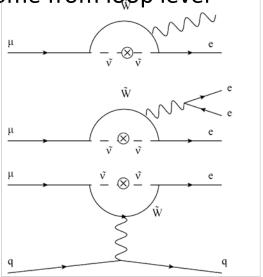
$$\mu^{-} N \rightarrow e^{-} N$$

Note: $\mu \to e\gamma$ and $\mu \to eee$ use a DC beam, and have *experimental* limitations (resolution, overlap, accidentals)

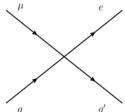
Ultimately Limits the measurement of: ${\rm Br}(\mu \to {\rm e} \ \gamma){\sim}10^{-14}$

No such limits on μ N \rightarrow eN channel

 New physics for these channels can come from loop level

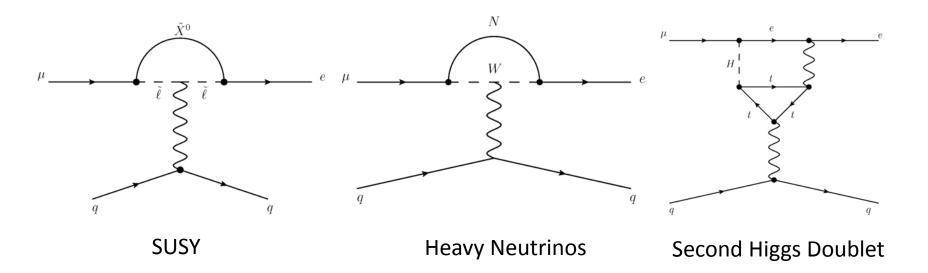


• For μ N \rightarrow e N and μ \rightarrow eee we also can have contact terms



Beyond the Standard Model

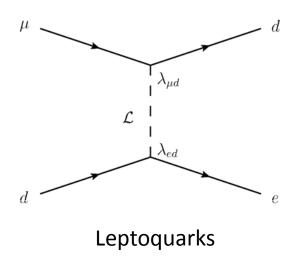
• The CLFV process can manifest in the μ N \rightarrow eN channel in many models with large branching fractions:

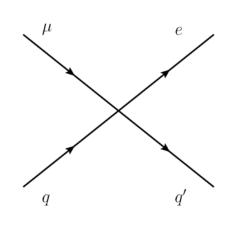


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Contact Terms





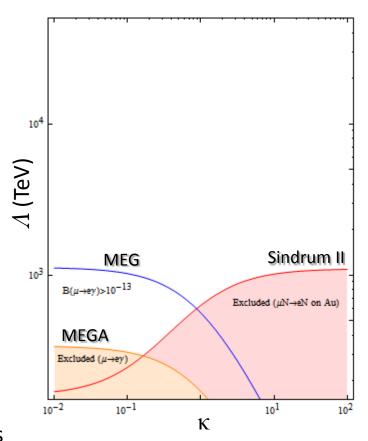
 $\begin{array}{c|c} & & & \\ & & \\ & & & \\ & & & \\ & & & \\ & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ &$

Compositeness

Anomalous Heavy Couplings

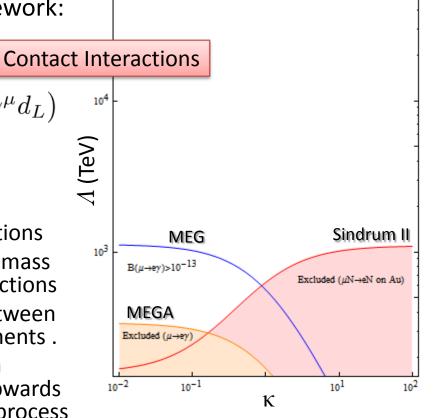
$$\mathcal{L}_{\mathcal{LFV}} = \frac{m_{\mu}}{(\kappa+1)\Lambda^{2}} \bar{\mu}_{R} \sigma_{\mu\nu} e_{L} F^{\mu\nu} + \frac{\kappa}{(1+\kappa)\Lambda^{2}} \bar{\mu}_{L} \gamma_{\mu} e_{L} \left(\bar{u}_{L} \gamma^{\mu} u_{L} + \bar{d}_{L} \gamma^{\mu} d_{L} \right)$$

- Splits CLFV sensitivity into
 - Loop terms
 - Contact terms
- Shows dipole, vector and scalar interactions
- Allows us to parameterize the effective mass scale Λ in terms of the dominant interactions
- The balance in effective reach shifts between favoring μ N \rightarrow eN and μ \rightarrow e γ measurements .
- For contact term dominated interaction (large κ) the sensitivity in Λ , reaches upwards of 10⁴ TeV for the coherent conversion process



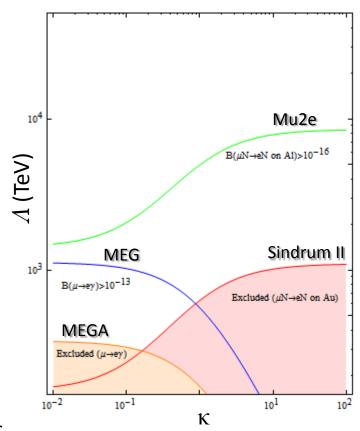
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 Gives the same dipole ensitivity into structure as g-2

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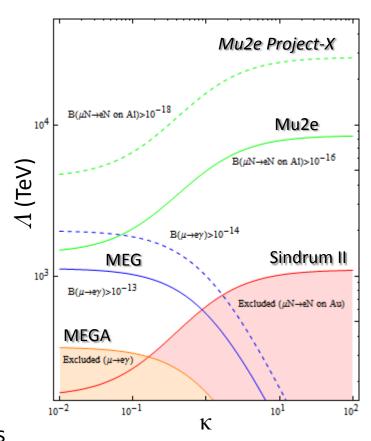
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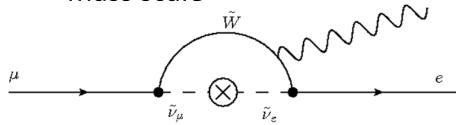
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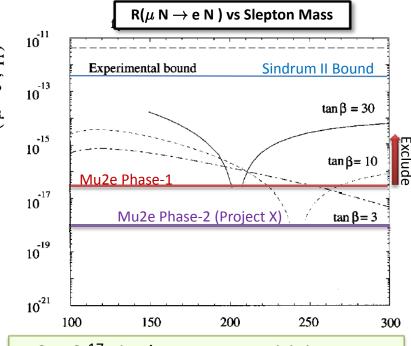
μ NightarroweN Sensitivity to SUSY

 Rates are not small because they are set by the SUSY mass scale



- For low energy SUSY like we would see at the LHC: ${\rm Br}(\mu\ {\rm N}
 ightarrow {\rm e\ N}) \sim 10^{-15}$
- Makes μ N \rightarrow eN compelling, since for Mu2e this would mean observation of

 $\approx \mathcal{O}(40)$ events [0.5 bkg]

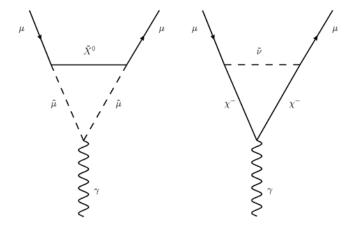


A 2x10⁻¹⁷ single event sensitivity,can exclude large portions of the available SUSY parameter spaces

Hisano et al. 1997

g-2 Sensitivity to SUSY

- SUSY contributes to a_{μ} = (g-2)/2: $a_{\mu}^{SUSY} \approx 130 \times 10^{-11} \left(\frac{100 \text{GeV}}{M_{SUSY}}\right)^2 \tan \beta \ \text{sign}(\mu)$
 - Gives direct access to $tan\beta$ and $sign(\mu)$
 - g-2 result rules out large classes of models
- a_{μ} 's sensitivity to SUSY is through the same loop interactions as CFLV channels



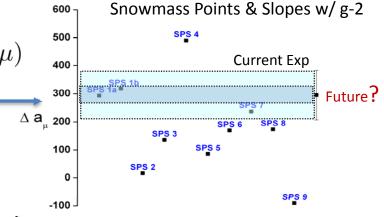
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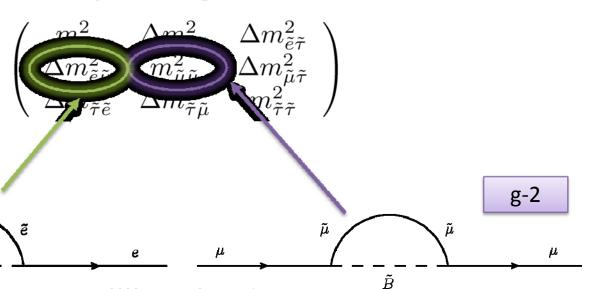


Slepton Mixing Matrix

Gives us access to **Slepton Mixing**

 $\mu \rightarrow e$

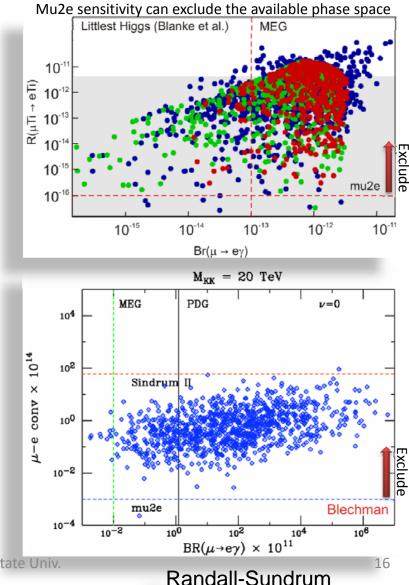
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μ N \rightarrow eN, $\mu\rightarrow$ e γ , g-2 Work Together

- Knowing μ N \rightarrow eN, μ \rightarrow e γ allow us to exclude SUSY phase space
- Also knowing the g-2 results allows us to then over constrain SUSY models
- In some cases this permits us to make strong, testable predictions for our models in terms of $Br(\mu \rightarrow e\gamma) \& R(\mu N \rightarrow eN)$



μ N \rightarrow eN, $\mu\rightarrow$ e γ , g-2 Work Together

10-11

10⁻¹³

10.14

10⁻¹

10-16

Mu2e sensitivity can exclude the available phase space

mu2e

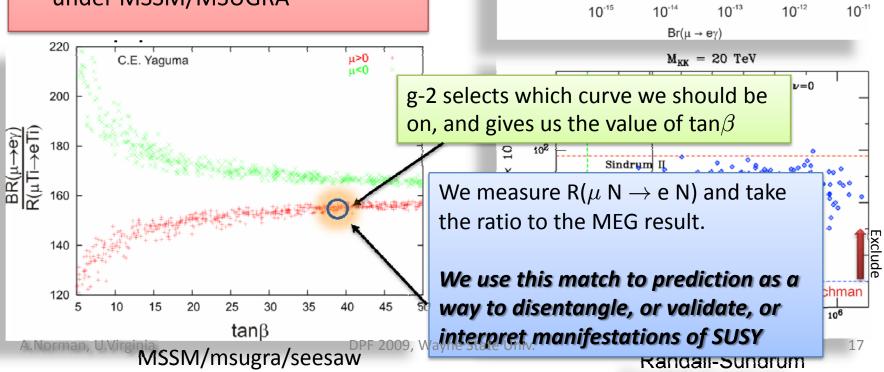
Littlest Higgs (Blanke et al.)

Example:

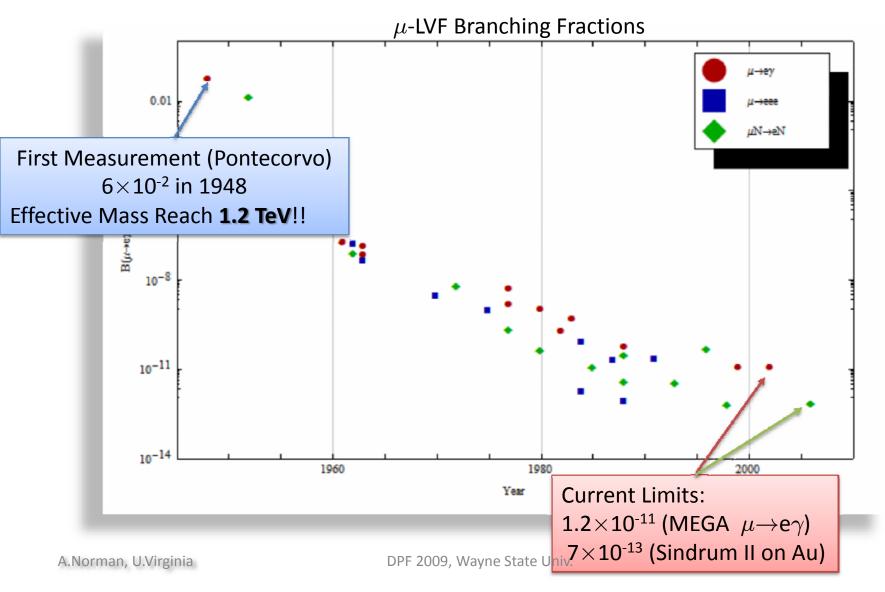
- From LHC we have the SUSY masses
- From g-2 we know $tan\beta$
- From g-2 we know also know μ >0
- Combining these we get an a priori PREDICTION for:

$$\frac{Br(\mu \to e\gamma)}{R(\mu N \to eN)}$$

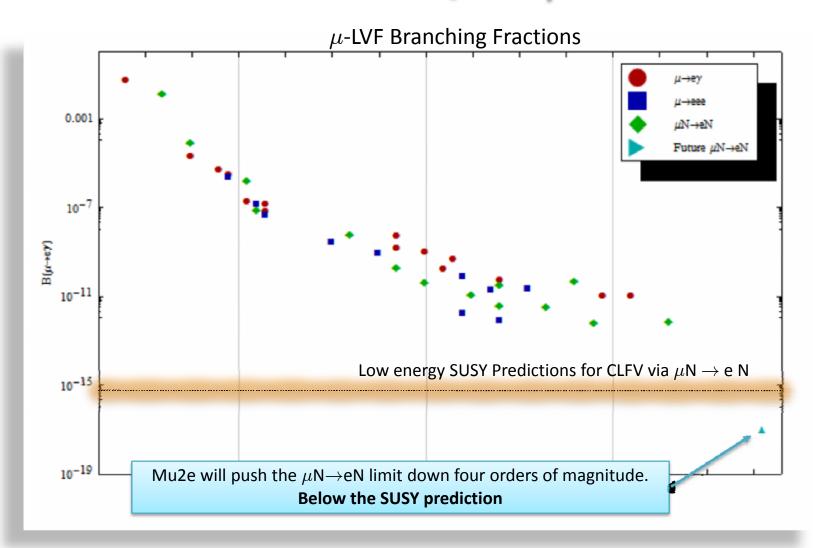
under MSSM/MSUGRA



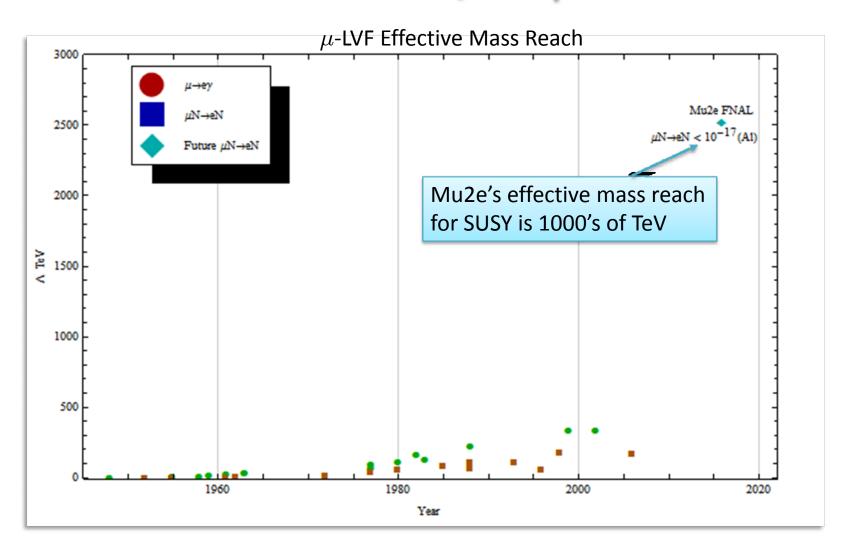
A Brief History of μ -cLFV

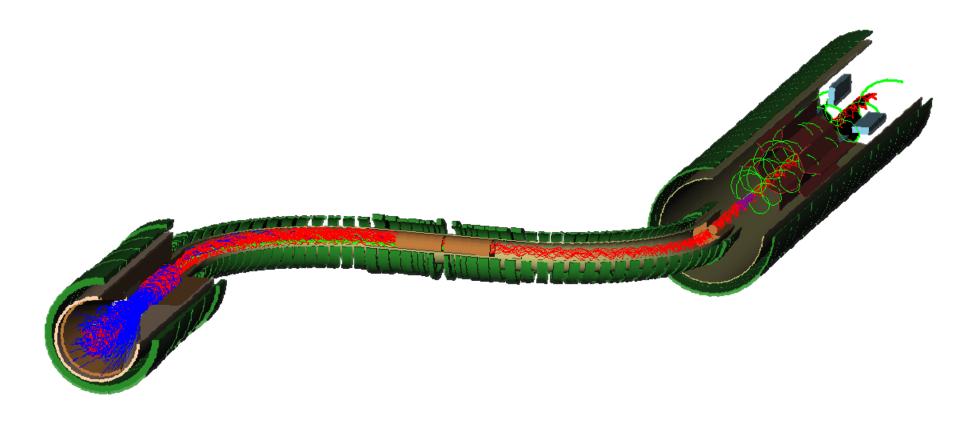


A Brief History of μ -cLFV



A Brief History of μ -cLFV





MU2E AT FNAL

The μ N \rightarrow eN measurement at Br(10⁻¹⁷) (in a nutshell)

- Stop $\sim \mathcal{O}(5 \times 10^{10})~\mu^{-}$ per pulse on a target (Al, Ti, Au)
- Wait 700ns (to let prompt backgrounds clear)
- Look for the coherent conversion of a muon to a monoenergetic electron:

$$E_e = M_{\mu} - N_{recoil} - (B.E.)_{\mu}^{1S}$$

= 104.96 MeV (on ²⁷Al)

Report the rate relative to nuclear capture

$$\mathcal{R} = \frac{\Gamma(\mu^- N \to e^- N)}{\Gamma(\mu^- N(Z) \to \nu_\mu N(Z-1))}$$

 If we see a signal, it's compelling evidence for physics beyond the standard model!

μ NightarroweN in Detail

Muonic Atom

- Start with a series of target foils
 - For Mu2F these are Al or Ti
- Bring in the low energy muon beam
 - \bullet We stop pprox 50% of μ 's
 - Stopped muons fall into the atomic potential
 - As they do they emit x-rays
- Muons fall down to the 1S state and a captured in the orbit
 - Muonic Bohr Radius

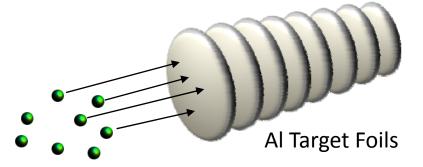
$$\langle r_{\mu} \rangle = \frac{n^2 \hbar}{m_{\mu} z e^2} \approx 19.6 fm \ (for \ Al)$$

• Nuclear Size

$$R \approx 1.2A^{1/3}fm = 3.6fm (forAl)$$

- Provides large overlap in the muon's wavefunction with the nucleous's
- For Z > 25 the muon is "inside" the nucleous
- Once captured 3 things can happen
 - Decay in Orbit: $\mu^-
 ightarrow e^-
 u ar{
 u}$

 \approx 50% stop in target



Target 200 µm, circular foils (²⁷Al) Radius tapers from 10 cm to 6.5 cm 5cm spacing between foils

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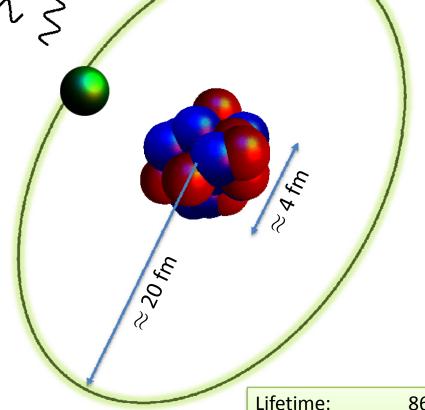
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We use the cascade of muonic x-rays and the well known spectrum to normalize the experiment.

(i.e. We measure N_{stop} in real time)



Lifetime: 864ns
DIO Fraction: 39.3%
Capture Fraction: 60.7%

μ NightarroweN in Detail

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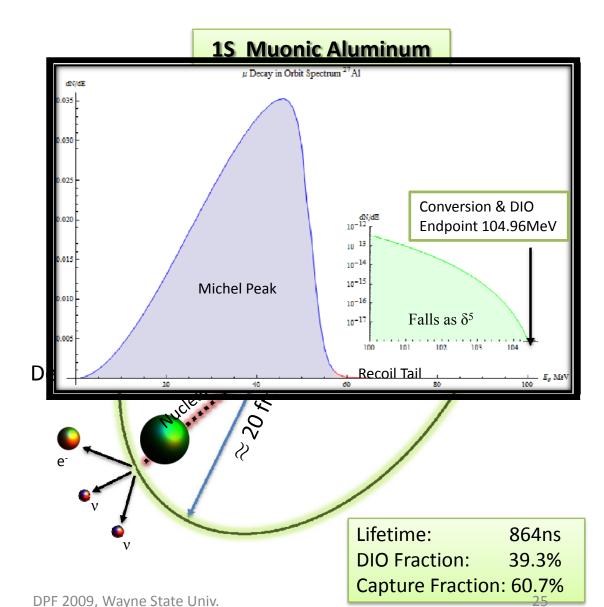
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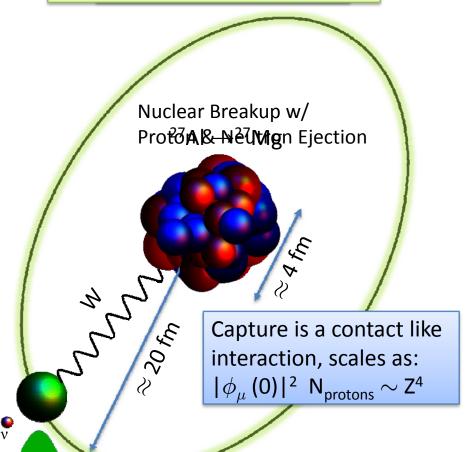
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Ordinary Muon Capture (OMC)



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DPF 2009, Wayne State Univ.

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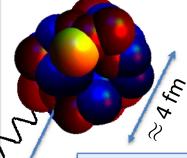
Problem

• As they do they These protons and neutrons constitute the largest source of rate in the detector (\sim 1.2 per μ)

The energy spectra for these $R \approx 1.2A^{1/3}f^{\eta}$ ejected particles is not well known.

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Nuclear Breakup w/ Proton Ejection



Capture is a contact like interaction, scales as:

 $|\phi_{\mu}\left(0
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Ordinary Muon Capture (OMC)

Nuclear Breakup w/

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2009 Measurements

Mu2E Collaboration will measure the spectra for muon capture induced nucleon emission this summer at the PSI test beam.

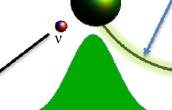
Contact: Peter Kammel (U. of Illinois) pkammel@illinois.edu for details

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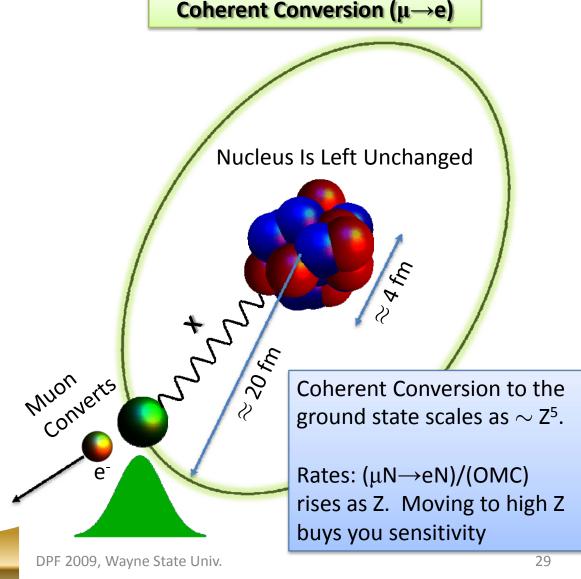
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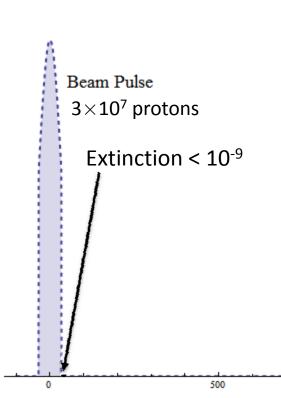
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 - New Physics! i.e. μ N \rightarrow e N



- μ 's are accompanied by "prompt" e, π ,
- These cause real background
- Must limit our beam extinction, and detector live window



Prompt Backgrounds

Radiative Pion Capture (RPC) presents the single most dangerous potential background to the Mu2e experiment.

$$\pi^- N \to \gamma N^* (Z-1)$$

The gamma can convert asymmetrically giving e^{-} with energy up to m_{π}

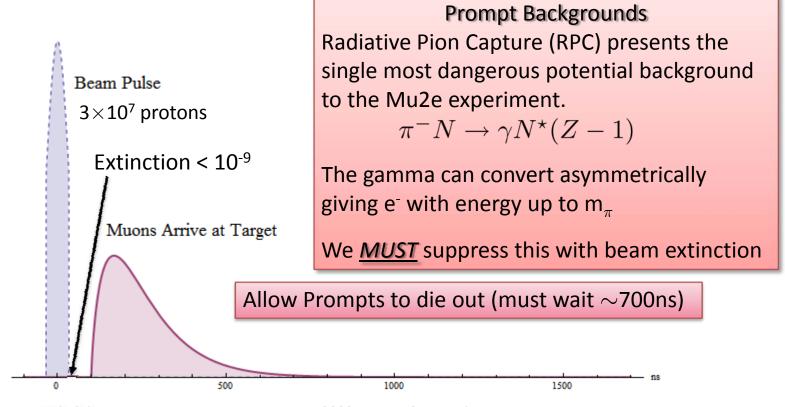
We **MUST** suppress this with beam extinction

1500

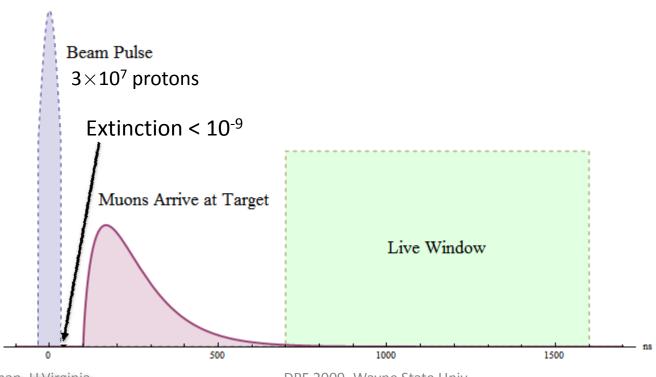
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1000

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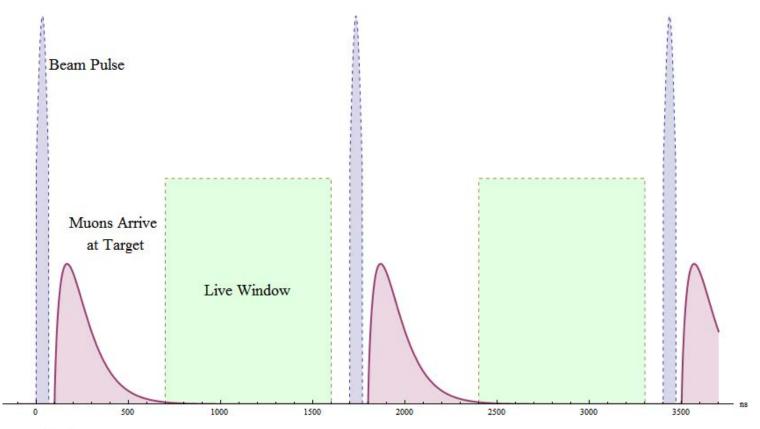
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- Must limit our beam extinction, and detector live window



A.Norman, U.Virginia

DPF 2009, Wayne State Univ.

Total Backgrounds

- Largest Background
 - Decay in Orbit (DIO)
 - Rad π Capture (RPC)
- Limiting Backgrounds
 - Can limit prompt backgrounds w/ extinction
 - In particular, Rad π Cap. drives the extinction requirement
 - Current Background Estimates require 10⁻⁹ extinction
 - BNL AGS already has demonstrated extinction of 10⁻⁷ with out using all the available tools

Background	Evts (2×10 ⁻¹⁷)
μ Decay in Orbit (DIO) Tail	0.225
Radiative pion capture	0.072
Beam Electrons	0.036

Total Backgrounds

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μ Decay in Orbit (DIO) Tail	0.225
μ Decay in flight w/ scatter	0.036
Beam Electrons	0.036
Cosmic Ray	0.016
μ Decay in flight (no scatter)	< 0.027
Anti-proton	0.006
Radiative μ capture	< 0.002
Radiative π capture	0.072
π Decay in flight	<0.001
Pat. Recognition Errors	<0.002
Total	0.415

Total Backgrounds

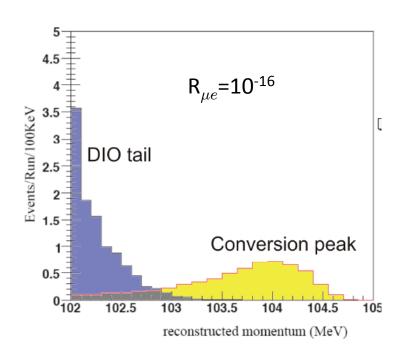
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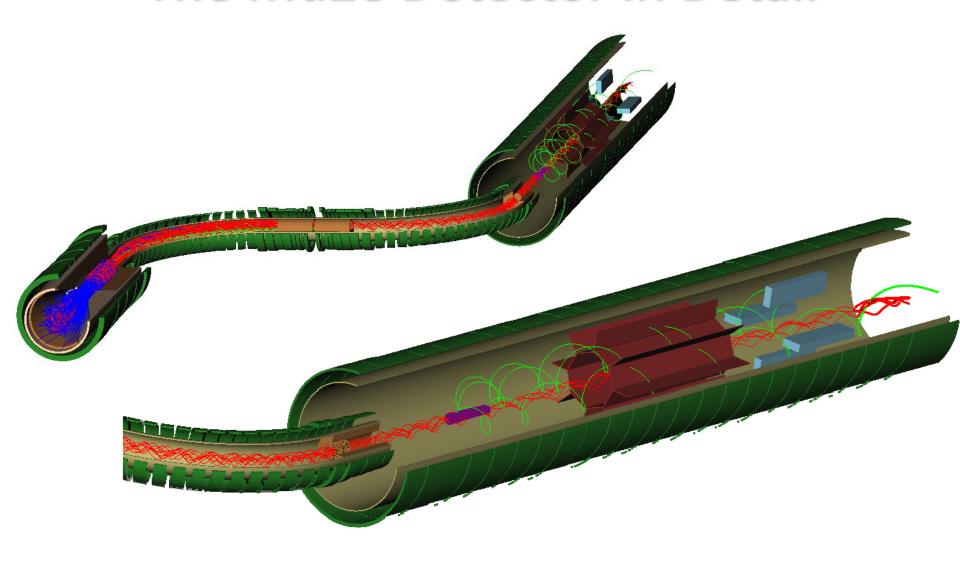
Signal to All Backgrounds

- Signal significance
 - If we assume SUSY accessible at the LHC:
 - Mu2e may see $\sim \mathcal{O}(40)$ events
 - On 0.5 event background
 - At $R_{\mu e}$ =10⁻¹⁶ (limit of sensitivity)
 - Mu2e sees \sim 4 events
 - on 0.5 event background
 - This is a Strong Signature

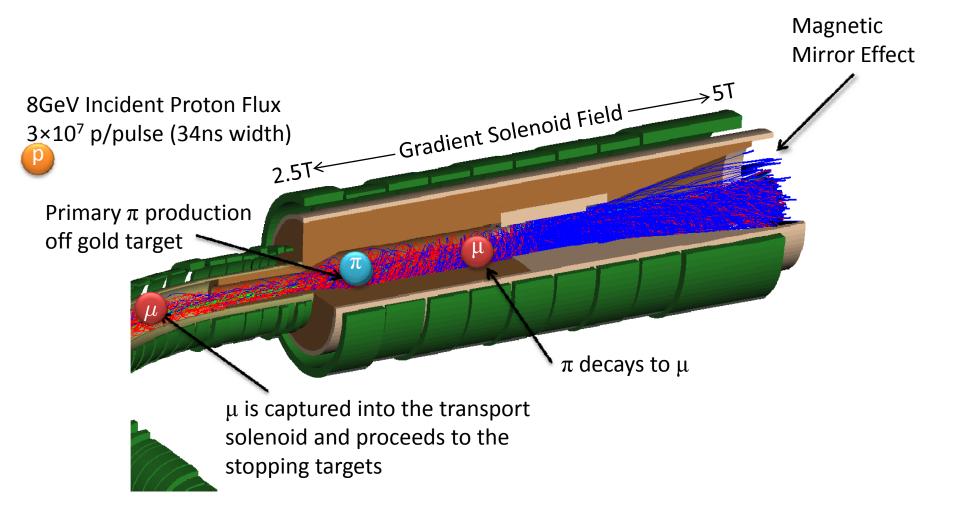
$$\frac{S}{\sqrt{B}} \sim 5.5$$



The Mu2e Detector in Detail

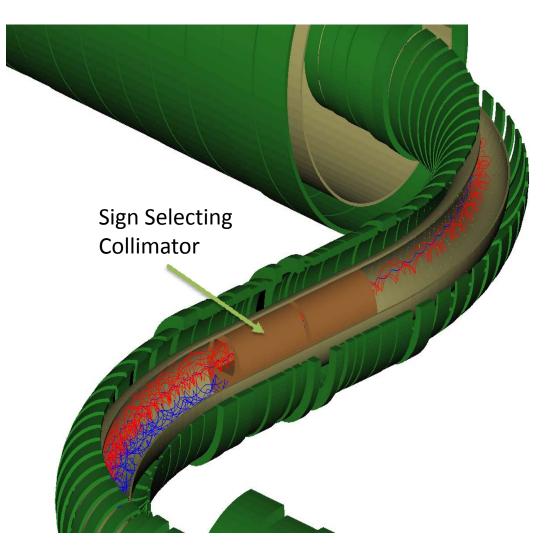


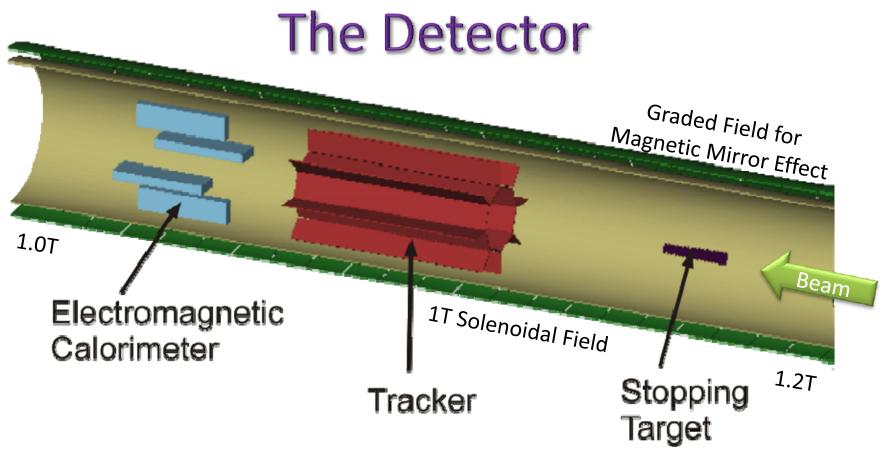
Production Solenoid



Transport Solenoid

- Designed to sign select the muon beam
 - Collimator blocks the positives after the first bend
 - Negatives are brought back on axis by the second bend

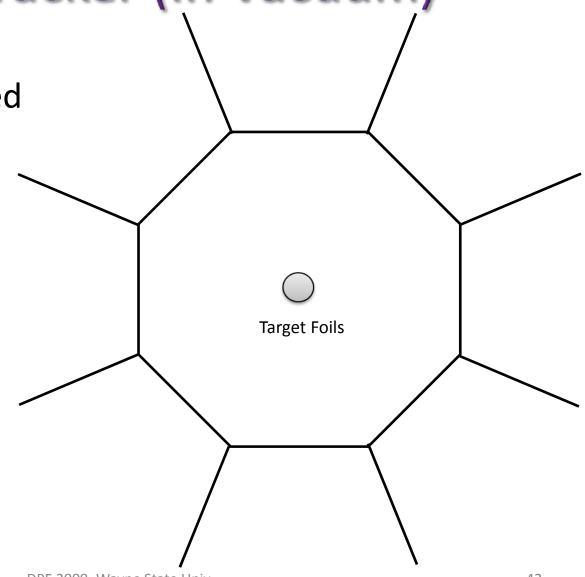




- The detector is specifically design to look for the helical trajectories of 105 MeV electrons
- Each component is optimized to resolve signal from the *Decay in Orbit* Backgrounds

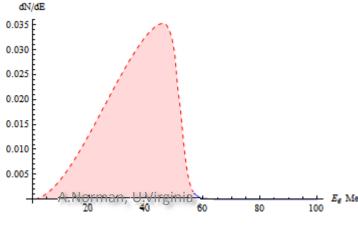
 Octagonal+Vanes geometry is optimized for reconstruction of 105MeV helical trajectories

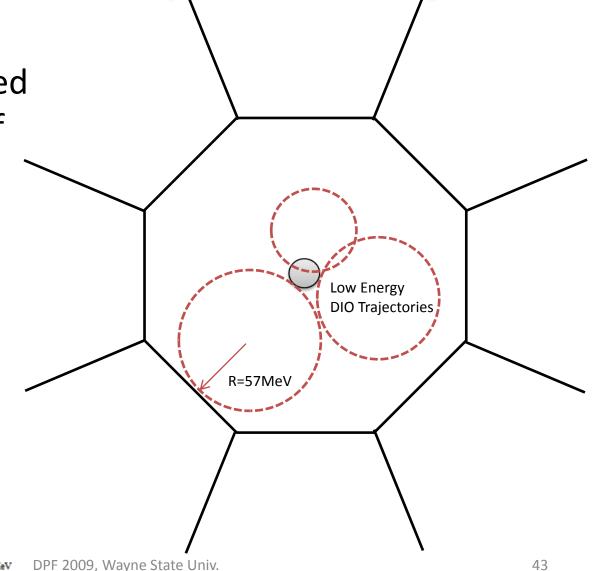
- Extremely low mass
- Acceptance for DIO tracks < 10⁻¹³



 Octagonal+Vanes geometry is optimized for reconstruction of 105MeV helical trajectories

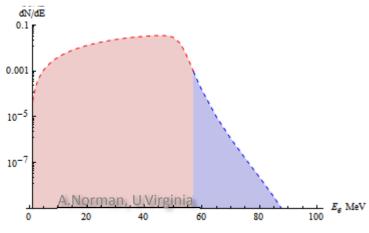
Extremely low mass

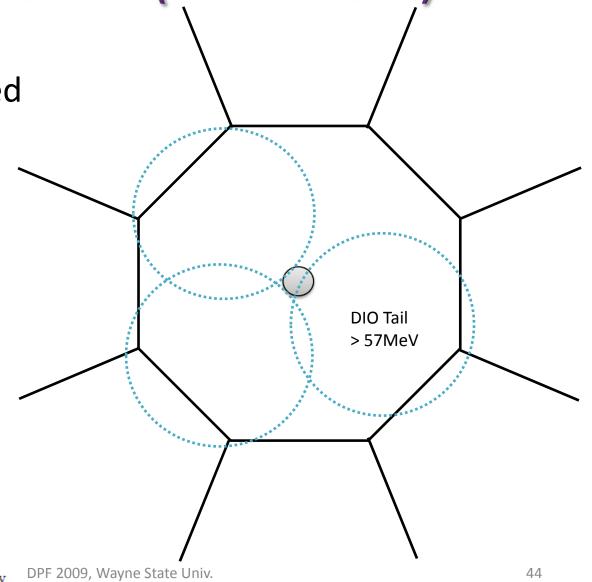




 Octagonal+Vanes geometry is optimized for reconstruction of 105MeV helical trajectories

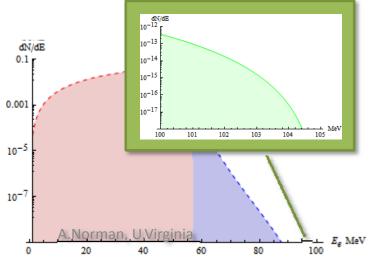
Extremely low mass

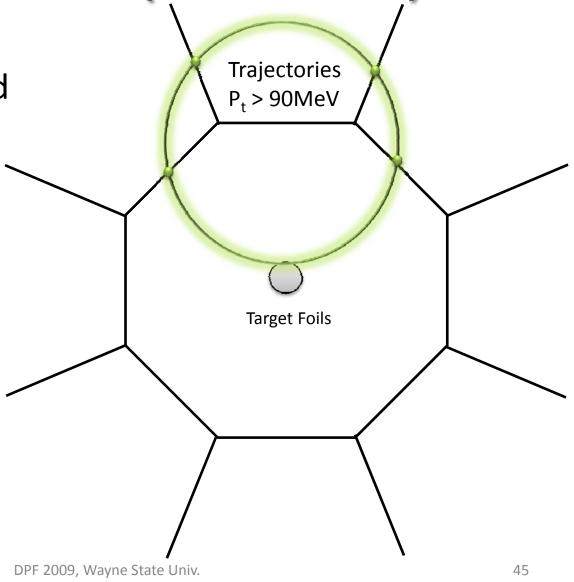


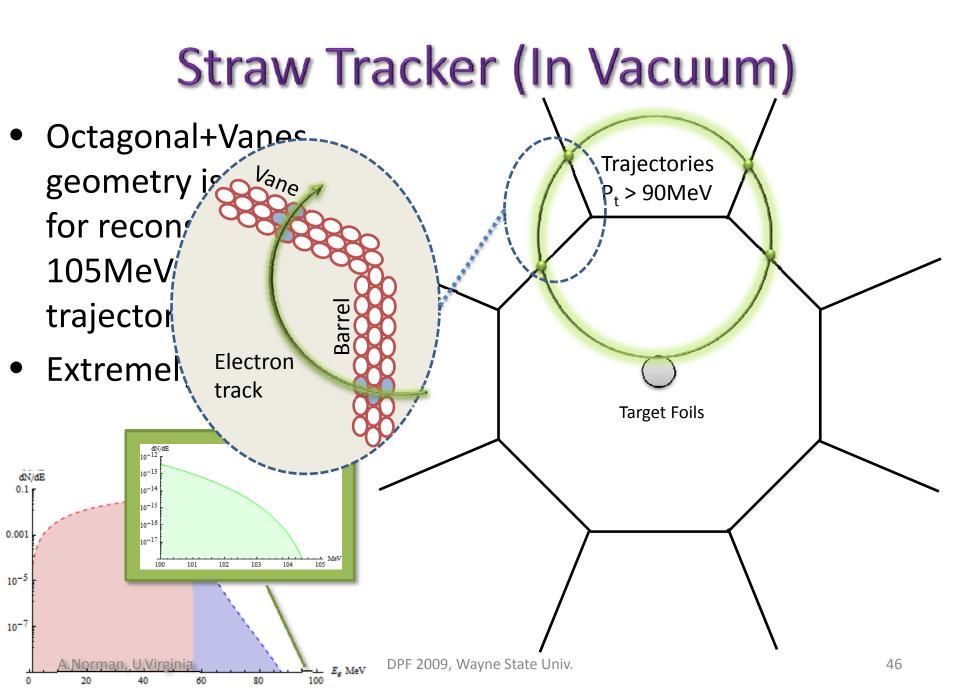


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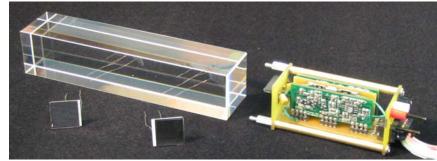


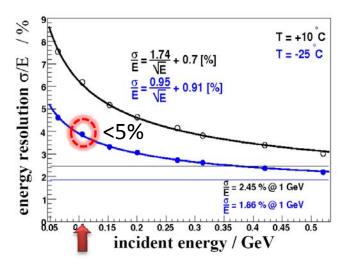


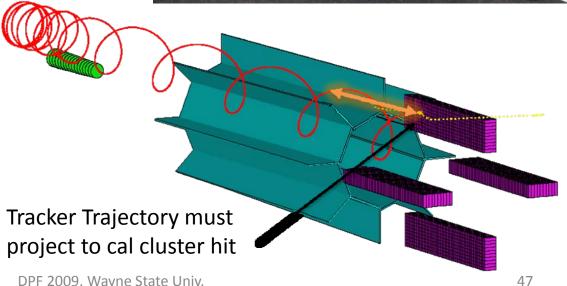
Crystal Calorimeter

Original Design:

- 5% energy measure for trigger decision (1kHz rate)
- Timing edge for event reconstruction
- Spatial match to tracker trajectory
- Low acceptance to Michel Peak





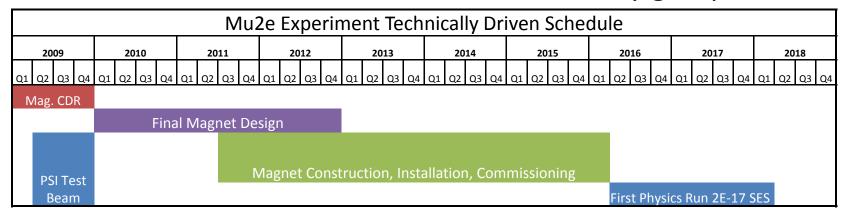


A.Norman, U.Virginia

DPF 2009, Wayne State Univ.

Cost and Schedule

- Total Project Cost Est. \$200M (fully loaded, escalated, appropriate contingencies)
- Received Stage-1 Approval and DOE's CD-0 anticipated shortly
- Technically Driven Schedule (wholly magnet driven) results in 2016 start of data taking
- Opportunities for Significant R&D, Test Beam, and Auxiliary
 Measurement work for students and university groups



Conclusions

- In an era where, we are poised to see our first direct evidence of physics beyond the standard model:
- We must pay special attention to precision measurements
- Charge Lepton Flavor Violation experiments have the ability, not only guide us as we begin to interpret and understand signs of new physics, But they naturally combine to:
 - Make elegant predictions
 - Probe large parameter spaces
 - and access physics beyond the Terascale
- Consider the possibilities and join us!
 - Mu2e: http://www-mu2e.fnal.gov
 - Spokesperson Contacts:
 - Robert Bernstein (rhbob@FNAL.gov)
 - James Miller (miller@buphy.bu.edu)

BACKUP SLIDES

Mu2e Collaboration

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Rice University
M.Corcoran

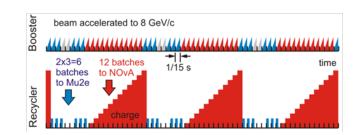
Syracuse University P.Souder, R.Holmes

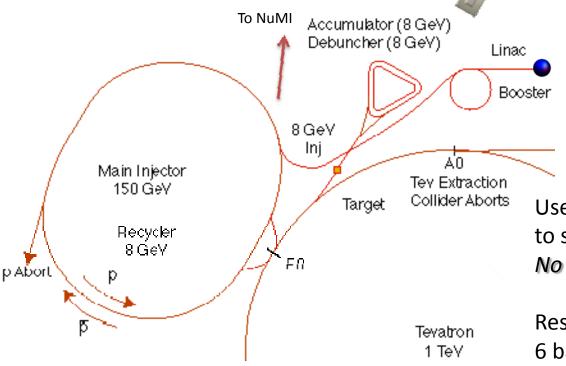
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R.Hirosky, A.Norman, K.Paschke,
D.Pocanic

College of William and Mary
J.Kane

Mu2E & NOvA/NuMI

• How do we deliver $\mathcal{O}(10^{18})$ bunched μ 's?





Use NuMI cycles in the Main injector to slow spill to Mu2e.

No Impact on NOvA

Results in:

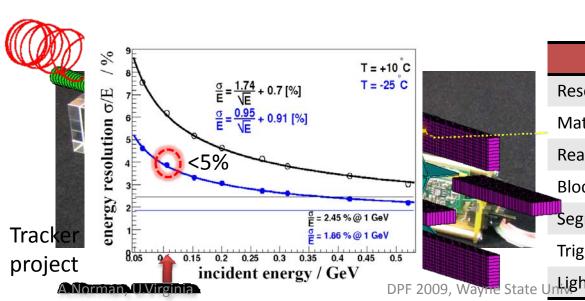
6 batches x $4x10^{12}/1.33$ s x $2x10^7$ s/yr

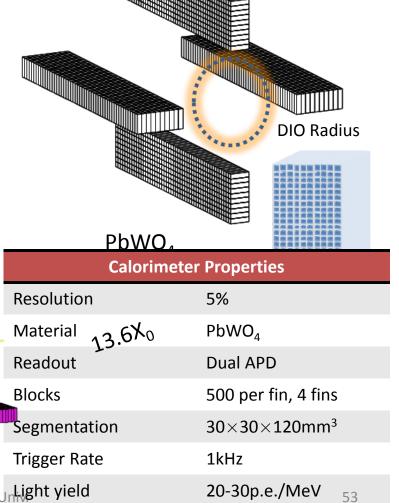
DPF 2009, Wayne State Univ. = 3.6 x10²⁰protons/yr

Crystal Calorimeter

Original Design:

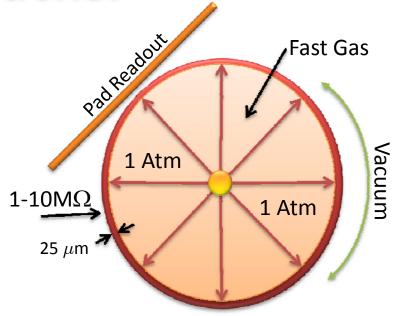
- 5% energy measure for trigger decision (1Hz rate)
- Timing edge for event reconstruction
- Spatial match to tracker trajectory
- Immune to DIO rates



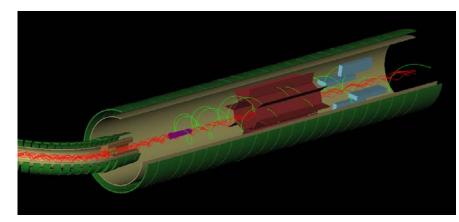


Straw Tracker

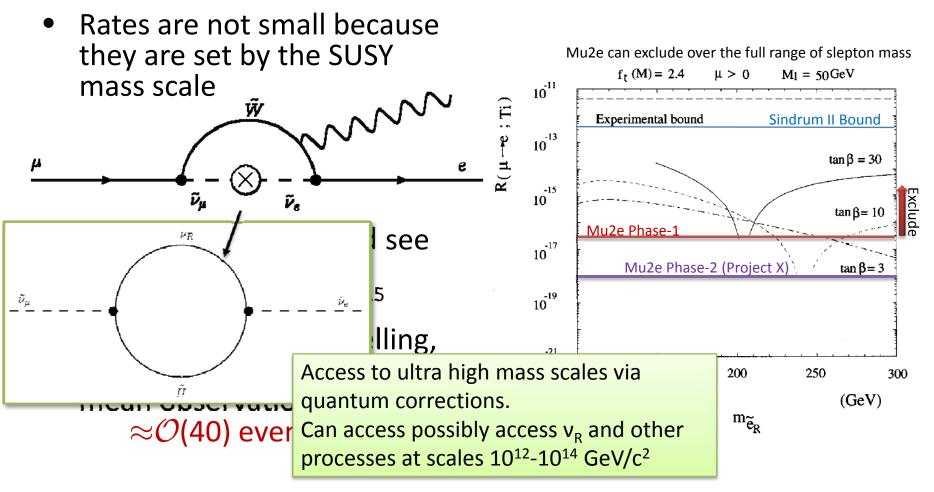
- Longitudinal Tracker
 Features:
 - 2800 straw tubes in vacuum
 - Utilize 17,000 pad readouts
 - 50% Geometric acceptance to signal (90 $^{\circ}$ \pm 30 $^{\circ}$)
 - Intrinsic resolution 200keV
 - Virtually Immune to DIO



Diameter Swell > 1%



Sensitivity to SUSY



Hisano et al. 1997

μ N ightarrow eN & SUSY Models

- Assuming we see a signal:
 - By changing target, we gain sensitivity to the scalar, vector or dipole nature of the interaction
 - Need to go to high Z
 - Hard because τ small for large Z (τ_{Au} =72ns)
 - But DIO backgrounds are suppressed and Conversion/OMC ratio scales as Z
- This is a unique feature of the μ N \rightarrow eN measurements

